Analysis of the Spectrum of Neutral Atomic Bromine (Br)

Jack L. Tech

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The spectrum of the neutral bromine atom, Bri, has been newly investigated by using electrodeless discharge tubes as light sources. The observations have led to a list of wavelengths and estimated intensities for 1253 spectral lines in the range 1067 to 24100 Å. number of known energy levels has been increased to 123 even and 128 odd levels, as compared with the 27 even and 33 odd levels previously known. All predicted energy levels of the $4s^2 4p^4 ns, np, nd, nf$ electron configurations from 0 to ~ 93250 K have been discovered. The observations in the vacuum ultraviolet establish that the positions of all the levels lying above those of the ground configuration as given in the compilation Atomic Energy Levels, Vol. II (1952) should be increased by 6.7 K. All but 26 faint lines of Br I have been classified. A total of 67 levels has been ascribed to the $4s^2$ $4p^4$ nf configurations. It is demonstrated that the nf configurations exhibit almost pure pair coupling. The very regular (3P2)nf[5]°11/2 series yields for the principal ionization energy of Brithe value 95284.8 K.

1. Introduction

It has been part of the recent program of this laboratory to obtain improved descriptions and analyses of the first and second spectra of the heavier halogens. Work on the BrII and ClI spectra is currently in progress, and the results for II and III have already been published [1, 2]. It is the purpose of the present paper to report the results for the first spectrum of bromine, Br1.

Iodine and bromine are very frequently used in electrodeless metal-halide lamps [3] serving as light sources for the study of rare-earth and other metallic spectra. Since the spectrum of both the metal and the halogen will appear when such a lamp is excited, it is essential that the user have available a complete and accurate description of the spectrum of the halogen in order to separate the halogen lines from those

of the metal under investigation.

The most thorough study of Br I hitherto available is that contained in the excellent paper on the structure of the arc spectrum of bromine published 30 years ago by C. C. Kiess and T. L. de Bruin [4]. That paper gives, also, a summary of the investigations of Bri carried out prior to 1930. The Bri analysis by Kiess and de Bruin appears with minor revisions in the compilation "Atomic Energy Levels" [5], and when we refer to the analysis given by those authors we shall mean the form as presented in AEL.

The most significant contributions to the study of Br I to appear since 1930 are those by P. Lacroute [6], who studied the Zeeman effect in this spectrum and made some wavelength measurements in the Schumann region; and by Tolansky and Trivedi [7],

who made a rather extensive study of hyperfine structure (hfs) in Br1. In the course of the latters' analysis of the structure of 64 lines in the range 4390 to 8700 Å it was found that a number of lines exhibited hfs patterns that could not be satisfactorily explained on the basis of the interpretation of these lines as offered by Kiess and de Bruin. Without going into details, we merely state that all the anomalies encountered by Tolansky and Trivedi have been eliminated by the extended analysis of this spectrum given in the present work.

This considerably improved analysis results largely from the use of more refined apparatus, photographic emulsions, and light sources than were formerly available. The analysis rests chiefly on the 1035 spectral lines that we have recorded for Bri by using electrodeless discharge sources in the photographic air region. In comparison, Kiess and de Bruin's line list contained only 330 lines, which is reduced to 274 when the 56 lines listed by them but not con-

firmed in our work are rejected.

The analysis of Bri as given in AEL has undergone a number of changes as a result of clues furnished by the greatly augmented line list. These changes can be summarized briefly as follows:

	Number		
	Even	Odd	
Original levelsRejected	39 12	39	
J or designation changed Total real	$\begin{array}{c} 7 \\ 27 \end{array}$	17 33	

¹ Figures in brackets indicate the literature references at the end of this paper.

The new wavelengths have permitted us to find 123 even and 128 odd levels for Br I compared with the 27 even and 33 odd levels listed in AEL.

2. Wavelength Material

(a) Vacuum region: We have observed the bromine spectrum below 2000 Å by use of a vacuum spectrograph having as dispersing element a concave grating of two-meter radius and ruled with 30,000 lines per inch directly on Pyrex. The plate factor at the normal (1920 Å) is 4.26 Å/mm, and the plateholder covers the range 0 to 2570 Å in the first order.

The light source was an end-on glass discharge tube having two side-arms set off by stopcocks and containing reservoirs of bromine and iodine. discharge tube was affixed to the slit housing of the spectrograph by means of an O-ring seal, so that the slit opened directly to the discharge. The bromine vapor was retarded from entering the main spectrograph chamber by continuous pumping through a trap that entered the discharge tube just before the slit. The bromine vapor pressure in the discharge, excited in the field of 2450 Mc/s radiation from a microwave oscillator, was partly controlled by regulating the temperature at the bromine reservoir. Helium was also admitted into the discharge. auxiliary iodine reservoir was used to provide standards [1] to supplement the impurity standards of hydrogen, helium, carbon, nitrogen, oxygen, and chlorine [8, 9]. We used EK 103a-O UV and SWR plates and measured the 1000 to 1280 Å region in the second order and the 1000 to 1650 Å region in the first order.

Table A1 in the appendix lists the 124 spectral lines in the vacuum region that we have attributed to Br I. The observed wavelengths in the first column of this table represent the weighted means of only two measurements in the second order for the range 1067 to 1120 Å and of three or more measurements in the first and second orders for the range 1120 to 1633 Å. We have also made use of measurements by W. C. Martin and C. H. Corliss of certain bromine lines that occurred on their iodine

The new wavelengths have made it possible to establish the ground Br I energy separation $4p^5 \, ^2P_{3/2}^{\circ} - ^2P_{1/2}^{\circ}$ as 3685.2 K with an estimated error of ± 0.3 K.² Since none of the higher Br I energy levels make transitions to the ground term that yield lines in the air region, it is necessary to fix the energies of these higher levels with respect to the ground $^2P_{3/2}^{\circ}$ level by use of the vacuum observations only. Our new vacuum wavelengths establish the absolute positions of all levels above the ground level to an estimated accuracy of ± 0.3 K and require that the corresponding values for Br I levels in

spectrograms.

² Throughout this paper we use the name kayser, with symbol K, for the unit

AEL be increased by 6.7 K.

Since the relative positions of the higher Br I levels have been accurately fixed by observations in the air region, we can calculate the Br I lines in the vacuum region on an accurate and uniform scale once we have made the energy connection between the ground and excited levels through our measured vacuum wavelengths. These calculated wavelengths are given in column 2 of table A1; they agree to a better extent with the measured wavelengths over the entire range than to the ± 0.01 Å we would have ventured to estimate solely on the basis of examination of the agreement of wavelengths derived from the different exposures, and especially in recognition of the inadequate standards in certain parts of the region.

(b) Air region: A preliminary list of Brī wavelengths in the photographic air region was given earlier by Tech and Corliss [10]. One will find in that paper a thorough description of the experimental procedure and the preparation of the light sources,

which we summarize briefly. Workers have encountered in the past two major difficulties in obtaining a light source suitable for the investigation of the bromine spectrum: (i) bromine is so chemically active a material that it attacks hot internal electrodes; and (ii) water vapor in the bromine causes the appearance of troublesome extraneous bands. Both these difficulties were eliminated in the present work by using electrodeless discharge lamps, excited by a Raytheon Microtherm microwave generator at 2450 Mc/s. For our first series of observations, we used lamps containing BeBr₂, but for the more recent series we used lamps containing pure bromine vapor at a pressure of the order of several millimeters of mercury and dried by means of several passes through P_2O_5 as described in [10]. The absence of water vapor in our lamps is attested to by figure (1), displaying the bromine spectrum in the vicinity of $H\alpha$ (6563 Å). It will be noticed that $H\alpha$ does not appear at all in the bromine exposure. On the other hand, this figure shows clearly the highly objectionable Br₂ background that occurs on spectrograms of long exposure and that reduces the effective resolving power. Furthermore, the presence of such bands often makes it difficult to decide whether certain faint lines are of atomic or molecular origin.

The spectrograms were made in the first and, wherever possible, second orders with concave gratings having 7,500, 15,000, and 30,000, lines per inch, and mounted in parallel light (Wadsworth mounting), giving first order plate factors of 10, 5, and 2.4 Å/mm respectively. All the spectrograms bore exposures to an electrodeless lamp containing thorium or iron, or to the iron arc, to serve as secondary standards in the determination of the bromine wavelengths. The plates were reduced either by linear interpolation and a correction curve, or by putting the micrometer readings of the lines on punched cards and carrying out the reduction by use of an IBM 704 computer, which adjusted all wavelengths to fit a fourth degree equation determined by least squares.

Table A2 contains the wavelengths, intensities, wave numbers, and classifications of 1035 Br I lines

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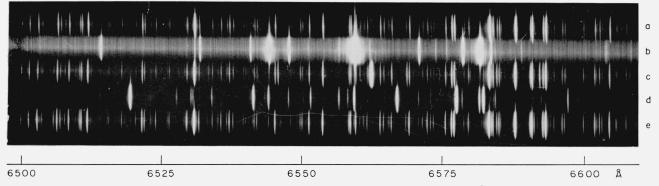


Figure 1. The bromine spectrum in the vicinity of $H\alpha$ (6563 Å).

Exposure (b) was made with an electrodeless discharge tube containing bromine vapor excited by a microwave diathermy unit. Exposure (d) is of bromine excited by a high voltage ring discharge apparatus that enhanced Br II. The other exposures are of thorium, which served as secondary standards. H α occurs only in the middle thorium exposure (e).

observed in the range 3325 to 12810 Å. Except for the weaker lines in the region 4264 to 5721 Å, and except for very faint lines over the whole range, the wavelengths in the table represent the weighted means of from two to eight measurements. Because of the inherent broadness of many of the lines, partly as a result of unresolved hyperfine structure, and because the weaker lines were only measured once or twice, the accuracy of the wavelengths is not as great as one would wish. The estimated error of the wavelengths according to spectral region is:

λ	Error
3325–4300 Å	\pm 0. 005 Å
4300-7400 Å	$\pm~0.~02~{ m \AA}$
7400–11000 Å	$\pm 0.03 \ \mathrm{\AA}$
11000–12900 Å	$\pm0.06\mathrm{\AA}$

Lines in table A2 whose wavelengths are connected by braces are individually measured components of resolved or partially resolved hyperfine structure. Most of the Br I lines exhibiting such wide, resolved hfs patterns result from transitions involving the $4p^4(^1\mathrm{D}_2)5s~^2\mathrm{D}_{5/2},~^2\mathrm{D}_{3/2}$ levels or the $4p^4(^3\mathrm{P}_2)5s~^4\mathrm{P}_{5/2}$ level, a clue that was very useful in the analysis.

The character of certain lines is indicated by inserting the following symbols after the intensity

of the line:

c almost resolved hyperfine structure

d unresolved double line

h hazy, diffuse

H very hazy, very diffuse

s shaded to shorter wavelengths

w wide

W very wide.

The numbers in column 2 of the table are visual estimates of the intensities on an approximately linear scale where the faintest line is assigned an intensity 1. It must be emphasized that these relative intensities are valid only over small wavelength ranges, and no attempt has been made to adjust the values to take into account the varying sensitivity of the photographic emulsions.

Lines of wavelength greater than 11316 Å were measured also on the infrared recordings described in the next section. For comparison, the intensities derived from these recordings are listed in parentheses following the photographic intensities for these lines in table A2. The symbol (m) indicates that the line was masked on the recorder charts either by a standard line or by a bromine line of another order.

Table A2 incorporates several improvements over our earlier list of wavelengths. Additional chlorine lines were deleted by consulting the unpublished chlorine line list of Corliss. Chlorine accounted for nearly all the impurity lines found on our spectrograms. Furthermore, several faint lines, measured only once, especially in the region 4725 to 4825 Å have been ascribed to Br₂ and removed from the list. The line at 4358.33 Å has been attributed to Hg and probably originated in the scattered light entering the slit from fluorescent illumination in our source room.

More importantly, additional plates have been measured in parts of the spectrum above 7400 Å. This has not only yielded a number of new lines in this region, but also has allowed us to correct a small systematic error of uncertain origin that was present in some of the earlier measurements in this region. The new wavelengths from 7400 to 9700 Å differ, if at all, by only about +0.01 Å from the earlier values. Above 9700 Å the correction averages somewhat less than +0.04 Å.

Finally, W. F. Meggers and R. Zalubas have generously provided us with some Br I wavelengths that they had measured on their Yb and Th spectrograms, respectively. These measurements have also been averaged into our new mean wavelengths.

(c) Lead sulfide region: The infrared description of Br I was extended to 24100 Å by the use of recording infrared spectrometers. The wavelengths of 94 new lines were measured by the author from radiometric records in the form of recorder charts generously supplied expressly for this investigation by C. J. Humphreys at the Naval Ordnance Laboratory,

Table 1. Wavenumbers and intensities of selected multiplets observed in the radiometric region

10										
	⁴ F _{9/2} 80587. 61	$^{4}_{81081.10}$	$^{4}\mathrm{F}_{5/2}$ 81672. 28	⁴ F _{5/2} 81842. 02	⁴ D _{7/2} 79044. 50	$^{4}_{79260.91}$	$d = {}^{4}D_{\epsilon/2} = 79630.67$	⁴ D _{1/2} 80026. 40	$^{4d}_{^{4}\mathrm{P}_{1/2}}$ 81429. 93	7
$\begin{array}{c} 5p \\ ^{4}\mathrm{P}_{5/2}^{s} \\ 74672.32 \\ ^{4}\mathrm{P}_{3/2}^{s} \\ 75009.13 \\ ^{4}\mathrm{P}_{1/2}^{s} \\ 75814.00 \end{array}$		(18) 6408. 78	(35) 6663. 15	(250) 6028. 02	(950) 4372. 18	(469) 4588, 59 (206) 4251, 78	(138) 4958, 35 (359) 4621, 54	(138) 5017. 27 (40) 4212. 40	(219) 6420. 84 (330) 5615. 97	4
$\begin{array}{c} 5p\\ ^{4}\mathrm{D}_{7/2}^{\circ}\\ 75521.\ 50\\ ^{4}\mathrm{D}_{5/2}^{\circ}\\ 75697.\ 05 \end{array}$	(3450) 5066. 11	(40) 5559. 60 (500) 5384. 05	(90) 6150. 78 (1800) 5975. 23	(200)			4 <i>d′</i> ² 8335.		d' ² F _{5/2} 723. 08	
⁴ D _{3/2} 76743. 08			(500) 4929. 20	(263) 5098. 94	<u></u>	$5p' ^2\mathrm{D}_5^{\circ}$ 78511.6 $5p' ^2\mathrm{D}_3^{\circ}$ 78676.6	$\frac{1}{10}$ 4847	7. 24 521	266 bl) 11. 48 452) 46. 43	

Corona, and by E. D. Tidwell at the Bureau. In both laboratories electrodeless discharge lamps, containing pure bromine and excited by 2450megacycle microwave diathermy units, served as light source. The records made at Corona were obtained by scanning the spectrum with the highresolution grating spectrometer designed and built at the National Bureau of Standards. This apparatus has been described by Humphreys and Kostkow-The bromine wavelengths were measured on these records by linear interpolation between krypton standards that were introduced by reflection from back of the various filters used to sort out the higher order bromine lines. This procedure does not yield the most precise wavelengths, but it was entirely satisfactory for our purposes. For nearly all lines, the departure of the measured wavelengths from those calculated by the combination principle was less than 1 Å. All lines (except 19317 Å and 21093 Å) of intensity 40 and above in the region beyond 19000 Å were also measured on the tracings made by Tidwell, who used the infrared spectrometer at NBS. These wavelengths could be measured with precision by interpolation between the white-light fringes of a Fabry-Perot interferometer that were recorded simultaneously with the bromine spectrum. This technique, as well as the spectrometer apparatus, has been described by Plyler, Blaine, and Tidwell [12].

With the exception of a single chlorine line,³ every line established from these records can be classified as a transition between Br I levels whose

values have been accurately determined from observations in the photographic air region. Precise wavelengths for these lines can therefore be calculated by use of the combination principle. It is these calculated wavelengths that are listed in table A3, which includes four lines above 12000 Å that were also observed photographically.

The relative intensities given in the second column of this table express, in tenths of an inch, the heights of the recorded line profiles above the continuum. One of the tracings furnished by Humphreys was selected to provide the standard scale to which the intensities of lines measured on other tracings would be referred. As the spectrum was being scanned, lines of intensity greater than 85 were attenuated by manual adjustment of a resistance network introduced in the input of the amplifier. The measured heights of these lines were multiplied by the known

linear within the limits of loading, and the response of the detector is almost flat within the range covered, the relative intensities are considered to be very reliable over intervals of several hundred angstroms, but not over the entire region covered, inasmuch as no correction was made for the spectral response of the detector or angular distribution of energy from

attenuation factors corresponding to the various resistance box dial settings. Since the amplifier is

In table 1 are displayed the wave numbers and intensities (in parentheses) of certain important multiplets observed in the radiometric region. Observed multiplets that helped to establish the important (3P)6s 4P levels are shown in table 2. In both these tables, transitions allowed by the selection rules, but unobserved, are indicated by horizontal

bars.

the grating.

 $^{^3}$ 15730.1 Å. This is the only impurity line found on the tracings; it occurred with intensity 5 on our scale. According to the recent work of Humphreys and Paul [13], it is the second strongest Cl line in the region studied. The strongest Cl line, at 15869.7 Å, was masked by a Br line of another order.

Table 2. Some observed multiplets involving the 6s 4P levels

$^{4}P_{5/2} \\82236. \ 17$	$6s\atop {}^{4}\mathrm{P}_{3/2}\\82661.\ 57$	$^{4}P_{1/2}_{85576.93}$
(1700)	(20)	
7563, 85	7989, 25	
	(110)	(500d)a
		10567, 80
		20001100
	6847. 57	-
(1250)		
0111.01	(1800)	
	0001. 02	$(10w)^{a}$
		8833, 85
		(55)
		6711. 21
		0.11.21
		(24)
		6398, 60
	82236. 17	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$

^a These two lines were observed photographically.

3. Term Structure of BrI

The unexcited state of the neutral bromine atom is characterized by the electron configuration $4s^2 4p^5$, which yields an inverted ²P term of odd parity. From the vacuum ultraviolet observations discussed above, the position of $4s^24p^5{}^2P_{1/2}^{\circ}$ is found to be 3685 K, relative to ${}^2P_{3/2}^{\circ}$ taken as zero. The $4s4p^6$ configuration gives one even level, 2S1/2, while all other known Bri levels arise from excited configurations of the type $4s^2 4p^4 nl$, where l stands for the letters s, p, d, f indicating the orbital angular momentum quantum number, and n is the principal quantum number. It is convenient to refer to $4s^2 4p^4 nl$ configurations simply as nl configurations. The terms that are expected theoretically for each of these configurations are displayed in table 3, built in the LS scheme. The nl levels are based on, and can be derived from, the parent ${}^{3}P_{2,1,0}$ ${}^{1}D_{2}$ and $^{1}S_{0}$ terms of the $4p^{4}$ configuration of the ion, Br⁺. The observed distribution of the levels in each of the Bri nl configurations indicates that the coupling energies of the $4p^4$ core dominate the structure of the Br I levels. With just a few exceptions the coordination of Br I levels to specific components of the limiting $4p^4$ terms is unambiguous. When discussing Br I levels in text, this limit is stated explicitly. However, in our tables of observed lines, the levels involved in the classification of the lines are written in an abbreviated notation, which affixes no prime, one prime, and two primes to the symbol for l-value to designate that the level is based on the 3P_2 , 3P_1 , and 3P_0 parent, respectively. For levels based on 1D_2 and 3S_0 , the parent is stated explicitly.

The positions of the Br II $4p^4$ levels have been accurately determined by Martin and Tech [14], who established the $4p^4$ $^3P_{2,1,0}$ levels from new observations of the vacuum ultraviolet spectrum and the 1D_2 and 1S_0 levels from observations in the air region of the "forbidden" magnetic dipole transitions, $4p^4$ $^3P_2-4p^4$ 1D_2 and $4p^4$ $^3P_1-4p^4$ 1S_0 . Since the distribution of these parent levels manifests itself in the structure of all the Br I configurations, the positions of the observed parent levels are repeated here in table 4, together with a slightly different fit to intermediate coupling theory from that given earlier.

The spin-orbit interaction energies in the $4p^4$ configuration may be obtained by reversing the sign of ζ_p in the matrices for p^2 given in *The Theory*

Table 4. Observed and calculated positions for the energy levels of the 4p4 configuration in Br II

En	ergies ar	nd counling	parameters	are stated	in kaysers

Level	Observed position	ObsCalc.	ObsCalc.
		$F_2 = 1690 \\ \zeta_{4p} = 2800$	$F_2 = 1698.1 \\ \xi_{4p} = 2795.0$
$^{^{3}P_{2}}_{^{^{3}P_{1}}}$ $^{^{3}P_{0}}_{^{^{1}}D_{2}}$	$\begin{array}{c} 0.0 \\ 3136.4 \\ 3837.5 \\ 12089.1 \\ 27867.1 \end{array}$	$\begin{array}{c} 0 \\ +6 \\ -18 \\ -111 \\ +112 \end{array}$	$\begin{array}{c} 0 \\ +13 \\ -14 \\ -153 \\ 0 \end{array}$

Table 3. Predicted terms of Br i

Configuration			Predicted term	ns	
$rac{4s^2}{4s}rac{4p^5}{4s}$	² P°				
	$ns \\ n \ge 5$	$np \\ n \ge 5$	$nd \\ n \ge 4$	$nf \ge 4$	$ng \atop n \ge 5$
$\begin{array}{c} 4s^2 \ 4p^4(^3\mathrm{P})nl \\ 4s^2 \ 4p^4(^1\mathrm{D})nl \\ 4s^2 \ 4p^4(^1\mathrm{S})nl \end{array}$	{	⁴ (SPD)° ² (SPD)° ² (PDF)° ² P°	$^{4}(\mathrm{PDF}) \\ ^{2}(\mathrm{PDF}) \\ ^{2}(\mathrm{SPDFG}) \\ ^{2}\mathrm{D}$	⁴ (DFG)° ² (DFG)° ² (PDFGH)° ² F°	⁴ (FGH) ² (FGH) ² (DFGHI) ² G

Parent	LS Designation	$\begin{bmatrix} \operatorname{J}_{c}l \\ \operatorname{Designation} \end{bmatrix}$	E	${T \choose n^*}$	E	${T \choose n^*}$	E	$\begin{cases} \mathbf{T}_{n^*} \end{cases}$	E	$ \begin{cases} T \\ n^* \end{cases}$	E	$\begin{cases} T \\ n^* \end{cases}$
	$\begin{bmatrix} 4s^24p^4 + \\ ^4D_{7/2}^2 \end{bmatrix}$	7/2	5p 75521. 50	19763. 30 2. 35638	6 <i>p</i> 85762. 98	9521. 82 3. 39481	7p 89636, 84	5647. 96 4. 40788	8 <i>p</i> 91538.33	3746. 47 5. 41208	9p 92620. 21	2664. 59 6. 41742
(07)	⁴ D ^o _{5/2} ⁴ P ^o _{5/2}	5/2 5/2	75697. 05 74672. 32	19587, 75 2, 36692 20612, 48 2, 30733	85820, 53 85526, 82	9464. 27 3. 40512 9757. 98 3. 35348	89669. 07 89548. 83	5615. 73 4. 42051 5735. 97 4. 37393	91557. 11 91489. 10	3727. 69 5. 42570 3795. 70 5. 37687	92626, 84 92584, 19	2657. 96 6. 42542 2700. 61 6. 37448
(3P ₂)	4P ³ / ₂ 4D ³ / ₂	3/2	75009. 13 76743. 08	20275. 67 2. 32642 18541. 72 2. 43277	85586. 30 85943. 86	9698. 50 3. 36375 9340. 94 3. 42752	89540. 42 89786. 18	5744. 38 4. 37073 5498. 62 4. 46734	91529. 22 91667. 59	3755. 58 5. 40552 3617. 21 5. 50794	92594. 47 92677. 31	2690. 33 6. 38665 2607. 49 6. 48731
	⁴ P _{1/2}	1/2	75814.00	19470. 80 2. 37402	85799. 19	9485. 61 3. 40129	89764. 81	5519. 99 4. 45868	91598. 86	3685. 94 5. 45634	92608. 84	2675. 96 6. 40377
	$\left\{\begin{array}{c} ^2\mathrm{D}_{5/2}^{\circ} \end{array}\right.$	5/2	78511.60	19909. 60 2. 34771	88848. 01	9573. 19 3. 38569	92755. 81	5665. 39 4. 40110			Parent	J_{cl} Designation
	$^2\mathrm{D}^{\circ}_{3/2}$	3/2	78676.65	19744. 55 2. 35750	88953. 76	9467. 44 3. 40455	92818. 85	5602. 35 4. 42579	,			11/2
(3P ₁)	4S3/2	3/2	79178. 33	19242. 87 2. 38804	89031. 84	9389. 36 3. 41867	92839. 36	5581. 84 4. 43391				[5]° 9/2
	$^4\mathrm{D}^\circ_{1/2}$	1/2	78865.72	19555, 48 2, 36887	89147. 30	9273. 90 3. 43989	92890. 80	5530. 40 4. 45448				9/2
	\ 2S _{1/2}	[0]°—1/2	78076.00	20345. 20 2. 32244	88712. 98	9708. 22 3. 36206	92728. 53	5692. 67 4. 39054	÷			[4]°′
(3P ₀)	$\left\{\begin{array}{c} {}^{2}\mathbf{P}_{3/2}^{\circ} \end{array}\right.$	3/2	79695, 89	19426. 41 2. 37673	89947, 93	9174. 37 3. 45850	93500. 15	5622. 15 4. 41799	Y		$(^{3}P_{2})$	[3]°
	$P_{1/2}^{\circ}$	1/2	79868.03	19254, 27 2, 38733	89905, 88	9216. 42 3. 45060	93491, 41	5630. 89 4. 41456				5/2
	(² F _{7/2}	7/2	87225. 16	20148. 74 2. 33374	A							[2]0
	$^2\mathrm{F}_{5/2}^{\circ}$	5/2	87061. 23	20312. 67 2. 32430				1,				3/2
(1D ₂)	$\int_{0}^{2} D_{5/2}^{\circ}$	[2]°	88483. 42	18890. 48 2. 41021								[1]°′ 1/2
	$^2\mathrm{D}^{\circ}_{3/2}$	3/2	88438.78	18935. 12 2. 40736			5					
	$^2\mathrm{P}_{3/2}^{\circ}$	3/2 [1]°	87259. 29	20114. 61 2. 33572								[4]°
×	$^{2}\mathrm{P}_{1/2}^{\circ}$	1/2	87499.03	19874. 87 2. 34976								7/2
											$({}^{3}\mathrm{P}_{1})$	[3]° 5/2
												[2]° 3/2
											$(^{3}P_{0})$	[3] 7/2

E	${T \choose n}$										
$10p \\ 93291.38$	1993. 42 7. 41953	,									
				Ground C	Configuration	of Br 1, $4s^24p^5$		² P _{3/2} ² P _{1,2}	E 0.00 3685.24		
93339. 31	1945. 49 7. 51037										
										1	I
E	$\{ _{n^{\ast }}^{T}$	E	$\left\{ _{n^{\ast }}^{T}\right.$	E	$egin{cases} T \\ n^* \end{cases}$	E	$\left\{ _{n^{\ast }}^{T}\right.$	E	$\left\{ _{n^{\ast }}^{T}\right.$	E	$\left\{ _{n}^{T}\right\}$
4f		5 <i>f</i>		6 <i>f</i>		7 <i>f</i>		8 <i>f</i>		9 <i>f</i>	
88392.18	6892. 62	90873.70	4411.10 4.98772	92222.74	3062.06 5.98644	93036.08	2248.72 6.98567	93563.77	$\begin{array}{c} 1721.03 \\ 7.98512 \end{array}$	93925, 35	1359. 45 8. 98450
	3. 99010	90873.78	4411.02 4.98777	92222, 80	3062.00 5.98650	93036.13	2248. 67 6. 98575	93563. 82	1720. 98 7. 98523	93925. 36	1359. 44 8. 98453
88339.16	6945. 64 3. 97484	90845. 55	4439. 25 4. 97188	92207.32	3077. 48 5. 97143	93026.10	2258. 70 6. 97022	93556. 94	1727. 86 7. 96932	93919.84	1364. 96 8. 96635
88340.71	6944. 09 3. 97528	90846.78	4438. 02 4. 97257	92208. 27	3076. 53 5. 97235	93026. 67	2258. 13 6. 97110	93557.34	1727. 46 7. 97025		
88351.10	6933. 70 3. 97826	90851.30	4433.50	92187. 93	3096. 87	93028.44	2256. 36	93558. 51	1726. 29		
88350.78	6934. 02 3. 97817)	4. 97511		5. 95270		6. 97383		7. 97295		
88371.32	6913. 48 3. 98407	90869. 64	4415.16 4.98543	92222.01	3062.79 5.98573	93035. 35	2249. 45 6. 98453				
88363.97	6920. 83 3. 98196	90868.76	4416.04 4.98493	92221.57	3063. 23 5. 98530	93034, 99	2249. 81 6. 98398	93562. 65	1722. 15 7. 98252		
88425.48	6859.32 3.99977	90891. 56	4393. 24 4. 99785	92233.08	3051.72 5.99658	93042.65	2242. 15 6. 99590				
88424.88	6859. 92 3. 99960	90890. 88	4393. 92 4. 99746	92232. 69	3052.11 5.99619	93042.39	2242. 41 6. 99549				
91496. 23	6924. 97 3. 98077	93994. 02	4427.18 4.97866								
91497.14	6924.06 3.98103	93994. 73	4426. 47 4. 97905							10 <i>f</i>	
91545.72	6875.48 3.99507	94017. 80	4403.40 4.99208					(³ P ₂) 10f	$[4]_{9/2}$	94180. 82	1103. 98 9. 97000
91542.44	6878.76 3.99412	94018. 21	4402.99 4.99231			,					
91483.01	6938.19 3.97697	93988. 36	4432. 84 4. 97548								
91482.30	6938. 90 3. 97677	93987. 61	4433, 59 4, 97506		,					. ,	
92228.30	6894.00 3.98970	94700. 23	4422.07 4.98153								
92228.60	6893. 70 3. 98978	94700. 41	4421. 89 4. 98163	A						*	

Parent	LS	$J_c l$	T.		-		1		1	TABLI	$E 6. Ev_0$
- arcin	Designation	Designation	<i>E</i>	$\begin{cases} T \\ n^* \end{cases}$	<i>E</i>	$\left\{ \begin{smallmatrix} T \\ n* \end{smallmatrix} \right.$	E	$\begin{cases} T \\ n^* \end{cases}$	E	$\begin{cases} T \\ n^* \end{cases}$	E
(3P ₂)	$\left\{\begin{array}{c} 4s^2 4p^4 + \\ {}^4\mathrm{P}_{5/2} \end{array}\right.$	5/2	58 63436. 45	31848. 35 1. 85623	6 <i>s</i> 82236, 17	13048. 63 2. 89997	7 <i>s</i> 88135. 07	7149. 73 3. 91770	8 <i>s</i> 90788. 90	4495. 90 4. 94046	98 92161. 45
	${ m P}_{3/2}$	3/2	64907. 19	30377. 61 1. 90064	82661.57	12623. 23 2. 94843	88275. 49	7009. 31 3. 95674	90797. 55	4487. 25 4. 94522	92182. 30
(3P ₁)	${ m P}_{3/2}$	[1]	67183. 58	31237. 62 1. 87429	85435.33	12985. 87 2. 90697	91284. 03	7137. 17 3. 92114	93906. 64	4514. 56 4. 93024	
	$^4\mathrm{P}_{1/2}$	1/2	66883. 87	31537. 33 1. 86536	85576.93	12844. 27 2. 92295	91347. 68	7073. 52 3. 93875	93983. 18	4438. 02 4. 97257	
(3P ₀)	$\left\{ \begin{array}{cc} {}^2P_{1/2} \end{array} \right.$	[0]—1/2	68970. 21	30152. 09 1. 90773	86279.41	12842. 89 2. 92310	92089. 61	7032. 69 3. 95016			
	$\begin{cases} 4s^2 & 4p^4 + \\ & ^4F_{9/2} \end{cases}$	9/2	4d 80587. 61	14697. 19 2. 73249	5d 87191. 97	8092. 83 3. 68235	6d 90244. 89	5039. 91 4. 66621	7d 91851. 14	3433. 66 5, 65323	8d 92807. 81
	4F _{7/2}	7/2	81081. 10	14203.70 2.77955	87440.02	7844. 78 3. 74012	90267. 85	5016. 95 4. 67688	91880. 64	3404. 16 5. 67768	92843. 81
	4D _{7/2}	[3]	79044. 50	16240. 30 2. 59943	86768. 28	8516. 52 3. 58959	89860.70	5424. 10 4. 49792	91755. 54	3529. 26 5. 57614	92729. 66
	$^{4}{ m D}_{5/2}$	5/2	79260. 91	16023. 89 2. 61693	86933. 28	8351. 52 3. 62487	89947. 97	5336. 83 4. 53455	91810. 29	3474. 51 5. 61990	92753. 69
(3P ₂)	${}^4{ m F}_{5/2}$	[2]	81672. 28	13612. 52 2. 83927	87754.92	7529. 88 3. 81752	90354. 93	4929. 87 4. 71800	91956. 55	3328. 25 5. 74206	92743, 16
	⁴ F _{3/2}	3/2	81842. 02	13442. 78 2. 85714	87769.96	7514. 84 3. 82134	90416. 44	4868. 36 4. 74771	91926. 88	3357. 92 5. 71664	92853, 62
	⁴ D _{3/2}	[1]	79630. 67	15654. 13 2. 64765	87131.65	8153. 15 3. 66871	89867. 69	5417. 11 4. 50082	91800. 15	3484. 65 5. 61172	92774. 91
	$^{4}\mathrm{D}_{1/2}$	1/2	80026.40	15258. 40 2. 68177	87271.73	8013. 07 3. 70063	89729. 40	5555. 40 4. 44445	91690. 94	3593. 86 5. 52580	92650. 4
	$4P_{1/2}$	[0]—1/2	81429. 97	13854. 83 2. 81433	87754.17	7530. 63 3. 81733	90428. 50	4856. 30 4. 75360	91823. 95	3460. 85 5. 63098	92824. 33
	$^{2}\mathrm{F}_{7/2}$	[3]	83358. 84	15062. 36 2. 69916	90427. 64	7993. 56 3. 70515	93244. 35	5176. 85 4. 60408			
	${}^2{ m F}_{5/2}$	5/2	83723. 08	14698. 12 2. 73240	90565.85	7855. 35 3. 73760	93342. 44	5078. 76 4. 64833			
(3P ₁)	${}^4{ m P}_{5/2}$	[2]	84305.76	14115. 44 2. 78823	90585, 26	7835, 94 3, 74223	93473. 77	4947. 43 4. 70962		,	
	⁴ P _{3/2}	3/2	84077. 52	14343. 68 2. 76596	90537. 52	7883. 68 3. 73088	93452. 77	4968. 43 4. 69966			
	² P _{3/2}	3/2	83578. 50	14842. 70 2. 71906	90459.75	7961. 45 3. 71261	93117. 18	5304. 02 4. 54855			,
	$^{2}\mathrm{P}_{1/2}$	1/2	83101.09	15320. 11 2. 67636	90292.34	8128. 86 3. 67418	92992. 67	5428. 53 4. 49609			
(3P ₀)	$\left\{ -{}^{2}D_{5/2}\right.$	5/2	84945. 16	14177. 14 2. 78216	91175.18	7947. 12 3. 71596	94088. 98	5033.32 4.66926	,		
	$^{2}D_{3/2}$	3/2	85191, 59	13930. 71 2. 80666	91001.35	8120. 95 3. 67597	94082.60	5039. 70 4. 66631			

{;	T_{n^*}	E	${T \choose n^*}$	E	$\left\{ _{n^{\ast }}^{T}\right.$	E	$\left\{ _{n^{\ast }}^{T}\right.$		Parent	LS Designation	E	${T \choose n^*}$
3128 5. 92	3. 35 2741	10s 92995. 81	2288. 99 6. 92395	11 <i>s</i> 93544. 55	1740. 25 7. 94090			4s ² 4p ⁴ + 5s	(1D ₂)	${ m 2D_{5/2}}$	75890. 33	31483. 57 1. 86695
3102 5. 94	2. 50 4730	93039. 29	2245, 51 6, 99066	93557. 98	1726. 82 7. 97172		,			$2D_{3/2}$	75908. 53	31465. 37 1. 86749
								58	(1S ₀)	$^{2}\mathrm{S}_{1/2}$	91937. 95	31213. 95 1. 87500
		$4s4p^6$	$^{2}\mathrm{S}_{1/2}$	84825. 60				68	$(^{1}D_{2})$	$\left\{ \begin{array}{c} ^{2}\mathrm{D}_{5/2} \end{array} \right.$		
										$^{2}D_{3/2}$	94470. 7	12903. 2 2. 91626
2476 6. 65	5. 99 5600	9d 93411. 48	1873. 32 7. 65367	10 <i>d</i> 93818. 96	1465. 84 8. 65232	11 <i>d</i> 94106. 70	1178. 10 9. 65127	12 <i>d</i> 94317. 40	967. 40 10. 65056	$\begin{bmatrix} 4s^2 4p^4 + \\ {}^2G_{9/2} \end{bmatrix}$	4 <i>d</i> 91921. 86	15452. 04 2. 66491
2440 6. 70		93431. 26	1853. 54 7. 69440	93859. 09	1425. 71 8. 77324	94126. 43	1158. 37 9. 73312			² G _{7/2}	91998. 55	15375. 35 2. 67155
2555 6. 55	5. 14	93403. 86	1880. 94 7. 63815	93798. 76	1486. 04 8. 59331					$^2\mathrm{D}_{5/2}$	93088. 62	14285. 28 2. 77160
2531 6. 58	1. 11 8446	93405. 91	1878. 89 7. 64231	93815, 57	1469. 23 8. 64233				$(^{1}\mathrm{D}_{2})$	$2D_{3/2}$	92516. 91	14856. 99 2. 71775
2541 6. 57	1. 64 7081									² F _{7/2}	93682. 82	13691. 08 2. 83111
2431 6, 71	1. 18 1842									$^2\mathrm{F}_{5/2}$	93671. 92	13701. 98 2. 82998
2509 6. 61	9. 89 1223	93335. 31	1949. 49 7. 50266	93821, 47	1463. 33 8. 65973					${}^{2}\mathrm{P}_{3/2}$	93609. 89	13764. 01 2. 82360
2634 6. 45	4. 35 5415	93282. 58	2002. 22 7. 40320							${}^{2}\mathbf{P}_{1/2}$	93650. 35	13723. 55 2. 82776
2460 6. 67	0. 47 7831	93451. 49	1833. 31 7. 73673							$^{2}S_{1/2}$	93701. 48	13672. 42 2. 83304
								×				,

of Atomic Spectra [15, p. 268]. The diagonal electrostatic energies to be added are ⁴

 $^{^{3}}P~:~6F_{0}-\!\!\!\!\!-15F_{2}\\ ^{^{1}}D~:~6F_{0}-\!\!\!\!\!-9F_{2}\\ ^{^{1}}S~:~6F_{0}$

where $F_0 = F_0(4p, 4p)$ and $F_2 = \frac{1}{25} F^2(4p, 4p)$. choice of coupling parameters in the previous calculation of the distribution of Br II $4p^4$ levels [14] resulted in the fit shown in the third column of table 4. Since these discrepancies reappear to some extent in the calculation of Br I $4p^45s$ levels presented below, it has seemed worthwhile to recalculate the parents in order to demonstrate that, with just a small change in parameters, the calculation of Br II 4p4 1S0 can be improved without appreciably affecting the fit for the ${}^{3}P_{2,1,0}$ levels. The new fit, shown in the last column of table 4, is a strict least squares fit to the four levels other than ¹D₂. It predicts the total energy spread of the configuration exactly, and emphasizes that the larger discrepancy for $4p^4$ ¹D₂ may actually be due to a perturbation by a higher level, since the discrepancy is in the right direction.

Tables 5 and 6 contain all the energy levels of the neutral bromine atom that we have been able to find by using the new wavelengths. The tables are arranged so as to display clearly the level parentage, configuration, and series. For each level are given the energy (E) relative to $4p^5 \, ^2\mathrm{P}^{\circ}_{3/2}$ taken as zero, the absolute energy value (T) measured from the principal ionization limit, i.e., from Br II $4p^4 \, ^3\mathrm{P}_2$, and the effective quantum number (n^*) calculated according to $n^* = \sqrt{R/T}$, where R is the Rydberg constant for bromine, 109736.56 K. The designations of the levels, usually in both the LS scheme and the $J_c l$ scheme, are given just to the left of the first member of each series and hold for all other levels of the series. In all except the nf configurations one might say that the coupling is "midway between" the LS, $J_c j$, and $J_c l$ coupling schemes. The almost pure pair-coupling in the nf configurations permits positive assignment of $J_c l$ symbols to the nf levels. All other levels are consistently referred to in terms of the adopted LS symbols, for reasons of tradition, of the general familiarity of this notation, and of the lack of any really satisfactory solution to the vexing problem of notation for intermediate couplings.

Figure 2 shows the distribution of ns, np, nd, nf levels through n=9. The small horizontal bars in each nl column represent the position of one or more levels. The scale of the diagram does not allow an individual representation for each level. The vertical lines in each column separate groups of levels having different parents, the lowest group of levels in each column being based on the $(^3P_2)$ parent. It should be remembered that when a level group appears to be intersected by a vertical line (by virtue

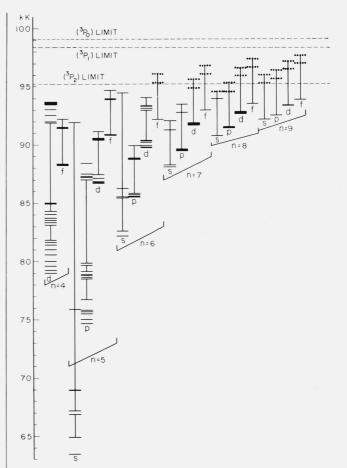


Figure 2. The distribution of Br1 energy levels for n=4 to n=9, as discussed in the text.

of being met from above and from below by such a line), then there is a change of parent in either going to the next higher or the next lower group. The separation of the level groups for each nl clearly reflects the energy differences of the parent limits shown by the dashed lines in the figure. Thus, except for absolute energy position, the levels of many nl sets have the same appearance in this figure: three groups of levels held apart by the parent differences. This "building-block" character is well observed at 4f, 5f, 6p, and 7s, for example, and can be expected to hold at all high n. The dotted level positions represent some predicted, but unobserved, levels based on (${}^{3}P_{1,0}$). It is seen that the only unobserved levels based on parents other than (${}^{3}P_{2}$) and lying below the $(^{3}P_{2})$ limit are those of $(^{3}P_{1})$ 8p, $(^{3}P_{1})7d$, and $(^{3}P_{0})8s$. The other dotted positions fall above the (³P₂) limit, and the atoms excited to these levels can be expected to autoionize.

It is convenient at this point to discuss the extent to which significance may be attached to the symbols used in this paper to describe the energy levels. The LS symbols are based on an examination of the intensities in the transition arrays, with symbols being assigned in such a way that the

 $^{^4}$ These electrostatic expressions differ by a constant amount from those given on p. 199 in TAS, which are not entirely correct. See footnote on page 102 of reference [22].

intensity relations most nearly conform to the wellknown LS rules. The arrays used were those containing levels of low principal quantum number n, the assignments being retained along series without regard to intensities. LS symbols assigned to levels in this way will not necessarily survive if the actual energy matrices in the LS scheme are diagonalized and the eigenvectors found, because when the coupling strongly departs from the LS scheme, the intensities then being governed more by, say, the J_{cl} rules, the LS symbols to which one is led depend more on J_c and K and will not with assurance correspond to the major LS contribution to the composition of the levels. A preliminary rough diagonalization of the Br I 5p energy matrices suggests, for example, that the "percentage-method" would require the $(^3P_2)5p \, ^4D_{5/2}^{\circ}$ and $(^3P_1)5p \, ^2D_{5/2}^{\circ}$ levels to have their designations interchanged, whereas the intensities would not suggest this. In any case, the departure from LS coupling is so great that L and S are not good quantum numbers anyway, making meaningless any involved procedure for naming the levels. In the absence of reliable eigenvectors, the "intensity-method" seems to be the most attractive. The np LS assignments in Br I turn out to correspond exactly to those in F I [16], but are quite different from those adopted for the I1 np levels by Kiess and Corliss [1], who assigned LS symbols by forming a one-to-one best correspondence of Landé g-factors to observed g-factors. For small departures from pure LS coupling, all three methods should result in the same designations.

Ideally, one should like to designate observed levels in the notation of the coupling scheme in which the energy matrices are already most diagonal, that is, in which there is the least nondiagonal contribution to the energies (eigenvalues). In this sense, the Br i nf configurations are very nearly pair-coupled, and J_{cl} symbols should certainly be used for the *nf* levels. For the sake of consistency, J_{cl} symbols have also been assigned to the other Br I levels, as alternatives to the $\overline{L}S$ symbols. is justifiable on the grounds that for ns the choice of $J_c l$ or $J_c j$ notation is arbitrary owing to the identity of the matrices built in the two schemes, while for np and nd the notation could be equally well J_cl or J_{cj} since the coupling is intermediate. In these cases only exact calculation would allow us to decide in favor of one or the other scheme, and this would still be difficult for nd because of the sizeable perturbations that afflict the *nd* configurations. J_{cl} designations for the Br 1 np levels have been chosen on the basis of approximate calculations by using the diagonal $J_c l$ energy expressions. Diagonal expressions are less valid for nd, so that only general considerations of level positions could be invoked in the assignment of $J_c \hat{l}$ notation, the primary emphasis being placed on the intensities of $(^3P_2)4d-(^3P_2)nf$ transitions for the $(^3P_2)4d$ designations and 5p-4d, 5p-5d transitions for the other nd designations.

The main features of $J_c l$ coupling and the theoretical treatment of electronic configurations exhibiting this type of coupling are discussed at length by Minnhagen [17] in his fine paper on Ar II. The theoretical

treatment is based upon the general formulas of Möller [18], who calculated the p^4l matrices of electrostatic energy in the J_d scheme. The corresponding matrices of spin-orbit interaction have recently been provided by Källén and appear in the appendix of reference [19].

A pair-, or $J_c l$ coupled configuration is characterized by (i) a domination by the core coupling energies, and (ii) an electrostatic interaction of the outer electron with the core that is stronger than the spin coupling of this electron. The second requirement means, in the case of the $4p^4nl$ configurations of Br I, that the contribution of $F^2(4p, nl)$ to the level energies must be greater than that of ζ_{nl} and the $G^k(4p, nl)$. The notation used for a level in a paircoupled configuration is that suggested by Racah [20], in which the level is designated as ()nl $[K]_J$, where the parent is placed in the parentheses, and where K, the intermediate quantum number, is the quantized resultant of the coupling of the total angular momentum J_c of the parent ion with the orbital angular momentum of the external electron. The coupling of K with the spin of this electron gives the J-value of the level. Each K-value thus has two J-values associated with it, $J=K\pm\frac{1}{2}$. Br I, J_c can take the values 2, 2, 1, 0, and 0, corresponding to the five levels of the ground configuration $4p^4$ in the ion. The selection rules governing electric dipole transitions between J_cl coupled levels are

$$\Delta J_c=0$$
; $\Delta K=0,\pm 1$; $\Delta l=\pm 1$; and $\Delta J=0,\pm 1$.

The last two are, of course, independent of coupling, but the first two are easily violated when departures from pure pair coupling occur, because the energy matrices are not diagonal in J_c or K.

Owing to the departures in BrI from any pure coupling scheme, the single level eigenfunctions of any such scheme are mixed. L, S, J_c , and K are not very good quantum numbers, and the only restrictions on transitions seem to be those on parity and ΔJ .

4. Ionization Limits

A very sharp determination of the principal ionization limit of the neutral bromine atom is possible now that the $(^{3}P_{2})nf[5]_{11/2}^{\circ}$ series has been observed. This series is found not only empirically to provide the best value for the limit; it is also the series one might expect on theoretical grounds to provide such a best value. The theoretical criterion one should use in the selection of the least perturbed series for the limit determination is that the effect of nondiagonal matrix elements connecting the levels to other levels of common l and J value shall be very small or zero. In general, this criterion restricts consideration to series of levels with high l and J values.

Since no level with J=11/2 can occur in the np configurations, there will be no perturbation of the $(^3P_2)nf[5]_{11/2}^{\circ}$ series by levels of those configurations. Indeed, the only nondiagonal element that makes a connection to this series is the ζ_{4p} element between

 $(^{3}P_{2})[5]_{11/2}^{\circ}$ and $(^{1}D_{2})[5]_{11/2}^{\circ}$, but the latter nf series is well above the $(^{3}P_{2})$ ionization limit, the lowest member of the series, $(^{1}D_{2})4f[5]_{11/2}^{\circ}$, being already more than 5000 K above the $(^{3}P_{2})$ limit.

The actual determination of the limit from the $({}^{3}P_{2})nf[5]_{11/2}^{\circ}$ series has been carried out graphically. The effective (principal) quantum number, n^{*} , of a series member is defined by the relation $T=R/(n^{*})^{2}$, where T is the position of the level measured from the series limit and R is the Rydberg constant, which for bromine equals 109736.56 K. If the series is Ritzian, then $n^{*}=n-\alpha-\beta T_{n}$, where n is the running principal quantum number and α and β are constants for a given series. The position of the series limit is taken as that value for which the plot of $(n-n^{*})$ versus T_{n} yields the best straight line. The quantity $(n-n^{*})$ defines the quantum defect, $\delta=\alpha+\beta T_{n}$. In practice it is usually more convenient to plot T_{n} against $(C-\delta)$, where C is some integer for the series such that $(C-\delta)$ is the part of n^{*} to the right of the decimal.

Figure 3 shows that a very nearly straight line results if the $4p^4(^3P_2)$ limit is taken to be 95284.80 K above the $4p^5(^3P_3)$ ground level of Br I. The vertical lines through the points represent the deviation of $(1-\delta)$ resulting from a change in the limit of ± 0.10 K. If the adopted series limit is reduced by only one kayser, we obtain the unsatisfactory and unreasonable behavior of this series depicted by the second plot in figure 3.

With the adopted IP=95284.80 K, we have made a least squares calculation of the constants α and β by using the first five series members. How well these constants reproduce the series members can be seen from table 7. The position of $(^3P_2)9f[5]^{\alpha}_{1/2}$ would have to be raised by 0.09 K in order to bring its point in figure 3 onto the line. This is certainly within the experimental error, since the level has been established on the basis of two measurements of a wide, hazy line at 7495.46 Å. A change of 0.05 Å

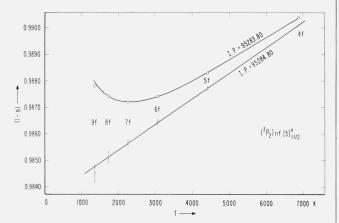


Figure 3. The $(^3P_2)nf[5]_{11/2}^{\circ}$ series plotted for two different values of the limit.

The upper curve results from changing the value of the BrI ionization potential by only 1.00 K from the value adopted for BrI on the basis of the lower plot. The vertical bars in the lower plot show the effect on $(1-\delta)$ of a change of \pm 0.1 K in the limit.

in this wavelength would bring the corresponding level into line.

The value of 95284.80 K for the principal ionization energy is well supported by other series. The $({}^{3}P_{2})nf[5]_{9/2}^{6}$ series gives an identical limit, although the point for 9f again falls about 0.1 K below the line, and again the spectral line establishing the level is wide and faint (7783.44 Å). With this limit, the plots of the higher members of the $({}^{3}P_{2})nd$ ${}^{4}F_{9/2}$ and $({}^{3}P_{2})np$ ${}^{4}D_{7/2}^{6}$ series are also very nearly linear, while the greater deviation from linearity in the first few members of each of these series is a result of the changing coupling conditions along the series.

Table 7. Observed Br i $(^3P_2)$ nf $[5]_{11/2}^{\circ}$ series members $\overline{^*}$ compared with positions calculated by use of the Ritz formula

 $\alpha = 0.016505$ $\beta = -9.585 \times 10^{-7}$

Config.	Observed	Calc.	Obscalc
$\begin{array}{c} 4f \\ 5f \\ 6f \\ 7f \end{array}$	88392. 18 90873. 70 92222. 74 93036. 08	88392. 19 90873. 70 92222. 73 93036. 07	$ \begin{array}{c} -0.01 \\ 0.00 \\ +0.01 \\ +0.01 \end{array} $
8 <i>f</i> 9 <i>f</i>	93563. 77 93925. 35	93563. 78 93925. 44	-0.01 -0.09

It may be pointed out that, although the apparent depression of $({}^{3}P_{2})9f[5]_{11/2}^{\circ}$ and $({}^{3}P_{2})9f[5]_{9/2}^{\circ}$ is within the experimental uncertainty, the depression of these levels might more interestingly be a result of an inter-ionic Stark effect. The conditions at 9f are ripe for such an effect, and the character of the observed lines from these levels are certainly suggestive of Stark broadening. Such an interpretation is strengthened by the fact that the lines (³P₂)4d $^4F_{9/2}-(^3P_2)7f[5]_{11/2}^\circ$ and $(^3P_2)4d$ $^4F_{7/2}-(^3P_2)7f[5]_{9/2}^\circ$, as well as 9 out of 13 of the observed $(^3P_2)4d-(^3P_2)8f$ lines have been recorded as hazy or wide. regularity of the $({}^{3}P_{2})nf[5]_{11/2}^{\circ}$ series used to determine the ionization limit leaves no doubt, however, that the determination is an excellent one, accurate to probably ± 0.15 K relative to the levels used in the determination, even allowing a small Stark effect on the n=7, 8 members. Such an effect would have no more influence than the normal experimental error on the limit determination.

It seems reasonable therefore to adopt 95284.8 $\pm 0.5 \, \mathrm{K}$ as the position of Br II $4p^{43} \mathrm{P}_2$ above the ground level of Br I. We have included in the stated error the estimated accuracy of the absolute positions of the higher levels. With the conversion factor $12395 \times 10^{-8} \, \mathrm{ev/K}$ as adopted in AEL, we obtain 11.811 ev for the principal ionization energy of Br I. If we use a more recent value [25] for the conversion factor, $8066.03 \, \pm 0.14 \, \mathrm{K/ev}$, we obtain 11.8131 $\pm 0.0003 \, \mathrm{ev}$ for this ionization energy.

Although several Br1 series having three members and based on the higher (${}^3P_{1,0}$) are known, none of them is unperturbed. More accurate values for the

them is unperturbed. More accurate values for the positions of the higher limits can be found by com-

bining the above (³P₂) limit with the positions of the Br II $4p^4$ levels as given by Martin and Tech [14]. It is estimated that these Br II levels are accurate to $\pm 0.5 \,\mathrm{K}$. The five limits are then as follows:

> $(^{3}P_{2})$: 95284.8 $(^{3}P_{1})$: 98421.2 $(^{3}P_{0})$: 99122.3 $(^{1}D_{2}) : 107373.9$ $(^{1}S_{0})$: 123151.9

These are the limits that were used in the calculations

of n^* for the Br I levels.

The (³P₂) limit found in this investigation is 265 K lower than that given in AEL, Vol. II. It is noteworthy that Catalán and Rico predicted a value of 95300 K for this limit on the basis of their study of the Ga_I-Rb_I sequence [21].

5. Levels of Even Parity

(a) $4p^4ns$: Except for $(^3P_0)$ $8s\,^2P_{1/2}$, which is calculated to fall around 94650 K, every $(^3P)\,ns$ level through 11s that is predicted to lie below the (3P₂) limit is now known. The (3P)5s group and the two (³P₂)78 levels were known before. With the exception of $({}^{3}P_{2})8s {}^{4}P_{5/2}$, assigned to the 6d configuration in AEL, all other ns levels are new. The (3P)7s

 $^{2}\mathrm{P}_{3/2,\,1/2}$ levels listed in AEL are spurious. The most important of the new $(^{3}\mathrm{P})ns$ levels are those belonging to (3P)6s. The strongest lines from these levels fall in the lead-sulfide and vacuum regions, but it has been possible to determine all the levels accurately from weaker lines observed in the photographic air region. Catalán and Rico [21], who made a graphical study of the GaI-RbI sequence by means of term values and the constants α and β appearing in the expression defining the quantum defect, predicted the positions of the then unknown $(^{3}P_{2})6s^{4}P_{5/2}$ and $(^{3}P_{1})6s^{2}P_{3/2}$ levels as 82300 K and 85649 K, respectively, which are remarkably close to their observed positions 82236 K and 85435 K. Catalán and Rico also came to the conclusion that (${}^{3}P_{2}$)5s ${}^{2}P_{3/2}$ at 67184 K was considerably perturbed, its "unperturbed" position being 67969 K. However, the results of certain calculations, presented below, do not support such a conclusion. It seems that the differing coupling conditions in the elements of the sequence studied by Catalán and Rico are responsible for the appearance that the 67184 level was perturbed.

In both the 5s and 6s configurations the ²D term based on the (¹D₂) parent falls below the (³P₂) limit. The term belonging to 5s is discussed below. The $(^{1}D_{2})6s$ $^{2}D_{3/2}$ level, 94470.7 K, has been established on the basis of only one line, identified as $4p^5$ $^2P^{\circ}_{\cancel{1}}$ — $(^1D_2)6s^2D_{\cancel{3}/2}$, at 1101.498 Å in the vacuum ultraviolet. This line was long regarded by us as belonging to BrII, but its intensity and the fact that it establishes the (1D2)6s 2D3/2 level within just a few kaysers of its predicted position (table 11), make us believe that our identification is almost certainly correct. One then wonders whether the unclassified line at 5139 Å might not be $({}^{3}P_{2})5p {}^{4}P_{3/2}^{\circ} - ({}^{1}D_{2})6s$ $^2\mathrm{D}_{5/2}$, placing the latter level at 94460.97 K.

The only ns configuration in which the 2S1/2 level based on the (1S₀) parent falls below the (3P₂) parent is 5s. The results of calculations presented below give ~ 91940 K for the position of $({}^{1}S_{0})5s {}^{2}S_{16}$. There are three eligible levels in this vicinity that have J=1/2, at 91691, 91824, and 91938 K. (For intensity reasons and on the basis of calculations, the level at 92090 K has been assigned to (³P₀)7s ²P₁₆.) Since the 91938 level agrees so well with the predicted position, we have interpreted this level as the (1S₀)5s 2S_{1/2}. The other two levels have been assigned to the $({}^{3}P_{2})7d$ configuration. It is felt that this interpretation of the two lower levels also leads to the most consistent nd series. An observed q-factor for the level at 91938 K will be required in order convince us that this interpretation is correct, however.

The (3P)5s levels were all known before, but a curious circumstance led Kiess and de Bruin to adopt incorrect $(^{1}D_{2} \, ^{1}S_{0})5s$ and $(^{1}D_{2} \, ^{1}S_{0})5p$ levels. This caused some confusion in the preparation of AEL about the correct order of the (¹D₂) and (¹S₀) limits in Br II. The 5s levels in question, adjusted to our energy scale, are:

	This work	\$ 1 x 3 x 4	Ref. [1]
$\begin{array}{ccc} (^1D_2) & ^2D_{5/2} \\ \\ (^1D_2) & ^2D_{3/2} \\ \\ (^1S_0) & ^2S_{1/2} \end{array}$	75890. 3 (-18. 2) 75908. 5 91938. 0	77330. 8 (+18. 2) 77312. 6 75908. 6	

It is seen that Kiess and de Bruin's ²S_{1/2} level is real and corresponds to our ${}^{2}D_{3/2}$. Although their ${}^{2}D$ term is inverted with respect to ours, it has the same separation, 18.2 K. This led to a qualitatively correct interpretation of many lines in Kiess and de Bruin's work, even though most of their $(^{1}D_{2})5s,5p$ levels are spurious. Both levels of their $(^{1}S_{0})5p$ $^{2}P^{\circ}$ term are real but are actually based on the ³P parent. Since Kiess and de Bruin established the (1D2)5p levels on the basis of strong lines having the 18.2 K separation of their spurious (¹D₂)5s levels, all (except ${}^{2}P_{3/2}^{\circ}$) of their $({}^{1}D_{2})5p$ levels are spurious as well, and are all (except ²P_{1/2}) raised above real levels by an amount corresponding to the difference $\Delta = 1422.2 \text{ K}$. The reason that this fortuitous system could have appeared so convincing—the transition array is displayed in reference [4]—is that Br I is a prime example of a spectrum in which chance numerical relationships occur among the levels and give support to gross errors in the analysis. It is through such relationships that additional errors can easily be generated; in this case they also account for some of the incorrect J-values given by Kiess and de Bruin.⁵

The energy levels of the Br I $4p^45s$ configuration have been calculated by diagonalizing the second-order energy matrices built in the LS scheme. The matrices of spin-orbit interaction are obtained from those of p^2s on page 268 of TAS by reversing the sign of the spin-orbit integral associated with the core, ζ_p . The diagonal electrostatic energies to be added to the spin-orbit matrices are as follows:

 $\begin{array}{l} ^4P: F_0 {-} 15 F_2 {-} 3 G_1 \\ ^2P: F_0 {-} 15 F_2 \\ ^2D: F_0 {-} 9 F_2 {-} 2 G_1 \\ ^2S: F_0 {-} 2 G_1 \end{array}$

where $F_2 = (1/25)F^2(4p, 4p)$, $G_1 = (1/3)G^1(4p, 5s)$,

and $F_0 = 6F_0(4p, 4p) + 4F_0(4p, 5s)$.

In the first diagonalization we adopted the values ζ_{4p} and F_2 found above for the $4p^4$ configuration of parent levels. In order to compensate approximately for the discrepancy in the calculated position of the $(^1D_2)$ parent, we have added an electrostatic correction term C=-150 K to the diagonal energy expression above for the 4D term, built on $(^1D_2)$. A least squares fit, holding C constant, was then made to the 5s -levels other than $^2D_{5/2,3/2}$, since this term is known to be perturbed. The calculation yielded

 $\substack{ (^3\mathrm{P}_2)7p \ ^4\mathrm{D}_{3/2}^5 \ 89786.2 - (^3\mathrm{P}_2)4f[2]_{3/2}^3 \ 88364.0 = 1422.2 \ \mathrm{K} } \\ \substack{ (^1\mathrm{D}_2)5p \ ^2\mathrm{D}_{3/2}^5 \ 88483.4 - (^1\mathrm{P}_2)5p \ ^2\mathrm{F}_{3/2}^5 \ 87061.2 = 1422.2 \ \mathrm{K} } \\ \substack{ (^3\mathrm{P}_0)6p \ ^2\mathrm{P}_{1/2}^2 \ 89965.9 - (^1\mathrm{D}_2)5p \ ^2\mathrm{D}_{3/2}^2 \ 88483.4 = 1422.5 \ \mathrm{K} }$

Owing to the appearance of the 88483.4 level in two of the pairs, the line at 7950.2 Å falls at a "junction" in Kiess and de Bruin's system, so that it could be, and indeed was, classified twice by them. The entire puzzle is compounded by the fact that the real pairs of levels,

 $\substack{ (^3P_0)5s\ ^2P_{1/2}\ 68970.2 - (^3P_1)5s\ ^2P_{2/2}\ 67183.6 = 1786.6\ \mathrm{K} \\ (^3P_1)6p\ ^2D_{3/2}^288848.0 - (^1D_2)5p\ ^2F_{3/2}^287061.2 = 1786.8\ \mathrm{K} },$

have separations almost equal to $\Delta \!=\! 1422.2$ added to the separation

 $(^{3}\mathrm{P}_{1})6p\ ^{2}\mathrm{D}_{5/2}^{\circ}\ 88848.0 + (^{1}\mathrm{D}_{2})5p\ ^{2}\mathrm{D}_{5/2}^{\circ}\ 88483.4 = 364.6.$

 G_1 =733.8. The results of the fit are given in table 8 and show that the 2D levels have apparently been pushed down by about 300 K. It is for this reason that any attempt to determine physically meaningful parameters from the diagonal-sum rule necessarily leads to failure. In fact, this method yields an imaginary value for ζ_{4p} and a negative G_1 .

In order to obtain better eigenvectors for a calculation of g-factors, we attempted to improve the parametric representation of the 5s levels by making a strict least squares fit to the levels other than $^2\mathrm{D}_{5/2,\,3/2}$ and by allowing all parameters (except C) to vary. The main effect on the parameters was a reduction of the value of ζ_{4p} . The fit is shown in table 8. By means of the eigenfunction expansion coefficients (table 9) obtained in this calculation, the theoretical g-factors can be calculated. These are compared in table 10 with the observed q-factors. The q-factors for these 5s levels are derived from measurements of the Zeeman patterns we obtained by exposing to an electrodeless lamp placed in a field of about 37000 oersteds. Since our exposure times were not long for these preliminary exposures, only the stronger Br I lines were recorded. These included most of the lines in the 5s-5p transition array, however. Also averaged with our own measurements were values generously furnished us by W. F. Meggers, who had measured some Br I Zeeman patterns occurring on his ytterbium spectrograms. The final observed g-factors should be accurate to better than 0.015. The agreement between observed and calculated values is seen to be very good. The observed g-sums also agree well with the theoretical sums. It should be noticed that the LS levels ${}^{2}P_{3/2}$ and ${}^{4}P_{3/2}$ are thoroughly mixed.

It is gratifying that two effects predictable on the basis of the perturbation of the $(^{1}D_{2})5s$ ^{2}D levels are actually observed. The first concerns the group of moderately strong lines in the region 7975 to 8035 Å. These lines all show resolved hyperfine structure and are $(^{1}D_{2})5s$ $^{2}D - (^{3}P_{2})4f$ transitions,

Table 8. Observed and calculated Br i $4\mathrm{p}^45\mathrm{s}$ energy levels

All energies and parameters are stated in kaysers.

Level desig.	Observed position	$G_1(4p,5s)$	0 = 1698.1 0 = 733.8 0 = 2795.0	$F_2(4p, 4p) = 1698.0$ $G_1(4p, 5s) = 748.5$ $\zeta_{4p} = 2712.0$	
		Calc.	Obs-Calc.	Calc.	Obs-Calc.
$ \begin{array}{c} (^3P_2) 5s ^4P_{5/2} \\ (^3P_2) 5s ^4P_{3/2} \\ (^3P_1) 5s ^4P_{1/2} \\ (^3P_1) 5s ^2P_{3/2} \\ (^3P_0) 5s ^2P_{1/2} \\ \end{array} \\ \begin{array}{c} (^1D_2) 5s ^2D_{5/2} \\ (^1D_2) 5s ^2D_{3/2} \\ \end{array} \\ (^1S_0) 5s ^2S_{1/2} \\ \end{array}$	63436 64907 66884 67184 68970 75890 75909	63389 64848 66938 67185 68984 76185 76239	$ \begin{array}{r} +47 \\ +59 \\ -54 \\ -1 \\ -14 \end{array} $ $ \begin{array}{r} -295 \\ -330 \\ -37 \end{array} $	63439 64905 66888 67173 68977 76172 76225	$ \begin{array}{r} -3 \\ +2 \\ -4 \\ +11 \\ -7 \\ -282 \\ -316 \\ 0 \end{array} $

 $^{^5}$ The difficulty rests on the fact that three pairs of real odd levels in Br I have almost the same separation, equal to the above difference Δ . They are

Table 9. Eigenvectors of the Br I 4p⁴5s levels, calculated with $F_2(4p, 4p) = 1698.0$, $G_1(4p, 5s) = 748.5$, $\zeta_{4p} = 2712$

	$(^{1}D_{2})$ $^{2}D_{5/2}$	$(^3P_2)$ $^4P_{5/2}$	
$^{(1}\mathrm{D}_{2})~^{2}\mathrm{D}_{5/2} \\ (^{3}\mathrm{P}_{2})~^{4}\mathrm{P}_{5/2}$	$ \begin{array}{c c} 0.9883 \\ -0.1524 \end{array} $	0. 1524 0. 9883	
	$(^{1}D_{2}) \ ^{2}D_{3/2}$	$(^{3}P_{2})$ $^{4}P_{3/2}$	$(^3P_1)$ $^2P_{3/2}$
$\begin{array}{ccc} (^1D_2) & ^2D_{3/2} \\ (^3P_2) & ^4P_{3/2} \\ (^3P_1) & ^2P_{3/2} \end{array}$		0.1581 0.7302 -0.6646	-0. 0851 0. 6807 0. 7276
	$(^3P_2)$ $^4P_{1/2}$	$(^3P_0)\ ^2P_{1/2}$	$(^{1}S_{0})$ $^{2}S_{1/2}$
$ \begin{array}{ccc} (^3P_1) & ^4P_{1/2} \\ (^3P_0) & ^2P_{1/2} \\ (^1S_0) & ^2S_{1/2} \end{array} $		0. 1561 0. 9805 -0. 1191	0. 1283 0. 0994 0. 9867

The eigenvector of each level is given in the column under the adopted designation of the level. The designations on the extreme left refer to pure LS states. The elements of the eigenvectors are the expansion coefficients that occur in the corresponding eigenfunction. For example, the eigenfunction of the level called $^2\mathrm{D}_{5/2}$ is a linear combination of those for $^2\mathrm{D}_{5/2}$ and $^4\mathrm{P}_{5/2}$ and is given by

$$\psi$$
 ("2D_{5/2}") = 0.9883 ψ (2D_{5/2}) - 0.1524 ψ (4P_{5/2}).

The ²D_{5/2} level thus has a purity of 97.7 percent in this approximation.

which point to apparent violations of the selection rule on the orbital angular momentum of the external electron, $\Delta l = \pm 1$. A reasonable interpretation of these lines is that $(^1D_2)5s$ 2D contains an admixture of $(^3P_2)4d$ eigenfunctions. Probably each of the $(^1D_2)5s$ 2D levels is mixed with both of the two possible $(^3P_2)4d$ levels of the same J-value, but the 5s $^2D_{5/2}\times 4d$ $^4D_{5/2}$ and 5s $^2D_{3/2}\times 4d$ $^4D_{3/2}$ interactions are perhaps the stronger because of the proximity of the levels. This would explain both the $\Delta l = 3$ transitions and the fact that the $(^1D_2)5s$ 2D levels have apparently been pushed downward. The situation is actually more complicated than this, because the strong $(^1D_2)5s$ $^2D - (^3P_2)4f[2]^\circ$ combinations are caused partly by the additional $(^1D_2)5p$ $^2D^\circ \times (^3P_2)4f[2]^\circ$ interaction.

The (${}^{3}P$)5s levels also combine weakly with (${}^{3}P$)nf levels, partly because of small $5s \times 4d$ interactions and also through the (${}^{1}D_{2}$)5s ${}^{2}D$ levels,

The other effect of the $({}^{1}\mathrm{D}_{2})5s \times ({}^{3}\mathrm{P}_{2})4d$ interaction is a reciprocal one to the above and concerns the unexpected intensity of the $({}^{3}\mathrm{P}_{2})4d - ({}^{1}\mathrm{D}_{2})5p$ transitions, particularly the line at 10840 Å, resulting from the transition $({}^{3}\mathrm{P}_{2})4d$ ${}^{4}\mathrm{D}_{5/2} - ({}^{1}\mathrm{D}_{2})5p$ ${}^{2}\mathrm{D}_{5/2}^{s}$.

Already at 5s the separation of the level groups according to parentage is obvious. For $n \ge 6$ the dominance of ζ_{4p} should permit the calculation and identification of the ns levels by use of simplified matrices for $J_c j$ or $J_c l$ coupling. Since these are identical in the two schemes for l=0, the coupling might be called $J_c s$ coupling. The complete $J_c s$ matrices for $p^4 s$ have been given by Minnhagen [19].

Table 10. Observed and calculated g-factors for the Br I 4p45s levels

Level design.	Energy	ergy g		
		Obs.	Calc.	Landé
$\begin{array}{c} (^3P_1)\ ^4P_{1/2} \\ (^3P_0)\ ^2P_{1/2} \\ (^1S_0)\ ^2S_{1/2} \end{array}$	$K \\ 66884 \\ 68970 \\ 91938$	2. 608 0. 734	2. 601 0. 734 (1. 998)	2. 667 0. 667 (2. 000)
	g-sum =	3. 342	3. 335	3. 333
$\begin{array}{c} (^3P_2)\ ^4P_{3/2} \\ (^3P_1)\ ^2P_{3/2} \\ (^1D_2)\ ^2D_{3/2} \end{array}$	64907 67184 75909	1. 532 1. 522 0. 826	1. 533 1. 515 0. 819	1. 733 1. 333 0. 800
	g-sum=	3. 880	3. 867	3. 867
$(^{3}P_{2})$ $^{4}P_{5/2}$ $(^{1}D_{2})$ $^{2}D_{5/2}$	63436 75890	1. 595 1. 210	1. 591 1. 209	1. 600 1. 200
	g-sum=	2. 805	2. 800	2. 800

⁶ Two typographical errors appear in those matrices: In the J=3/2 matrix the off-diagonal $(-\zeta_p/\sqrt{2})$ element should connect the levels based on the $(^3P_2)-(^1D_2)$ parents rather than $(^3P_1)-(^1D_2)$, which should be zero. Further, the coefficient of G^1 in the diagonal element for the J=1/2 level based on $(^3P_1)$ should be 2/3 rather than 1/3.

As was the case in II, the relative magnitudes of the J_{cs} matrix elements permit a simplification such that the p^4s levels in BrI may be calculated by use of the following approximate expressions:

$$(^{3}P_{2}) \ ^{4}P_{5/2} = B + (^{3}P_{2})$$

$$(^{3}P_{2}) \ ^{4}P_{3/2} \qquad (^{3}P_{1}) \ ^{2}P_{3/2}$$

$$(^{3}P_{2}) \ ^{4}P_{3/2} \qquad B + (^{3}P_{2}) + \frac{5}{6} G^{1} \qquad \frac{\sqrt{5}}{6} G^{1}$$

$$(^{3}P_{1}) \ ^{2}P_{3/2} \qquad \frac{\sqrt{5}}{6} G^{1} \qquad B + (^{3}P_{1}) + \frac{1}{6} G^{1}$$

$$(^{3}P_{1}) \ ^{4}P_{1/2} \qquad (^{3}P_{0}) \ ^{2}P_{1/2}$$

$$(^{3}P_{1}) \ ^{4}P_{1/2} \qquad B + (^{3}P_{1}) + \frac{2}{3} G^{1} \qquad \frac{\sqrt{2}}{3} G^{1}$$

$$(^{3}P_{0}) \ ^{2}P_{1/2} \qquad \frac{\sqrt{2}}{3} G^{1} \qquad B + (^{3}P_{0}) + \frac{1}{3} G^{1}$$

$$(^{1}D_{2}) \ ^{2}D_{5/2,3/2} = B + (^{1}D_{2}) + \frac{1}{3} G^{1}$$

$$(^{1}S_{0}) \ ^{2}S_{1/2} \qquad = B + (^{1}S_{0}) + \frac{1}{3} G^{1},$$

where the parent symbols stand for the observed values: (${}^{3}P_{2}$)=0, (${}^{3}P_{1}$)=3136, (${}^{3}P_{0}$)=3838, (${}^{1}D_{2}$)=12089, and (${}^{1}S_{0}$)=27867 K. LS symbols are given the Br I $p^{4}s$ levels in the company of $J_{c}s$ energy expressions, but the $J_{c}l$ correspondence can be found in table 6.

The LS eigenvectors for (3P)5s show that the J=1/2 levels are relatively pure in the LS scheme while the J=3/2 are thoroughly mixed. As might be expected, the calculation of (3P)5s from the above J_cs expressions gives just the opposite results, i.e., the J=1/2 levels are thoroughly mixed in the J_cs scheme while the J=3/2 levels are pure. This same behavior was found by Minnhagen in I I. The large electrostatic interaction of the J=1/2 levels is also present for $n \ge 6$ and reveals itself through a larger

nondiagonal contribution to the energy than for the (3 P) J=3/2 levels. The comparison of observed 6s and 7s levels with those calculated from the $J_{c}s$ expressions is given in table 11. Since the agreement is excellent for a rather large range of G^{1} , and since small perturbations are present at both 6s and 7s, it makes little sense to debate the exact value of G^{1} in this approximation. A strict least squares fit gives $G_{1}=\frac{1}{3}G^{1}$ as 172 K and 62 K for 6s and 7s, respectively. In the preparation of table 11, however, we have assumed that G_{1} varies approximately as $(n^{*})^{-3}$ in accordance with theory and have taken $G_{1}(4p, 6s)=190$ K and $G_{1}(4p, 7s)=78$ K, both based on the value of $G_{1}(4p, 5s)$ found above. The non-diagonal contributions in kaysers to the (3 P)6s, 7s energies for 4 P are as follows:

$$J = 1/2$$
 ± 115 ± 19 $J = 3/2$ ± 16 ± 3

where the upper sign is always taken with the higher level of a given J. It is seen that even at 7s a positive confirmation of the level assignments can be effected by use of diagonal expressions only.

Coupling diagrams of the type found in TAS, chapter 11, cannot be constructed for a p^4s configuration because the number of parameters necessary to determine the level intervals is greater than two. However, since the splitting of Br II (3P) is small compared to the distance from (3P) to (1D_2) or (1S_0), the Br I (3P)ns levels can be described to a satisfactory approximation as a function of the ratio $G_1(4p,ns)/\zeta_p$. This is accomplished by omiting in the energy matrices the elements associated with (1D_2) and (1S_0). A coupling diagram showing the structure of $p^4(^3P)ns$ in the transition from LS to J_cj - or J_cl -coupling has been constructed by Edlén [22, p. 132]. An equivalent diagram for the observed Br I levels is presented in figure 4 and strikingly agrees with the theoretical behavior of these levels.

Table 11. Observed Br i 4p⁴6s, 7s energy levels compared with calculations using the simplified J_cs matrices

All energies and parameters are stated in kaysers.

Level design.	Observed position	Calc.	ObsCalc.	Parameters
$ \begin{array}{c} (^3P_2) 6s ^4P_{5/2} \\ (^3P_2) 6s ^4P_{3/2} \\ (^3P_1) 6s ^4P_{1/2} \\ (^3P_1) 6s ^2P_{3/2} \\ (^3P_0) 6s ^2P_{1/2} \\ (^1D_2) 6s ^2D_{3/2} \end{array} $	82236 82662 85577 85435 86279 94471	82199 82658 85601 85447 86342 94478	+37 $+4$ -24 -12 -63 -7	$B = 82199$ $G_1 = 190$
$\begin{array}{c} (^{3}P_{2})7s\ ^{4}P_{5/2} \\ (^{3}P_{2})7s\ ^{4}P_{3/2} \\ (^{3}P_{1})7s\ ^{4}P_{1/2} \\ (^{3}P_{1})7s\ ^{2}P_{3/2} \\ (^{3}P_{0})7s\ ^{2}P_{1/2} \end{array}$	88135 88275 91348 91284 92090	88100 88292 91373 91278 92035	+35 -17 -25 $+6$ $+55$	B=88100 G ₁ =78

Since G_1 varies approximately as $(n^*)^{-3}$ and ζ_p remains essentially constant along the series, we have plotted in figure 4 the ns levels, referred to the center of gravity of each group, against $(n^*)^{-3}$. The mean values $\frac{1}{2}(^4P_{\frac{1}{2}}+^2P_{\frac{1}{2}})$ and $\frac{1}{2}(^4P_{\frac{3}{2}}+^2P_{\frac{3}{2}})$ are also plotted and found to give two approximately straight parallel lines, as expected. The slight irregularities are caused partly by small perturbations of the levels. The observed crossover of the $^4P_{\frac{1}{2}}$ and $^2P_{\frac{3}{2}}$ levels is in accordance with theory. It should be noticed that the plot yields an unambiguous coordination of the levels to specific levels of the (^3P) limit.

From a comparative study of the plots in figures 4 and 5 one can obtain a qualitative picture of the perturbations present in the $(^{3}P)ns$ series, mostly caused by $ns \times nd$ interactions. The displacements of the perturbed levels are all small, being of the order

of 20 to 50 K.

The $(^{3}P_{0})$ $6s^{2}P_{1/2}$ level seems to be about 75 K too low, but the cause is not clear, unless this displacement results from the influence of the $J=\frac{1}{2}$ levels in the $(^{3}P_{2})5d$ configuration that lie 1000 to 1500 K higher. The $(^{1}D_{2})4d^{2}P_{1/2}$, $^{2}S_{1/2}$ levels are probably responsible for the perturbation of $(^{3}P_{1})8s^{4}P_{1/2}$, which is about 50 K too high. The higher $(^{3}P_{2})$ ns series members all suffer slight perturbations owing to $ns \times nd$ interactions. While the plots in figure 5 seem to be highly irregular at the end corresponding to high n, it should be remembered that T is there relatively small, and n^{*} is therefore quite sensitive to changes in T, as is evident from the expression

$$\Delta n^* \sim -\frac{R^{1/2}}{2T^{3/2}} \Delta T.$$

(b) $4p^4nd$: Sorting out and interpreting the Br I nd energy levels is perhaps the most intricate problem in the analysis of this spectrum. Abundant perturbations and an intermediate type of coupling severely complicate the analysis. We have in the analysis attempted to make the best possible use of the observations. The interpretation of the Br I nd levels as presented in table 6 seems to be the most reasonable derivable from general arguments, observed line intensities, application of the Ritz formula, approximate calculations, and the like. A more detailed study of the nd configurations based on accurate calculations including interconfigurational interaction might well necessitate the revision of some of the assignments, particularly of those levels having low J-values. Reconsideration of the nd $J=\frac{1}{2}$ levels might even bring the interpretation of $\binom{1}{3}$ 5s $\binom{2}{3}$ into question, as pointed out above.

The separation of the nd levels of each configuration into groups according to the parent limits is quite apparent. The $(^3P_{1,0})4d$, 5d, 6d levels all lie below the $(^3P_2)$ limit and have been found. The $(^3P_1)7d$ levels are predicted to fall below the $(^3P_2)$ limit but are so close (~ 94900 K) to this limit that they have not been discovered. All $(^3P_2)nd$ series are complete through 8d; each series thus contains

five members. The $(^{3}P_{2})nd^{4}F_{9/2}$ series is known through 12d. Of the nd levels based on the high $(^{1}D_{2})$ and $(^{1}S_{0})$ parents, only the $(^{1}D_{2})4d$ group falls below the $(^{3}P_{2})$ limit. All levels in this group have

been identified at least tentatively.

The grouping of the nd levels according to parent suggests that the coupling of these configurations is intermediate between the J_cl and J_cj schemes. The diagonal energy expressions for p^4d in these two schemes have been calculated by Möller and Källén, respectively, and can be found in the paper by Minnhagen [19]. A comparison of the J_cl expressions with the observed levels reveals that even the observed ordering of K-pairs with respect to energy is not reproduced by the diagonal J_cl expressions on the assumption that the dominant contribution to the energy is due to $F^2(4p, nd)$.

It is felt that the reason for this lies, at least in the case of 4d, in the magnitude of the G^1 integral, whose effect on the energies is probably greater than that of F^2 , implying a large departure from $J_c l$ coupling, which requires a small G^k contribution. The large G^1 contribution probably accounts in part for the high position of the $(^3P_2)4d[2]_{5/2,3/2}$ levels. It may be that the levels $(^3P_2)4d[3]_{5/2}$ and $[2]_{5/2}$ should have their designations interchanged. The interaction of the 4d levels having J=5/2, 3/2 with the $(^1D_2)5s$ $^2D_{5/2,3/2}$ levels contributes to the high position of the 4d pair just mentioned, but this effect is probably less significant than the size of the G^1 integral.

The large departures from any pure coupling scheme in the nd configurations, as well as sizeable perturbations, render the assignment of designations to the levels a rather difficult problem. Both the LS and $J_c l$ symbols assigned these levels in table 6 are based on the intensities of the 4d-nf, 5p-4d, and 5p-5d combinations. There still remain several inconsistencies; in particular the combining properties of some levels vary irregularly along a series. Only a very detailed consideration of perturbations, together with accurate calculations, can aid in the solution of this problem. Even a cursory examination of the effective quantum numbers of the nd levels reveals marked irregularities along series. Apart from small interaction with some ns levels, the perturbations are primarily a result of the overlapping or proximity of groups of levels with different n and based on different parents. The most striking case of this is the $({}^{3}P_{1})5d\times({}^{3}P_{2})6d$ interaction, which causes the $({}^{3}P_{1})5d$ group of levels to be slightly higher than expected. It is very possible that the $(^{3}P_{2})6d$ $^{4}F_{7/2}$ and $(^{3}P_{1})5d$ $^{2}F_{7/2}$ levels should have their designations interchanged. A $(^{3}P_{1})6d \times (^{3}P_{2})9d$ interaction has apparently pushed the (³P₁)6d group down. The separation of these latter groups is difficult, however, and the assignments should be regarded as tentative. Their further interaction with the slightly higher (1D2)4d levels is also probable.

The $(^{1}D_{2})4d$ $^{2}G_{9/2,7/2}$ levels have been established with certainty (see table 6). The other seven levels of the $(^{1}D_{2})4d$ group have received their designations primarily on the basis of their position being such that they do not fit well into any (^{3}P) groups.

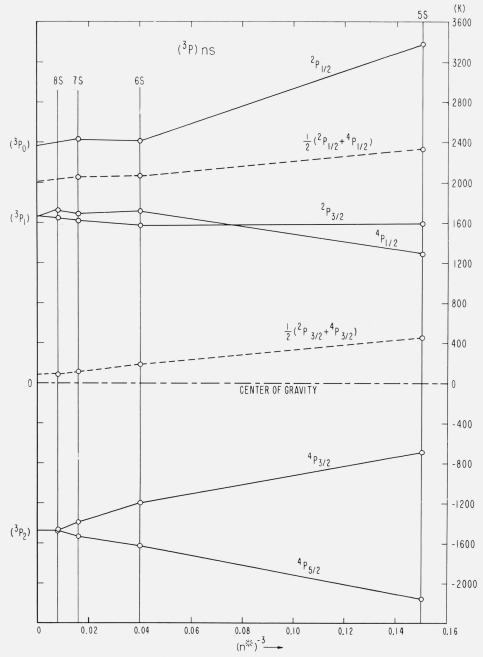


Figure 4. Observed levels of $4p^4(^3P)ns$, referred to center of gravity, and plotted against $(n^*)^{-3}$ to show the convergence toward the Br II $4p^4$ $^3P_{2,1,0}$ limits.

(Furthermore, we had just seven levels left over after unraveling the $(^{3}P)nd$ levels.) The order of the $(^{1}D_{2})4d$ levels bears a striking resemblance to the corresponding group in Kr II, but is in sharp disagreement with the Jd diagonal expressions (assuming dominant F^{2}), which predict the $J=\frac{1}{2}$ levels to be the lowest.

The most important nd series showing any degree of regularity is the $(^{3}P_{2})nd$ $^{4}F_{9/2}$ series, for which $(2-\delta)$ is plotted against T in figure 6. The nonlinearity of this plot is interpreted not as a failure to have the correct ionization potential, but as the

effect of a gradual change in coupling along the series. The members from 9d through 12d are almost exactly linear. A change in the assumed ionization potential would cause the line connecting these levels to curve up or down. A least squares calculation of the Ritz constants α and β from these levels gives $\alpha = 1.35276$, $\beta = -3.442 \times 10^{-6}$, from which the entire series has been calculated. The results, given in table 12, reveal not only by how much the early series members lie too high because of the different coupling conditions, but also by how

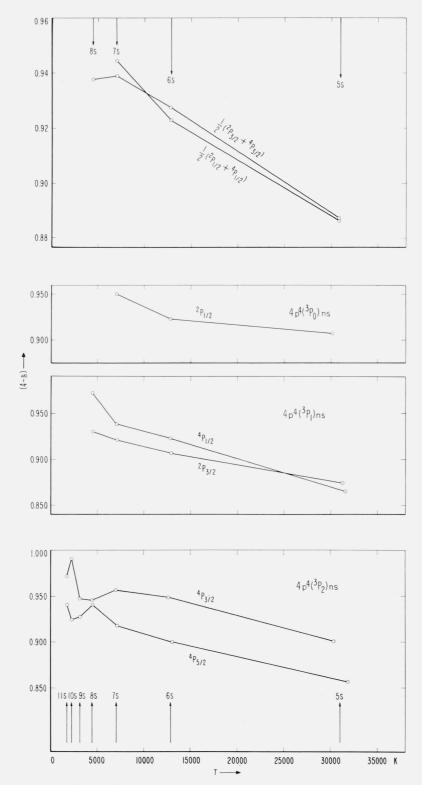


Figure 5. Diagram showing the observed Brins series.

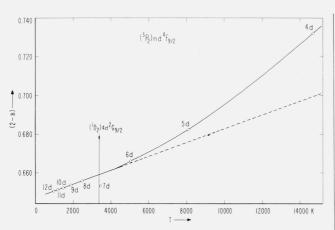


Figure 6. Diagram showing the very regular (3P2)nd 4F9/2 series.

The (^1D_2)4d ^2G _{9/2} level was found from its influence on the 7d member of this series.

much the 7d and 8d levels have been pushed down and up, respectively, by the $(^{1}D_{2})4d\ ^{2}G_{9/2}$ level between them.

Table 12. Observed energy levels of the $4 \mbox{$^{\circ}$}(^3P_2) nd \,^4F_9/_2 \, series compared with calculations using the Ritz formula, as described in the text.$

All energies are stated in kaysers. $\alpha = 1.35276$ $\beta = -3.442 \times 10^{-6}$

Config.	Observed	Calculated	Obscale
4d	80587. 61	80221. 57	+366. 04
5d	87191. 97	87160. 44	+31.53
6d	90244. 89	90241. 40	+3.49
7d	91851. 14	91858. 17	-7.03
	$4p^4(^1{ m D}_2)4d^{\;2}$	$G_{9/2} = 91921.86$	
8d	92807. 81	92807. 63	+0.18
9d	93411. 48	93411. 49	-0.01
10d	93818. 96	93818. 95	+0.01
	94106, 70	94106.70	0. 00
11d	0 1100. 10		

These slight deviations of the points representing $(^3P_2)7d$, $8d\ ^4F_{9/2}$ indicate with certainty the presence of the nearby perturbing level, which can be only $(^1D_2)4d\ ^2G_{9/2}$. Until this perturbation was recognized, the only strong Br I lines still remaining unclassified were those at 6096 and 6133 A. That the line at 6096 Å results from the transition $(^3P_2)5p\ ^4D_{7/2}^2-(^1D_2)4d\ ^2G_{9/2}$ is unquestionably correct, because (i) it fortunately happens to be one of the two unclassified lines whose hfs were reported by Tolansky and Trivedi [7], its observed structure fitting fairly well the splitting of $(^3P_2)5p\ ^4D_{7/2}^2$ derived from other lines, and (ii) it yields a value 91922 K for $(^1D_2)4d\ ^2G_{9/2}$, which is the proper position to account for the perturbation of the $nd\ ^4F_{9/2}$ series.

The other line, at 6133 A, is interpreted as the transition $(^{3}P_{2})5p$ $^{4}D_{5/2}^{\circ}-(^{1}D_{2})4d$ $^{2}G_{7/2}$. One additional line has been observed in the radiometric region for each of the ^{2}G levels. These are combinations with $(^{1}D_{2})5p$ $^{2}F_{7/2,5/2}^{\circ}$ and serve as further evidence for the reality of the ^{2}G levels.

(c) $4s4p^6$: The $4s4p^6$ configuration yields only one level, $^2S_{54}$. Bromine is the only halogen for which this level has been found with certainty. A calculation based upon the irregular-doublet law predicts the $^2S_{54}$ term to lie near 87000 K. Of the nine even levels with J=1/2 that were found in the range 80000-90000 K, only eight can arise from configurations other than $4s4p^6$. These can all be unambiguously identified on the basis of position and intensity considerations. The remaining level, at 84825.60 K, must be $4s4p^6$ $^2S_{54}$. This identification explains the high intensity of the ultraviolet lines at 1232 Å, resulting from the transition $4s^24p^5$ $^2P_{54}^9-84825$, and at 1179 Å, $4s^24p^5$ $^2P_{3/2}^9-84825$: the ground $^2P^\circ$ term is the only known term to which the $4s4p^6$ $^2S_{54}$ can make transitions in a one-electron jump.

In addition to the ultraviolet lines and a faint, wide line at 12369 A, the only other lines from ${}^2S_{>2}$ that have been observed are two moderately strong doublets at 10184 Å and 11094 Å. The doublet character of each of these lines is assumed to result from the wide hfs splitting of ${}^2S_{>2}$, which is expected, since the unpaired 4s electron in the $4s4p^6$ configuration should couple strongly with the nucleus. The classifications of these lines are as follows:

10184 Å
$$5p \, {}^{4}P_{3/2}^{*} - sp^{6} \, {}^{2}S_{1/2}$$

11094 Å $5p \, {}^{4}P_{1/2}^{*} - sp^{6} \, {}^{2}S_{1/2}$

The components of 11094 Å appear single with our resolution but there is slight shading in the two components of 10184 Å. This results from the hfs splitting of $5p \, ^4P_{3/2}^{\circ}$, which according to Tolansky and Trivedi has a structure (0.127, 0.086, 0.043) in K. To get an estimate of the splitting of the ²S₁₂ level, we note that the separations of the components of 10184 Å and 11094 Å are 0.45 and 0.59 K, respectively. The large difference between these two values might at first seem disturbing. But a simple sketch of the appropriate transitions, taking into account the known hyperfine structure of ${}^{4}P_{3/2}^{\circ}$ (four F-levels) and adopting any reasonable value for the unknown structure of ${}^4P^{\circ}_{\frac{1}{2}}$ (two F-levels), reveals that one would expect the components of 10184 Å to be about 0.1 K closer together than the components of 11094 A. Furthermore, it is apparent from such a sketch that the measured separation of the centers of gravity of the 11094 Å components very nearly represents the true hfs splitting of the ²S₁₆. We therefore adopt 0.59 K as the splitting of the ${}^{2}S_{\frac{1}{2}}$ level.

A calculation of the theoretical structure of this level is revealing. Following Slater [23], we can

write the expectation value of the nuclear interaction contribution to the total Hamiltonian as

$$\langle H' \rangle = \frac{\alpha^2}{3} g_N R^2(0) K$$
 rydbergs.

(This formula holds only for configurations containing a single unpaired s-electron.) ⁷ Here α is the fine structure constant (1/137), g_N is the weighted average bromine nuclear g-factor (= 7.93×10^{-4}), R(0) is the value of the normalized s-eigenfunction at the origin, and

$$K = F(F+1) - I(I+1) - J(J+1)$$
.

Since J=1/2 and the nuclear spin I=3/2, the energy separation of the two F-levels will be

$$\Delta E = \frac{4\alpha^2}{3} g_N R^2(0)$$
 rydbergs.

We have calculated R(0) from recently published [24] Hartree-Fock radial wave functions for the bromine $4s^24p^5$ configuration. The 4s-orbital of this configuration will not be very different from that of the $4s4p^6$ configuration. We find R(0) = 15.6. Inserting numerical values in the above equation, we finally obtain

$$E=1.37\times10^{-5}$$
 rydbergs
=1.50 kaysers.

(A relativity correction would increase this value by about 14%.) The above result is to be compared with the observed value 0.59 K. Some of the discrepancy between these values must arise from configuration interaction. Such an interaction is undoubtedly present, because the two lines under discussion are examples of so-called two-electron jumps. The occurrence of these lines at all indicates that the sp^6 configuration is mixed with a configuration that would make the transition in a one-electron jump. The interaction involved here is probably with one of the J=1/2 levels of the 4d configuration, possibly 4d $^4P_{1/2}$. One would expect the level with which the sp^6 $^2S_{1/2}$ is mixed to share some of the latter's hyperfine structure. Each of the 4d levels with J=1/2 was checked for hfs, but none could definitely be established from our observations.

6. Levels of Odd Parity

(a) $4p^4np$: The analysis of the system of odd levels of Br I was enormously simplified by the occurrence of almost pure $J_c l$ -coupling in the nfconfigurations. Upon the identification of a few of the nf levels, the rest could easily be calculated and identified. We could then safely regard all odd levels left over as belonging to excited npconfigurations.

With the exception of the $({}^{3}P_{1})8p$ group, which is predicted to center around 94675 K, all np levels for n=5, 6, 7, 8, 9 that fall below the ($^{3}P_{2}$) limit have been found. Two levels belonging to ($^{3}P_{2}$)10p have been identified also. Of the levels based on the $({}^{1}D_{2}, {}^{1}S_{0})$ parents, only the $({}^{1}D_{2})5p$ levels are below the (³P₂) limit. As in the ns and nd configurations, the np levels fall into groups that reflect the structure of the Br II p^4 parents. The coupling of the higher np configurations is between the $J_c l$ and $J_c j$ schemes, but apparently closer to $J_c l$. Happily, configuration interaction of the np set of levels with the nf set is extremely small. The very different coupling exhibited by these two sets of odd levels may partly account for this minimal interaction.

The matrices of spin-orbit interaction for a p^4p configuration have been calculated in an LS basis and are presented in the following paper in this issue of the Journal of Research. A study of the Br I np configurations by means of these matrices has not yet been completed. The results of the calculations will appear later, together with the

results of Zeeman observations in progress.

It may be mentioned that the electrostatic [18] and spin-orbit [19] elements of the p^4l matrices for pair-coupling have just been published, as well as the p^4l matrices for J_{cj} -coupling [19], calculated by Källén. The assignment of K-values to the Bri np levels is based on intensities and approximate calculations with the J_{cl} diagonal energy expressions given in reference [19]. The LS symbols are based on the intensities of 5p—5s, 6s, 4dtransitions.

The Bri 5p configuration is rather isolated and except for a few small interactions is regarded as a pure configuration. It should therefore be possible to represent quite accurately in parametric form the distribution of the 5p levels. It was for this reason, in fact, that the p^4p matrices in intermediate coupling were calculated. Since the (1S_0) 5p $^2P^\circ$ term is above the (3P2) limit and has not been observed, it is not possible to use the diagonal sum rule in the calculation of the parameters. The parameters may, however, be estimated in a variety of ways based on diagonal energy expressions in the several coupling schemes for levels of high J-value, on approximately calculated positions for (1S_0) 5p 2P , and on level differences. The approximate parameters found in this way can be inserted in the energy matrices and improved by iteration.

The most interesting interaction occurring with the 5p configuration is perhaps the $(^1D_2)$ 5p $^2D_{5/2,3/2}$ $\times (^3P_2)4f[2]_{5/2,3/2}^3$ interaction. The $(^3P_1)6p^2D^\circ$ term lies somewhat higher but may also be involved. Accurate calculations given below show that the two 4f levels have been pushed down by 13 K and 21 K for J=5/2 and J=3/2, respectively. This partly accounts for the surprisingly strong intensity of $(^1D_2)5s$ $^2D-(^3P_2)4f[2]^\circ$ transitions, as pointed out above.

The $({}^{3}P_{2})5p {}^{4}D_{3/2}^{\circ} = [1]_{3/2}^{\circ}$ level seems to lie abnormally high, judging from the diagonal J_{cl} energy

⁷ It should also be mentioned that a factor of 2/3 is missing from this equation in Slater's work, Vol. II, p. 266.

expressions. This same behavior was noticed in the case of I I by Minnhagen [19], who calculated the I I 6p levels by use of the $F^2(5p, 6p)$ parameter only. It was suggested that configuration interaction might be responsible for the high position of the level in question. As evidence for such an interaction, Minnhagen cited the fact that the intensity of the I I line classified as $(^3P_2)6s[2]_{5/2}-(^3P_2)6p[1]_{3/2}^{\circ}(=8044\text{ Å})$ is much stronger than that of the line $(^3P_2)6s[2]_{5/2} (^{3}P_{2})6p[3]_{7/2}^{\circ}$ (=9058 Å) whereas the latter line should be stronger for all unperturbed couplings. The only J=3/2 level below I i (3P_2) 6p is the ground level $^5p^5$ ²P_{3/2}, which renders perturbation difficult to accept. It seems reasonable that the intensity ratio of I i lines just mentioned is, rather, a result of the fact that 8044 Å falls at the wavelength of peak sensitivity of EK—"N" photographic plates, while 9058 Å was undoubtedly photographed on "Q" or "Z" plates, which are less sensitive than "N" plates by about a factor of 10. In Br 1, the corresponding lines both occur in the "N" region, and the relative intensities are in the expected order. It is felt that the high position of the level in question, both in I 1 and Br 1, is to be explained by the magnitude of the G^0 (np, n'p) integral, which contributes sizeably to the Br 1 5p and I i 6p energies and should not be neglected. The $J_c l \ p^4 p$ diagonal energy expressions for $(^3P_2)np$ [2]°, [1]° are, except for an additive constant:

$$\begin{split} &[2]_{5/2}^{\circ} - \frac{7}{50} \; \mathrm{F}^2 \qquad \qquad + \frac{18}{600} \; \mathrm{G}^2 + \frac{1}{6} \; \zeta_{n'p} \\ &[2]_{3/2}^{\circ} - \frac{7}{50} \; \mathrm{F}^2 + \frac{45}{24} \; \mathrm{G}^0 + \frac{153}{600} \; \mathrm{G}^2 - \frac{1}{4} \; \zeta_{n'p} \\ &[1]_{3/2}^{\circ} + \frac{7}{50} \; \mathrm{F}^2 + \frac{5}{24} \; \mathrm{G}^0 + \frac{17}{600} \; \mathrm{G}^2 - \frac{1}{4} \; \zeta_{n'p} \\ &[1]_{1/2}^{\circ} + \frac{7}{50} \; \mathrm{F}^2 + \frac{20}{24} \; \mathrm{G}^0 + \frac{50}{600} \; \mathrm{G}^2 + \frac{1}{2} \; \zeta_{n'p} \end{split}$$

where all integrals refer to the (np, n'p) interaction. Approximate Br I 5p parameters, derived as described above, indicate that the effect of G^0 on $[2]_{3/2}^s$ brings its diagonal energy very close to that of $[1]_{3/2}^s$. Since the two levels are connected by G^0 , G^2 , and $\zeta_{n'p}$, they repel each other in such a way, neglecting the action of the other levels in the configuration, as to make $[1]_{3/2}^s$ appear high and the K=2 levels appear as a closer pair. The $[1]_{1/2}^s$ level is pushed down by all 5p J=1/2 levels above it. This same type of effect is probably operating also in the case of the high Br I $(^3P_2)4d[2]$ pair, as mentioned previously.

A qualitative picture of the run of Br 1 np levels and perturbations can be obtained from figure 7, in which the quantum defect is plotted against absolute level value. The points representing the $(^{3}P_{2})np$ $^{4}D_{7/2}$ series are very nearly linear for n=6, 7, 9, 10. The slight depression at n=8 is due to the interaction with $(^{3}P_{1})4f[3]_{7/2}$. An application of the Ritz formula with $\alpha=2.57410$, $\beta=+3.190\times10^{-6}$ derived by least squares from $(^{3}P_{2})$ 7, 9, 10p $^{4}D_{7/2}^{*}$,

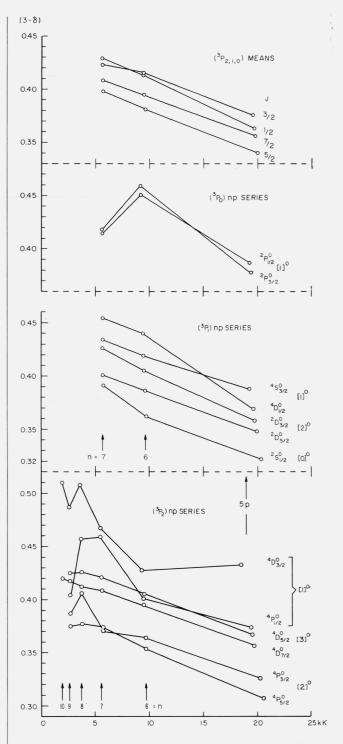


Figure 7. The $(^{3}P)np$ series: $(3-\delta)$ plotted against T.

reproduces these levels almost exactly and shows that the level at 8p is depressed by about 2.6 K, its "unperturbed" position being 91540.9 K. The same Ritz formula predicts (3P_2)5p $^4D_{7/2}^{\circ}$ to fall 114 K higher than observed, the departure being a result of the regular coupling change along the series.

The great deviation from linearity exhibited in figure 7 by the early members of the $(^3P_2)np$ $^4D_{3/2}^\circ$, $^4P_{1/2}^\circ$ and $(^3P_0)np$ $^2P_{3/2,1/2}^\circ$ series is probably not entirely due to perturbations, since the plots of mean term energies for each J-value are much more regular. However, the possibility of a configuration interaction affecting these levels in such a way as approximately to cancel out in the means should not be overlooked. It seems fairly certain that there is a $(^3P_0)6p\times(^3P_2)7p$ interaction, and the $(^3P_2)6p$ $^4D_{3/2}^\circ$, $^4P_{1/2}^\circ$ levels may be influenced by $(^1D_2)5p$ $^2P_{3/2,1/2}^\circ$. The $(^3P_2)7p$, 8p, $^4P_{3/2}^\circ$ levels seem to have been pushed down and up, respectively, through interaction with $(^3P_0)6p$ $^2P_{3/2}^\circ$, which lies between them.

The most obvious wholesale mutual interactions in the np configurations affect the close (${}^{3}P_{1}$)7p and (${}^{3}P_{2}$)9p groups, the levels of the first group being higher than expected and those of the second group

(except $9p \, ^4\mathrm{D}_{7/2}^{\circ}$) lower.

(b) $4p^4nf$: All the 67 levels ascribed to nf configurations in table 5 are new. The pronounced J_cl coupling prevailing in these configurations permits the levels to be calculated very exactly by the theoretical expressions of electrostatic energy given by Möller [18]. The J_cl designations of the levels as presented in table 5 are therefore, with one or two exceptions discussed below, definite. The pair structure of these levels is very marked, as can be seen by examining the pair splittings in table 5. All nf levels based on limits other than $(^3P_2)$ and predicted to lie below the $(^3P_2)$ limit have been found. These are the 16 levels of the $(^3P_{1,0})4f$, 5f groups. All $(^1D_2, ^1S_0)nf$ levels lie above the $(^3P_2)$ limit.

Most of the observed transitions involving the

Br I nf levels occur in the $({}^{3}P_{2})4d - ({}^{3}P)nf$ arrays. As mentioned above, 5s-nf transitions are also observed, primarily as a result of a $5s\times4d$ interaction, but nearly all other lines involving nf levels would be difficult to observe since they fall in a very unfavorable region of the spectrum, the far infrared and radiometric regions. The $({}^{3}P_{2})4d-({}^{3}P_{2})nf$ arrays are similar for each value of n as regards the intensity relationships of the various lines in the array, but the intensity of any particular transition decreases as a rule with increasing n. In general, for transitions between two pair-coupled configurations the lines will be stronger when l, K, and Jchange by unity in the same direction, or, with a given $\Delta l (= \pm 1)$, when $\Delta J = \Delta K (= \pm 1, 0)$. As usual, also, the stronger lines involve the higher J-values. The representative $({}^{3}P_{2})4d - ({}^{3}P_{2})5f$ array is shown in table 13.

In this table, the adopted K-values of the $(^3P_2)4d$ levels are found together with the corresponding LS designations. The only rigorous selection rule here seems to be that on J. It will be noticed, however, that three transitions from 5f in this table that would be allowed by the selection rule on J, but not by the selection rule on K in pure J_cl coupling have not been observed. The same is true in the case of these three transitions for other nf, except that $(^3P_2)4d[3]_{7/2}-(^3P_2)4f[5]_{3/2}$ occurs weakly, owing to the "mixing" of the two $(^3P_2)4d$ levels having J=7/2. On the other hand, for the 5f case in table 13 as well as for other nf, there appear strongly two lines that do "violate" the $\Delta K=0$, ± 1 rule. These are the $(^3P_2)4d$ $^4P_{3/2}-(^3P_2)$ nf $[3]_{5/2}^2$ and $(^3P_2)4d$ $^4F_{5/2}-(^3P_2)nf[4]_{7/2}^2$ transitions. For no value of n have the two $(^3P_2)4d$ $^4F_{5/2,3/2}-(^3P_2)nf[2]_{3/2}^2$ lines been ob-

Table 13. Wavenumbers and intensities of the observed (3P2)4d-(3P2)5f transitions in Br 1

$4d \rightarrow 5f \downarrow$	[4] ⁴ F _{9/2} 80587. 61	$[4] \\ {}^{4}F_{7/2} \\ 81081. \ 10$	[3] ⁴ D _{7/2} 79044. 50	[3] ⁴ D _{5/2} 79260. 91	[2] ⁴ F _{5/2} 81672. 28	$[2] \\ {}^{4}F_{3/2} \\ 81842. \ 02$	[1] ⁴ D _{3/2} 79630. 67	[1] ${}^{4}D_{1/2}$ 80026. 40	[0] ⁴ P _{1/2} 81429. 97
[5]° _{11/2} 90873.70 [5]° _{9/2} 90873.78	(600) 10286. 09	(300) 9792. 68				,			
$ \begin{array}{c} [4]^{\circ}_{9/2} \\ 90845.55 \\ [4]^{\circ}_{7/2} \\ 90846.78 \end{array} $	(100) 10257. 95	(75) 9765. 68	(500) 11801. 03 (100) 11802. 27	(200) 11585. 86	(200) 9174, 50				
[3] _{7/2, 5/2} 90851.30	(15) 10263. 69	(15) 9770. 21	(300) 11806. 78	(100) 11590. 39	(250) 9179. 02	(20) 9009, 31	(300) 11220, 60		
$\begin{array}{c} [2]_{5/2}^{\circ} \\ 90869.64 \\ [2]_{3/2}^{\circ} \\ 90868.76 \end{array}$			(75) 11825. 12	(250) 11608. 72 (75) 11607. 87	(100) 9197. 37	(40) 9027. 62	(20) 11238, 92 (150) 11238, 09	(200) 10842. 35	
[1] _{3/2} 90891.56 [1] _{1/2} 90890.88	,			(50) 11630, 57	(2w) 9219. 28	(10) 9049. 58	(90) 11260. 85 (90) 11260. 20	(80w) $10865. 15$ (80) $10864. 48$	(250) 9461.59 (100) 9460.95

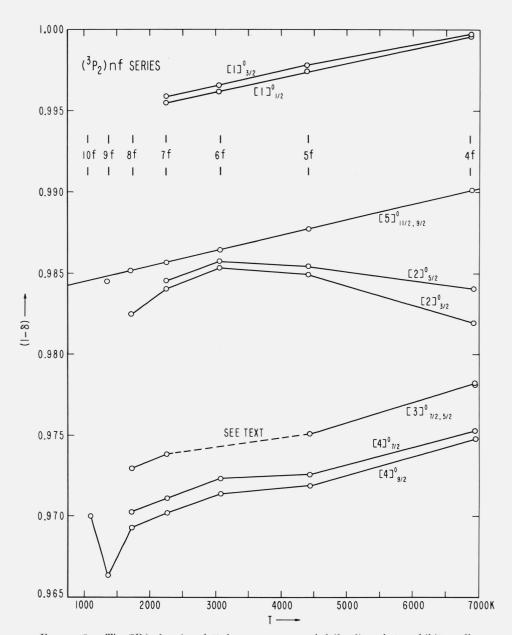


Figure 8. The $(^3P_2)$ nf series plotted on a very expanded $(1-\delta)$ scale to exhibit small perturbations.

served. The perturbations that affect the $(^{3}P_{2})4d$ levels having J=5/2, 3/2 and the difficulty at present of appraising the quantum significance of the K-values assigned these levels, as mentioned in another section, render premature any detailed correlation of intensity characteristics in the arrays under discussion.

The several perturbations that afflict the (${}^{3}P_{2}$)nf series are vividly revealed by a plot of quantum defects against relative term values for these series (fig. 8). In examining this figure, one should bear in mind that the ordinate scale is greatly expanded, and the irregularities in the plots are caused by level perturbations of only a few kaysers and less. An

idea of the vertical scale in kaysers can be gained for each value of n by noting the separation of the various pairs. For example, at n=7, the vertical separation between $7f[4]_{3/2}^{\circ}$ and $7f[4]_{7/2}^{\circ}$ corresponds to only 0.57 K, an amount that is probably greater than that by which either level is perturbed. The perturbations in the Br I nf configurations are most conveniently discussed in terms of the J_cl non-diagonal matrix components. We shall therefore consider first the energy matrices and the theoretical calculation of the nf energy level distribution.

The matrices of electrostatic energy are diagonal in J. The diagonal elements of the matrices for p^4f have the form E=E (parent level)— $2F^0$ —

 $2\bar{f}_2'F^2+2\bar{g}_2'G^2+2\bar{g}_4'G^4$, where \bar{f}_2' , \bar{g}_2' , \bar{g}_3' are numerical coefficients whose values can be found in Möller's paper, and F^2 , G^2 , G^4 are the familiar integrals defined in TAS. Levels of a given K-pair have a common \bar{f}_2' coefficient; the pair splitting with respect to J arises from different \bar{g}_2' and \bar{g}_4' coefficients. Small pair splittings, therefore, imply small G^2 and G^4 integrals. In the Brinf configurations, the contribution to the level energies by the F^2 integral is much greater than that by the G^2 and G^4 integrals. Nondiagonal F^2 is found only between levels with equal K-value and based on the different parents, ${}^1S_0-{}^1D_2$, ${}^3P_0-{}^3P_2$, and ${}^3P_1-{}^3P_2$. Also appearing off the diagonal are G^2 and G^4 , connecting levels whose K-values are equal or are different by ± 1 . In order that K be a good quantum number, it is thus essential that G^2 and G^4 be small compared with F^2 .

In the actual calculation of the nf level distribution, it is convenient and suitable to make certain approximations. All $({}^{1}S_{0}, {}^{1}D_{2})nf$ levels fall above the principal ionization limit and none was observed. But since these levels are distant and are connected to the levels based on (3P) only through nondiagonal G^2 and G^4 (and ζ_{4p}), the effect of omitting these levels from the energy matrices will cause negligible error in the calculated (3P)nf energies. Furthermore, greater accuracy in the calculated distribution of these $(^{3}P)nf$ levels (18 levels for each n) can be expected if we employ the observed positions of the Br II (3P) levels rather than adopting the parametric form for these parents. We thus replace the term E (parent level) $-2{\rm F}^{\rm 0}$ in the above diagonal expression by the quantities: B, B+3136.4, and B+3837.5, for levels based on $(^{3}P_{2})$, $(^{3}P_{1})$, and $(^{3}P_{0})$, respectively, where B is a constant for any given n. This procedure automatically accounts for nondiagonal ζ_{4p} , which may then be disregarded.

We thus have 18 levels for each n to be represented in parametric form by four parameters, B, F², G², and G⁴. The determination of meaningful parameters cannot be accomplished by a simple application of the diagonal sum rule, because every observed Br I nf level, with the exception of members of the $(^3P_2)nf[5]^{\circ}$ series, suffers interaction with other configurations. The extent of this interaction can be reasonably well estimated only in the case of the $({}^{3}P_{2})nf[4]_{9/2}^{\circ}$ levels, because almost the entire interconfigurational contribution to these must come from $(^{3}P_{1})nf[4]_{9/2}^{\circ}$ levels, there being no possibility of a J=9/2 in the np configurations. After G^4 has been estimated as described below, F^2 can be found from the $(^3P_2)nf$ $[5]_{11/2}^\circ-[4]_{5/2}^\circ=\frac{1}{6}F^2-\frac{1}{6}G^4$ diagonal interval, provided the nondiagonal intraconfigurational contribution to [4]_{9/2}, as well as the interconfigurational contribution, have first been removed from the energies as observed. The first of these contributions can be determined by performing a preliminary diagonalization of the J=9/2 matrix with a good estimate of F^2 . The second contribution can be obtained if a reasonable "unperturbed"

 $(^{3}P_{2})$ nf [4] $_{9/2}^{\circ}$ series can be deduced. If an interaction parameter (see ref. [22], p. 141) γ =23 K is assumed between the $(^{3}P_{2})$ nf [4] $_{9/2}^{\circ}$ and $(^{3}P_{1})$ nf [4] $_{9/2}^{\circ}$ series, the former series becomes satisfactorily Ritzian, and after an additional "smoothing" of the series, unperturbed level energies can be estimated to sufficient accuracy, and then F² can be derived.

Since nondiagonal contribution to the (³P₂)nf[5]° pair is negligible, the G⁴ parameter can be obtained directly from the observed separation of this pair, whose diagonal energy expressions, including the spin-orbit interaction energy of the f-electron, are

$$[5]_{11/2}^{\circ}: B + \frac{1}{15} F^2 + \frac{3}{2} \zeta_f$$

$$[5]^{\circ}_{9/2}:B+\frac{1}{15}F^2+\frac{11}{54}G^4-\frac{9}{5}\zeta_f.$$

The fact that for $n \ge 5$, the $[5]_{9/2}^{\circ}$ falls higher than $[5]_{11/2}^{\circ}$ confirms the dominance of $\frac{11}{54}G^4$ over $\frac{33}{10}\zeta_f$ and partially justifies the omission of ζ_f in our calculations. The G^4 parameter for 4f has been extrapolated from those for higher n, since the observed coincidence of the 4f pair, suggesting an equality of $\frac{11}{54}G^4$ and $\frac{33}{10}\zeta_f$, does not permit a direct calculation of G^4 .

If the $({}^{3}P_{2})nf[1]^{\circ}$ levels were unperturbed, G^{2} could be calculated fairly accurately from the pair splitting, $\frac{9}{50}G^{2}$. Even though this series as plotted in figure 8 appears Ritzian, there is some evidence that the series members have all been perturbed upwards in such a way that linearity of the plot was preserved. Nevertheless, G^{2} was crudely taken as given by the pair splittings for $n \ge 5$ and extrapolated for n = 4, since any other procedure meets with equivalent uncertainty. The remaining parameter, B, was adjusted so as to predict the $({}^{3}P_{2})$ $nf[5]_{1}^{\alpha}$ levels exactly.

 $nf[5]_{11/2}^{\circ}$ levels exactly. The levels as calculated by use of the parameters found above are compared with the observed positions in table 14 for 4f, 5f and in table 15 for 6f, 7f. The B and F^2 parameters are given in table 16, along with the almost constant products $F^2 \times n^3$. The agreement between observed and calculated energies is seen to be very good, and it is probably significant that the departures in almost every case agree with the direction of perturbation one might expect on the basis of the plots in figure 8. The calculations predict the observed ordering of K-pairs with respect to energy for all except the perturbed $(^3P_2)4f[3]^{\circ}$ and $(^3P_1)4f[3]^{\circ}$ pairs. The perturbations of the Br $_1$ nf levels arise from

The perturbations of the Br i nf levels arise from $nf \times np$ as well as $nf \times n'f$ interactions. Considering first the latter type, we remember that since the coefficients of the Slater integrals in the nf energy matrices are based entirely upon the angular parts of the eigenfunctions, these same coefficients are

Table 14. Comparison between observed and calculated (3P) 4f and 5f levels

Level design.	4	f	5f		
	Calc.	Obs.	Calc.	Obs.	
$(^{3}P_{2})$ $[5]^{\circ}_{11/2}$ $[5]^{\circ}_{9/2}$	88392. 2	88392. 2	90873. 7	90873. 7	
	392. 3	392. 2	873. 8	873. 8	
$[4]_{9/2}^{9/2} \ [4]_{7/2}^{8}$	339. 3	339. 2	846. 4	845. 6	
	340. 4	340. 7	847. 3	846. 8	
$\begin{bmatrix} 3 \\ 7/2 \end{bmatrix}$	347. 9	351. 1	851. 1	851. 3	
$\begin{bmatrix} 3 \\ 5/2 \end{bmatrix}$	348. 5	350. 8	851. 5		
$[2]_{5/2}^{8/2}$	383. 6	371. 3	869. 7	869. 6	
$[2]_{3/2}^{8}$	383. 4	364. 0	869. 5	868. 8	
$[1]_{3/2}$	422. 3	425. 5	889. 6	891. 6	
$[1]_{1/2}$	421. 5	424. 9	888. 9	890. 9	
$(^{3}P_{1})$ $[4]_{9/2}^{\circ}$	91497. 7	91496. 2	93994. 0	93994. 0	
$[4]_{7/2}^{\circ}$ $[3]_{7/2}^{\circ}$	498. 4	497. 1	994. 5	994. 7	
	539. 3	545. 7	94015. 6	94017. 8	
	539. 9	542. 4	016. 1	018. 2	
$[3]_{5/2}^{\circ}$ $[2]_{5/2}^{\circ}$ $[2]_{3/2}^{\circ}$	484. 4	483. 0	93987. 3	93998. 4	
	483. 7	482. 3	986. 7	987. 6	
$[2]_{3/2}$ $(^{3}P_{0})$ $[3]_{7/2}^{\circ}$	92209. 5	92228. 3	94700. 7	94700. 2	
$[3]_{5/2}^{\circ}$	209. 9	228. 6	701. 0	700. 4	

Table 15. Comparison between observed and calculated (3P₂) 6f and 7f levels

Level design.	6	δf	7f		
	Calc.	Obs.	Calc.	Obs.	
$ \begin{array}{c c} (^3P_2) & [5]_{11/2}^{\circ} \\ [5]_{3/2}^{\circ} \\ [4]_{3/2}^{\circ} \\ [4]_{7/2}^{\circ} \\ [3]_{7/2}^{\circ} \\ [3]_{5/2}^{\circ} \\ [2]_{5/2}^{\circ} \\ [2]_{3/2}^{\circ} \\ [1]_{1/2}^{\circ} \end{array} $	92222. 7 222. 8 206. 9 207. 4 209. 6 209. 9 220. 5 220. 4 232. 0 231. 6	92222. 7 222. 8 207. 3 208. 3 } 187. 9 222. 0 221. 6 233. 1 232. 7	93036. 1 036. 1 026. 4 027. 8 028. 0 034. 7 034. 6 041. 7	93036. 1 036. 1 026. 1 026. 7 } 028. 4 035. 4 035. 0 042. 6 042. 4	

Table 16. The B and F² parameters for 4f, 5f, 6f, and 7f configurations, calculated as explained in the text

Conf.	В	F ²	$\mathrm{F}^2 imes n^3$
$\begin{array}{c} 4f \\ 5f \\ 6f \\ 7f \end{array}$	88371. 24 90862. 84 92216. 42 93032. 09	314. 1 162. 9 94. 8 59. 9	$\begin{array}{c} 201\!\times\!10^2 \\ 204 \\ 205 \\ 205 \end{array}$

valid in the discussion of $nf \times n'f$ interaction. Since this interaction is expected to be greater if a nondiagonal f'_2 is involved, we should expect that the strongest interactions would be of the types $(^{3}P_{1})nf[K] \stackrel{\circ}{\sim} \times (^{3}P_{2})n'f[K'] \stackrel{\circ}{\circ}$, where K=K'=2, 3, or 4, and $({}^{3}P_{0})nf[3]^{\circ}\times({}^{3}P_{2})n'f[3]^{\circ}$. Indeed, the plots in figure 8 show quite clearly that (3P₁)4f falls just between $(^{3}P_{2})5f$ and 6f and pushes $(^{3}P_{2})4f,5f$ levels down and $({}^{3}P_{2})6f,7f$. . . levels up, for K=2, 3, 4. Also, the $({}^{3}P_{0})4f[3]^{\circ} \times ({}^{3}P_{2})6f[3]^{\circ}$ interaction is, comparatively, very strong, owing to their almost identical unperturbed positions, as indicated by the calculations. It is possible that the designations of these levels should be interchanged. The points corresponding to (3P2)6f[3]° fall outside the range of figure 8 and have been omitted. Another striking feature of this figure is that it suggests, correctly, that $({}^{3}P_{1})5f[4]_{9/2}$ falls between $({}^{3}P_{2})9f,10f[4]_{9/2}$.

Superposed on the above interactions are those of the type $nf \times np$, which also seem to be stronger if the interacting levels have the same K-value.

The $({}^{1}D_{2})5p$ group appears to interact with nf levels having the appropriate K- and J-values, the largest such interaction being $({}^{1}D_{2})5p {}^{2}D_{5/2,3/2}^{\circ} \times ({}^{3}P_{2})4f[2]_{5/2,3/2}^{\circ}$, owing to the proximity of the levels. The several interactions affecting $({}^{3}P_{2})nf[3]^{\circ}$ pairs make the levels for $n \geq 5$ coincide within the experimental error. Since it is not possible to decide with certainty from which level of the pair a particular line arises, when that line is classifiable by either of the levels, the two levels are represented by a single energy in table 5.

The calculations suggest that the observed ($^{3}P_{2}$) $nf[1]^{\circ}$ pair is uniformly high. Since a K=1 is not possible from the ($^{3}P_{1}$) and ($^{3}P_{0}$) parents, we should expect the observed [1]° series to be almost Ritzian in the absence of interaction with np levels. Now the series is almost Ritzian, so that if the series is perturbed, it must be of a complicated sort, involving several np levels, since a single interaction with ($^{1}D_{2}$)5p2 $^{2}P_{3/2,1/2}^{\circ}$ would not account for the preservation of the rather strict linearity found in figure 8.

The $(^3P_2)8p$ $^4D_{7/2}^{\circ}\times(^3P_1)4f[3]_{7/2}^{\circ}$ interaction, complicated by additional influences on these levels, should also be mentioned, since it is possible that the two designations should be interchanged. It is not entirely clear what causes the $(^3P_1)4f$,5 $f[3]^{\circ}$ levels to be observed so "much" higher than their calculated positions.

7. Conclusion

Since the Br I line list has now been essentially exhausted—only 26 of 1253 lines remain unclassified—it is clear that any significant additions to the analysis of Br I will demand a thorough, new observation of the spectrum, made with a much stronger source for reasonable exposure times in the red, and one that produces Br₂ emission of much less intensity than the source used in this investigation. Even without making accurate calculations, one can predict very closely, barring strong interactions, the positions of the remaining unknown ns, np, nd,

and nf levels that fall below the principal ionization This can be done almost by inspection of tables 5 and 6. One can, then, make educated guesses as to the wavelength region where the stronger lines involving the unknown levels will lie. Of course many of the new lines that would be included in any augmented line list would be accounted for by transitions predicted in the present square array, and might help us in a possible revision of the more insecure level designations mentioned above. A study of the Zeeman effect in Br I is possibly the only and certainly the best method of resolving a few difficulties of interpretation, such as that of the J=1/2 levels lying in the vicinity of $({}^{1}S_{0})5s {}^{2}S_{1/2}$ as discussed above.

Perhaps the most demanding study that could be initiated on Br i is an extensive analysis of the nd configurations and their coupling and perturbations. If some of the levels still missing from the $({}^{3}P_{2})9d$ -11d configurations could be found, a more definite decision could possibly be reached as to whether some levels presently assigned to (³P₁)6d might not

more properly belong to $({}^{3}P_{2})9d$.

Of the 26 lines in our list that have defied classification, one has intensity 50, one intensity 15 w, and the rest have intensities 10 and less. It is probable that as many as one-half of these lines actually do not arise from neutral atomic bromine. For example, since the six unclassified lines of wavelengths less than 4100 Å would have to result from transitions to the even (³P)5s levels in order for the upper levels to lie below the (³P₂) limit, and since no new odd levels predictable on this basis fall in acceptable positions, it is doubtful that these six lines belong to Br 1.

There seems to be little hope of finding any Br 1 levels above the (3P2) limit unless autoionization from these levels is somehow avoided. Although we have predicted and looked in our list for many lines that we thought most likely to appear from such levels, none could be established definitely.

The unknown ng levels falling below the $({}^{3}P_{2})$ limit are predicted to lie in the range 90890 to 95285. Their stronger combinations will generally be with nf levels. Since these latter levels are in the range 88300 to 95285 K combinations with ng will produce lines that fall in the infrared beyond the region accessible by photography.

For suggesting this problem and offering advice throughout, I thank C. C. Kiess. His interest in the bromine spectrum has lasted over 30 years, and his excitement and enthusiasm as he watched our analysis progress was a constant source of encouragement.

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R. E. Trees examined and improved upon the discussion of the sp^6 configuration. W. R. Bozman frequently advised and instructed me in the various computations by the electronic computer and also took charge of the preparation of prints for the figures in this paper. W. C. Martin critically read the manuscript and offered many useful suggestions. Mrs. Ruth Peterson patiently and accurately carried out some of the final hand calculations, and Mrs. Sylvia Shure carefully typed the manuscript. It is a pleasure to thank each of these colleagues for his or her contribution to this paper.

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length observations.

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9. Appendix

 ${\bf Table \ A1.} \quad {\it Wavelengths \ of \ Br \ i \ in \ the \ vacuum \ ultraviolet}$

	111. 77	accenge	ms of Di i	in the vacaum attraviolet
Observed wave- length	Calc. wave- length	Intensity	Observed wave number	Classification
Å			K	
1067. 559 1067. 805 1068. 256 1068. 849 1073. 912	0. 556 . 802 . 263 . 856 . 916	$ \begin{array}{c} 10 \\ 15 \\ 10 \\ 3 \\ 70 \end{array} $	93671. 6 93650. 1 93610. 5 93558. 6 93117. 5	$\begin{array}{c} 4p^5 ^2\mathrm{P_{3/2}} \!$
1074. 243 1074. 803 1075. 345 1076. 964 1077. 873	$\begin{array}{c} .\ 245 \\ .\ 815 \\ \{\ .\ 317 \\ .\ 354 \\ .\ 964 \\ .\ 878 \end{array}$	70 10 15 20 15	93088. 8 93040. 3 92993. 4 92853. 6 92775. 3	$\begin{array}{c} 4p^5 {}^2\mathrm{P}_{3/2} \!$
1078. 124 1078. 242 1079. 320 1080. 882 1084. 810	. 124 . 247 . 326 . 883 . 807	15 30 30 50 10	92753. 7 92743. 6 92650. 9 92517. 0 92182. 0	$\begin{array}{l} 4p^5 {}^2\mathrm{P}_{3/2}^* - 8d {}^4\mathrm{D}_{5/2} \\ 4p^5 {}^2\mathrm{P}_{3/2}^* - 8d {}^4\mathrm{F}_{5/2} \\ 4p^5 {}^2\mathrm{P}_{3/2}^* - 8d {}^4\mathrm{D}_{1/2} \\ 4p^5 {}^2\mathrm{P}_{3/2}^* - ({}^1\mathrm{D}_2) 4d {}^2\mathrm{D}_{3/2} \\ 4p^5 {}^2\mathrm{P}_{3/2}^* - 9s {}^4\mathrm{P}_{3/2} \end{array}$
1085. 050 1085. 896 1087. 468 1087. 687 1087. 819	. 052 . 899 . 470 . 690 . 821	$ \begin{array}{r} 15 \\ 5 \\ 30 \\ 40 \\ 20 \end{array} $	92161. 6 92089. 8 91956. 7 91938. 2 91927. 1	$\begin{array}{c} 4p^{5} {}^{2}\mathrm{P}_{3/2}^{*} - 9s {}^{4}\mathrm{P}_{5/2}^{*} \\ 4p^{5} {}^{2}\mathrm{P}_{3/2}^{*} - 7s'' {}^{2}\mathrm{P}_{1/2}^{*} \\ 4p^{5} {}^{2}\mathrm{P}_{3/2}^{*} - 7d {}^{4}\mathrm{F}_{5/2}^{*} \\ 4p^{5} {}^{2}\mathrm{P}_{3/2}^{*} - ({}^{1}\mathrm{S}_{0})5s {}^{2}\mathrm{S}_{1/2} \\ 4p^{5} {}^{2}\mathrm{P}_{3/2}^{*} - 7d {}^{4}\mathrm{F}_{3/2}^{*} \end{array}$
1089. 039 1089. 203 1089. 322 1090. 623 1094. 722	. 040 . 203 . 323 . 620 . 719	30 15 20 85 100	91824. 1 91810. 2 91800. 2 91690. 7 91347. 4	$\begin{array}{l} 4p^{5} {}^{2}P_{3/2} \!$
1095. 481 1096. 788 1098. 881 1101. 347 a 1101. 456	. 482 . 790 . 885 . 351	100 100 100 50	91284. 1 91175. 3 91001. 7 90797. 9 90788. 9	$\begin{array}{l} 4p^5 {}^2\mathrm{P}_{3/2} \!$
1101. 498 1103. 924 1104. 168 1105. 460 1105. 844	. 932 . 169 . 464 . 846	250 60 60 50 50	90785. 5 90585. 9 90565. 9 90460. 1 90428. 7	$\begin{array}{l} 4p^5 ^2\mathrm{P}_{3/2}^\circ \! - \! (^1\mathrm{D}_2) 6s ^2\mathrm{D}_{3/2} \\ 4p^5 ^2\mathrm{P}_{3/2}^\circ \! - \! 5d' ^4\mathrm{P}_{5/2} \\ 4p^5 ^2\mathrm{P}_{3/2}^\circ \! - \! 5d' ^2\mathrm{F}_{5/2} \\ 4p^5 ^2\mathrm{P}_{3/2}^\circ \! - \! 5d' ^2\mathrm{P}_{3/2} \\ 4p^5 ^2\mathrm{P}_{3/2}^\circ \! - \! 6d ^4\mathrm{P}_{1/2} \end{array}$
1105. 994 1107. 442 1107. 512 1109. 422 1110. 904	. 994 . 445 . 514 . 432 . 911	70 50 50 20 100	90416. 4 90298. 2 90292. 5 90137. 0 90016. 8	$\begin{array}{c} 4p^5 {}^2\mathrm{P}_{3/2}^* \! - \! 6d {}^4\mathrm{F}_{3/2} \\ 4p^5 {}^2\mathrm{P}_{1/2}^* \! - \! 8s' {}^4\mathrm{P}_{1/2} \\ 4p^5 {}^2\mathrm{P}_{3/2}^* \! - \! 5d' {}^2\mathrm{P}_{1/2} \\ 4p^5 {}^2\mathrm{P}_{1/2}^* \! - \! 10d {}^4\mathrm{D}_{3/2} \\ 4p^5 {}^2\mathrm{P}_{1/2}^* \! - \! ({}^1\mathrm{D}_2) 4d {}^2\mathrm{S}_{1/2} \end{array}$
b 1111. 579 1111. 751 1112. 743 1115. 448 1116. 105	. 542 . 754 . 747 . 448 . 105	50 40 20 30 40	89962. 1 89948. 2 89868. 0 89650. 1 89597. 3	$\begin{array}{l} 4p^5 ^2\mathrm{P}_{1/2} \!$
1118. 173 1119. 140 1119. 725 1121. 473 1121. 839	. 169 . 143 . 728 . 474 . 842	90 60 90 120 80	89431. 6 89354. 3 89307. 6 89168. 4 89139. 4	$\begin{array}{l} 4p^5 ^2\mathrm{P}_{1/2}^\circ \! - \! 6d' ^2\mathrm{P}_{3/2} \\ 4p^5 ^2\mathrm{P}_{1/2}^\circ \! - \! 10s ^4\mathrm{P}_{3/2} \\ 4p^5 ^2\mathrm{P}_{1/2}^\circ \! - \! 6d' ^2\mathrm{P}_{1/2} \\ 4p^5 ^2\mathrm{P}_{1/2}^\circ \! - \! 8d ^4\mathrm{F}_{3/2} \\ 4p^5 ^2\mathrm{P}_{1/2}^\circ \! - \! 8d ^4\mathrm{P}_{1/2} \end{array}$
1124. 038 1125. 728 1129. 979 1131. 171 1132. 822	. 035 . 725 . 981 . 166 . 817	120 200 35 80 100	88965. 0 88831. 4 88497. 2 88404. 0 88275. 1	$\begin{array}{c} 4p^5 {}^2\mathrm{P}_{1/2}^\circ - 8d {}^4\mathrm{D}_{1/2} \\ 4p^5 {}^2\mathrm{P}_{1/2}^\circ - ({}^1\mathrm{D}_2) 4d {}^2\mathrm{D}_{3/2} \\ 4p^5 {}^2\mathrm{P}_{1/2}^\circ - 9s {}^4\mathrm{P}_{3/2} \\ 4p^5 {}^2\mathrm{P}_{1/2}^\circ - 7s'' {}^2\mathrm{P}_{1/2} \\ 4p^5 {}^2\mathrm{P}_{3/2}^\circ - 7s {}^4\mathrm{P}_{3/2} \end{array}$

Table A1. Wavelengths of Br 1 in the vacuum ultraviolet—Con.

Observed wave- length	Calc. wave- length	sity	Observed wave number	Classification
Å			K	
1133. 116 1133. 251	0. 110 . 252	200 80	88252. 2 88241. 7	$rac{4p^{52}\mathrm{P}_{1/2}^{\circ}-\!$
1134. 588 1134. 888 1136. 294	\begin{cases} \ \ . 575 \ . 622 \ \ . 882 \ \ . 290 \end{cases}	300 200 250	88137. 7 88114. 4 88005. 4	$\begin{array}{l} \left\{4p^{5} {}^{2}\mathrm{P}_{1/2}^{\circ} - 7d {}^{4}\mathrm{P}_{1/2} \right. \\ \left. \left\{4p^{5} {}^{2}\mathrm{P}_{3/2}^{\circ} - 7s {}^{4}\mathrm{P}_{5/2}^{\circ} \right. \\ \left. \left. \left\{4p^{5} {}^{2}\mathrm{P}_{1/2}^{\circ} - 7d {}^{4}\mathrm{D}_{3/2}^{\circ} \right. \right. \\ \left. \left. \left\{4p^{5} {}^{2}\mathrm{P}_{1/2}^{\circ} - 7d {}^{4}\mathrm{D}_{1/2}^{\circ} \right. \right. \end{array} \right. \end{array}$
1139. 350 1139. 544 1140. 732 1141. 564 1145. 268	. 342 {. 537 {. 547 . 739 . 568 . 264	100 120 20 25 80	87769. 3 87754. 4 87663. 0 87599. 1 87315. 8	$\begin{array}{c} 4p^5 {}^2\mathrm{P}_{3/2} - \!$
1145. 854 1147. 689 1147. 943 1150. 312 1151. 381	. 846 . 689 . 943 . 307 . 380	60 80 60 80 100	87271. 2 87131. 6 87112. 3 86932. 9 86852. 2	$\begin{array}{l} 4p^5 {}^2\mathrm{P}_{3/2}^* \! - \! 5d {}^4\mathrm{D}_{1/2} \\ 4p^5 {}^2\mathrm{P}_{3/2}^* \! - \! 5d {}^4\mathrm{D}_{3/2} \\ 4p^5 {}^2\mathrm{P}_{1/2}^* \! - \! 8s {}^4\mathrm{P}_{3/2} \\ 4p^5 {}^2\mathrm{P}_{3/2}^* \! - \! 5d {}^4\mathrm{D}_{5/2} \\ 4p^5 {}^2\mathrm{P}_{1/2}^* \! - \! 5d' {}^4\mathrm{P}_{3/2} \end{array}$
1152. 418 1152. 833 1152. 989 1154. 640 1159. 030	. 412 . 827 . 988 . 640 . 025	$100 \\ 40 \\ 60 \\ 15 \\ 80$	86774. 1 86742. 8 86731. 1 86607. 1 86279. 0	$\begin{array}{c} 4p^5 {}^2\mathrm{P}_{1/2}^\circ \! - \! 5d' {}^2\mathrm{P}_{3/2}^\circ \\ 4p^5 {}^2\mathrm{P}_{1/2}^\circ \! - \! 6d {}^4\mathrm{P}_{1/2}^\circ \\ 4p^5 {}^2\mathrm{P}_{1/2}^\circ \! - \! 6d {}^4\mathrm{F}_{3/2}^\circ \\ 4p^5 {}^2\mathrm{P}_{1/2}^\circ \! - \! 5d' {}^2\mathrm{P}_{1/2}^\circ \\ 4p^5 {}^2\mathrm{P}_{3/2}^\circ \! - \! 6s'' {}^2\mathrm{P}_{1/2}^\circ \end{array}$
1160. 332 1168. 542 1170. 479 1173. 827 1177. 233	. 329 . 539 . 476 . 825 . 230	$100 \\ 150 \\ 150 \\ 175 \\ 250$	86182. 2 85576. 7 85435. 1 85191. 4 84945. 0	$\begin{array}{c} 4p^5^2\mathrm{P}_{1/2}^*\!-\!6d^4\mathrm{D}_{3/2} \\ 4p^5^2\mathrm{P}_{3/2}^*\!-\!6s'^4\mathrm{P}_{1/2} \\ 4p^5^2\mathrm{P}_{3/2}^*\!-\!6s'^2\mathrm{P}_{3/2} \\ 4p^5^2\mathrm{P}_{3/2}^*\!-\!4d''^2\mathrm{D}_{3/2} \\ 4p^5^2\mathrm{P}_{3/2}^*\!-\!4d''^2\mathrm{D}_{5/2} \end{array}$
1178. 895 1182. 171 1186. 161 1189. 279 1189. 378	. 889 . 169 . 159 . 277 . 379	$ \begin{array}{c} 400 \\ 150 \\ 150 \\ 1000 \\ 250 \end{array} $	84825. 2 84590. 1 84305. 6 84084. 6 84077. 6	$\begin{array}{c} 4p^5 {}^2\mathrm{P}_{3/2} - 4s4p^6 {}^2\mathrm{S}_{1/2} \\ 4p^5 {}^2\mathrm{P}_{1/2}^\circ - 7s {}^4\mathrm{P}_{3/2} \\ 4p^5 {}^2\mathrm{P}_{3/2}^\circ - 4d' {}^4\mathrm{P}_{5/2} \\ 4p^5 {}^2\mathrm{P}_{1/2}^\circ - 5d {}^4\mathrm{F}_{3/2} \\ 4p^5 {}^2\mathrm{P}_{3/2}^\circ - 4d' {}^4\mathrm{P}_{3/2} \end{array}$
1189. 498 1194. 413 1196. 370 1196. 477 1198. 371	. 500 . 414 . 366 . 480 . 374	1000 200 200 200 500	84069. 1 83723. 1 83586. 2 83578. 7 83446. 6	$\begin{array}{c} 4p^5 {}^2\mathrm{P}_{1/2}^{\circ} - 5d {}^4\mathrm{P}_{1/2} \\ 4p^5 {}^2\mathrm{P}_{3/2}^{\circ} - 4d' {}^2\mathrm{F}_{5/2} \\ 4p^5 {}^2\mathrm{P}_{1/2}^{\circ} - 5d {}^4\mathrm{D}_{1/2} \\ 4p^5 {}^2\mathrm{P}_{3/2}^{\circ} - 4d' {}^2\mathrm{P}_{3/2} \\ 4p^5 {}^2\mathrm{P}_{1/2}^{\circ} - 5d {}^4\mathrm{D}_{3/2} \end{array}$
1203. 353 1209. 756 1210. 734 1216. 006 1221. 128	. 354 . 752 . 739 . 010 . 125	750	83101. 1 82661. 3 82594. 5 82236. 4 81891. 5	$\begin{array}{c} 4p^5 {}^2\mathrm{P}_{3/2}^* - 4d' {}^2\mathrm{P}_{1/2} \\ 4p^5 {}^2\mathrm{P}_{3/2}^* - 6s {}^4\mathrm{P}_{3/2} \\ 4p^5 {}^2\mathrm{P}_{1/2}^* - 6s'' {}^2\mathrm{P}_{1/2} \\ 4p^5 {}^2\mathrm{P}_{3/2}^* - 6s {}^4\mathrm{P}_{5/2} \\ 4p^5 {}^2\mathrm{P}_{3/2}^* - 6s' {}^4\mathrm{P}_{1/2} \end{array}$
1221. 870 1223. 240 1224. 408 1226. 899 1228. 049	. 866 . 240 . 406 . 898 . 049	1200	81841. 8 81750. 1 81672. 1 81506. 3 81430. 0	$\begin{array}{c} 4p^5 ^2\mathrm{P}_{3/2}^{\circ} - 4d ^4\mathrm{F}_{3/2} \\ 4p^5 ^2\mathrm{P}_{1/2}^{\circ} - 6s' ^2\mathrm{P}_{3/2} \\ 4p^5 ^2\mathrm{P}_{3/2}^{\circ} - 4d ^4\mathrm{F}_{5/2} \\ 4p^5 ^2\mathrm{P}_{1/2}^{\circ} - 4d'' ^2\mathrm{D}_{3/2} \\ 4p^5 ^2\mathrm{P}_{3/2}^{\circ} - 4d ^4\mathrm{P}_{1/2} \end{array}$
1232. 431 1243. 897 1249. 589 1251. 664 1255. 799	. 432 . 901 . 588 . 670 . 798	$1200 \\ 800 \\ 1500$	81140. 4 80392. 5 80026. 3 79893. 6 79630. 6	$\begin{array}{c} 4p^5 ^2\mathrm{P}_{1/2}^{\circ} - 484p^6 ^2\mathrm{S}_{1/2} \\ 4p^5 ^2\mathrm{P}_{1/2}^{\circ} - 4d' ^4\mathrm{P}_{3/2} \\ 4p^5 ^2\mathrm{P}_{3/2}^{\circ} - 4d ^4\mathrm{D}_{1/2} \\ 4p^5 ^2\mathrm{P}_{3/2}^{\circ} - 4d' ^2\mathrm{P}_{3/2} \\ 4p^5 ^2\mathrm{P}_{3/2}^{\circ} - 4d' ^4\mathrm{D}_{3/2} \end{array}$
1259. 199 1261. 658 1266. 200 1279. 477 1286. 259	. 194 . 656 . 202 . 480 . 261	$1200 \\ 1200 \\ 1000$	79415. 6 79260. 8 78976. 5 78156. 9 77744. 8	$\begin{array}{c} 4p^5 {}^2\mathrm{P}_{1/2}^{\circ} - 4d' {}^2\mathrm{P}_{1/2} \\ 4p^5 {}^2\mathrm{P}_{3/2}^{\circ} - 4d {}^4\mathrm{D}_{5/2} \\ 4p^5 {}^2\mathrm{P}_{1/2}^{\circ} - 6s {}^4\mathrm{P}_{3/2} \\ 4p^5 {}^2\mathrm{P}_{1/2}^{\circ} - 4d {}^4\mathrm{F}_{3/2} \\ 4p^5 {}^2\mathrm{P}_{1/2}^{\circ} - 4d {}^4\mathrm{P}_{1/2} \end{array}$

Table A1. Wavelengths of Brin the vacuum ultraviolet—Con. | Table A1. Wavelengths of Brin the vacuum ultraviolet—Con

Observed wave- length	Calc. wave- length	sity	Observed wave number	Classification
Å			K	
1574. 841 1576. 387 1582. 312 1633. 404	. 381	$30000 \\ 20000 \\ 25000 \\ 75000$	63498. 5 63436. 2 63198. 7 61221. 8	$rac{4p^5}{4p^5} rac{^2 ext{P}_{3/2}^{5/2} - 5s}{4p^5} rac{^4 ext{P}_{5/2}^{5/2}}{5s'} rac{^4 ext{P}_{1/2}}{4}$

Observed wave- length	wave-	sity	Observed wave number	Classification
$ m \mathring{A}$			K	
1309. 908 1316. 735 1317. 372 1317. 695 1384. 598	. 735 . 375 . 691		76341. 2 75945. 4 75908. 7 75890. 1 72223. 1	$\begin{array}{c} 4p^5 \ ^2P_{1/2}^{\circ}-4d \ ^4D_{1/2} \\ 4p^5 \ ^2P_{1/2}^{\circ}-4d \ ^4D_{3/2} \\ 4p^5 \ ^2P_{3/2}^{\circ}-(^1D_2) 5s \ ^2D_{3/2} \\ 4p^5 \ ^2P_{3/2}^{\circ}-(^1D_2) 5s \ ^2D_{5/2} \\ 4p^5 \ ^2P_{1/2}^{\circ}-(^1D_2) 5s \ ^2D_{3/2} \end{array}$
1449. 903 1488. 452 1495. 132 1531. 743 1540. 654	. 459 . 129 . 746	$ \begin{array}{r} 3000 \\ 50000 \\ 5 \\ 30000 \\ 25000 \end{array} $	68970. 1 67183. 9 66883. 7 65285. 1 64907. 5	$4p^5 {}^2P_{3/2} \! - \! 5s'' {}^2P_{1/2} \ 4p^5 {}^2P_{3/2} \! - \! 5s' {}^2P_{3/2} \ 4p^5 {}^2P_{3/2} \! - \! 5s' {}^4P_{1/2} \ 4p^5 {}^2P_{1/2} \! - \! 5s'' {}^2P_{1/2} \ 4p^5 {}^2P_{3/2} \! - \! 5s {}^4P_{3/2}$

 $^{^{\}rm a}\,\lambda 1101.456$ computed wavelength. The line is masked by $\lambda 1101.498.$

Table A2. Observed lines of Br I

TABLE RZ. Observed times by BI I				
Wavelength	Intensity	Wave number	Classification	
Å		K		
3325. 307 3348. 566 3400. 030 3400. 060 3402. 411	$ \begin{array}{c} 10w \\ 15 \\ 25 \\ 20 \\ 15 \end{array} $	30063, 76 29854, 95 29403, 07 29402, 81 29382, 49	$\left\{egin{array}{lll} 5s\ ^4 ext{P}_{5/2}7p''\ ^2 ext{P}_{3/2}^3 \ 5s\ ^4 ext{P}_{5/2}10p\ ^4 ext{D}_{1/2}^2 \ 5s\ ^4 ext{P}_{5/2}7p'\ ^4 ext{S}_{3/2}^2 \end{array} ight\}$	
3402. 436 3409. 728 3409. 753 3418. 888 3425. 577	$ \begin{cases} 10 \\ 15 \\ 10 \\ 15w \\ 60 \end{cases} $	29382, 28 29319, 44 29319, 23 29240, 89 29183, 79	$\left. egin{array}{lll} 58\ ^4 ext{P}_{5/2} ext{$-7p'\ ^2D_{3/2}^\circ$} \ & 58\ ^4 ext{P}_{5/2} ext{$-7p'\ ^2D_{5/2}^\circ$} \ & 58\ ^4 ext{P}_{5/2} ext{$-9p\ ^4D_{3/2}^\circ$} \ & 58\ ^4 ext{P}_{5/2} ext{$-9p\ ^4D_{7/2}^\circ$} \end{array} ight.$	
3428. 605 3429. 809 3472. 188 3472. 216 3496. 366	$ \left. \begin{array}{c} 15 \\ 8 \\ 20 \\ 10 \\ 5 \end{array} \right. $	29158. 02 29147. 79 28792. 04 28791. 81 28592. 94	$\left.\begin{array}{c} 58{}^{4}\mathrm{P}_{5/2} - 9p{}^{4}\mathrm{P}_{3/2}^{\circ} \\ 58{}^{4}\mathrm{P}_{5/2} - 9p{}^{4}\mathrm{P}_{5/2}^{\circ} \\ 58{}^{4}\mathrm{P}_{5/2} - 4f''[3]_{7/2}^{\circ} \\ 58{}^{4}\mathrm{P}_{3/2} - 7p''{}^{2}\mathrm{P}_{3/2}^{\circ} \end{array}\right\}$	
3497. 433 3516. 144 3541. 173 3556. 518 3556. 540	$ \begin{array}{c} 1 \\ 40 \\ 100 \\ 40 \\ 30 \end{array} $	28584. 22 28432. 12 28231. 16 28109. 36 28109. 19	$\left.\begin{array}{c} 5s\ ^4\mathrm{P}_{3/2}-7\ p''\ ^2\mathrm{P}_{1/2}^\circ\\ 5s\ ^4\mathrm{P}_{3/2}-10\ p\ ^4\mathrm{D}_{3/2}^\circ\\ 5s\ ^4\mathrm{P}_{5/2}-8\ p\ ^4\mathrm{D}_{3/2}^\circ\\ 5s\ ^4\mathrm{P}_{5/2}-4f'\ [3]_{7/2}^\circ\end{array}\right\}$	
3557, 461 3558, 616 3563, 706 3564, 468 3564, 498	$ \begin{array}{c} 250 \\ 75 \\ 50 \\ 15w \\ 10 \end{array} $	28101. 91 28092. 79 28052. 66 28046. 67 28046. 43	$\left.\begin{array}{c} 58^{4}\mathrm{P}_{5/2} - 8p^{4}\mathrm{D}_{7/2}^{\circ} \\ 58^{4}\mathrm{P}_{5/2} - 8p^{4}\mathrm{P}_{3/2}^{\circ} \\ 58^{4}\mathrm{P}_{5/2} - 8p^{4}\mathrm{P}_{3/2}^{\circ} \\ 58^{4}\mathrm{P}_{5/2} - 4f'[2]_{5/2}^{\circ} \end{array}\right\}$	
3579, 081 3581, 712 3589, 818 3599, 962 3606, 518	$ \begin{array}{r} 40 \\ 15 \\ 100 \\ 100 \\ 20 \end{array} $	27932. 16 27911. 64 27848. 62 27770. 15 27719. 67	$\begin{array}{c} 58\ ^4\mathrm{P}_{3/2} -\!-\!7p'\ ^4\mathrm{S}_{5/2}^{*} \\ 58\ ^4\mathrm{P}_{3/2} -\!7p'\ ^2\mathrm{D}_{3/2}^{*} \\ 58\ ^4\mathrm{P}_{3/2} -\!7p'\ ^2\mathrm{D}_{3/2}^{*} \\ 58\ ^4\mathrm{P}_{3/2} -\!9p\ ^4\mathrm{D}_{5/2}^{*} \\ 58\ ^4\mathrm{P}_{3/2} -\!9p\ ^4\mathrm{D}_{5/2}^{*} \end{array}$	
3608, 864 3644, 165 3644, 198 3646, 605 3646, 626	$ \left. \begin{array}{c} 7 \\ 15 \\ 10 \\ 100 \\ 80 \end{array} \right. $	27701. 65 27433. 31 27433. 06 27414. 96 27414. 80	$ \begin{cases} 5s \ ^{4}\mathrm{P}_{3/2} - 9p \ ^{4}\mathrm{P}_{1/2}^{\circ} \\ 5s \ ^{4}\mathrm{P}_{5/2} - 5f \ [2]_{5/2}^{\circ} \\ \\ 5s \ ^{4}\mathrm{P}_{5/2} - 5f \ [3]_{7/2,5/2}^{\circ} \end{cases} $	
3658. 486 3659. 970 3735. 800 3745. 419 3751. 287	$ \begin{array}{c c} 2h \\ 20 \\ 500 \\ 50 \\ 10 \end{array} $	27325. 93 27314. 85 26760. 42 26691. 70 26649. 94	$\begin{array}{c} 5s\ ^4\mathrm{P}_{3/2}-6f\ [1]_{3/2}^3\\ 5s\ ^4\mathrm{P}_{3/2}-6f\ [2]_{5/2}^3\\ 5s\ ^4\mathrm{P}_{3/2}-8p\ ^4\mathrm{P}_{3/2}^3\\ 5s\ ^4\mathrm{P}_{3/2}-8p\ ^4\mathrm{P}_{7/2}^3\\ 5s\ ^4\mathrm{P}_{3/2}-8p\ ^4\mathrm{P}_{5/2}^6\end{array}$	

^b λ1111.579 this may be a Br II line.

Table A2. Observed lines of Br I—Continued

	TABLE A2.	Ooserved times of 1	
Wavelength	Intensity	Wave number	Classification
Å		K	
3753. 351 3756. 032 3757. 265 3761. 746 3761. 847	5 30 5 2 1	26635. 29 26616. 28 26607. 54 26575. 85 26575. 14	$\begin{array}{c} 58\ ^4\mathrm{P}_{3/2} -\!$
3770. 88 3794. 03 3798. 805 3810. 95 3815. 650	$ \begin{array}{c} 150c \\ 900c \\ 5w \\ 1h \\ 1200 \end{array} $	26511. 48 26349. 72 26316. 60 26232. 73 26200. 42	$\begin{array}{c} 5s\ ^4\mathrm{P}_{5/2}\!\!-\!6p''\ ^2\mathrm{P}_{3/2}^*\\ 5s\ ^4\mathrm{P}_{5/2}\!\!-\!7p\ ^4\mathrm{D}_{3/2}^*\\ 5s'\ ^2\mathrm{P}_{3/2}\!\!-\!7p''\ ^2\mathrm{P}_{3/2}^*\\ 5s\ ^4\mathrm{P}_{5/2}\!\!-\!7p\ ^4\mathrm{D}_{5/2}^*\\ 5s\ ^4\mathrm{P}_{5/2}\!\!-\!7p\ ^4\mathrm{D}_{7/2}^*\\ \end{array}$
3828. 505 3829. 75 3844. 035 3847. 37 3850. 619	$700 \ 200c \ 100 \ 5h \ 10$	26112, 45 26103, 96 26006, 95 25984, 41 25962, 49	$\begin{array}{c} 5s\ ^4\mathrm{P}_{5/2} \!$
3850, 740 3851, 654 3854, 702 3896, 653 3899, 767	$ \begin{array}{c} 1\\ 3\\ 90\\ 200\\ 25 \end{array} $	25961. 67 25955. 51 25934. 99 25655. 78 25635. 29	$\begin{array}{c} 5s\ ^4\mathrm{P}_{3/2} -\!$
3905. 823 3905. 858 3909. 385 3913. 560 3917. 775	$ \begin{cases} 25w \\ 20 \\ 175 \\ 12 \\ 200 \end{cases} $	25595. 55 25595. 32 25572. 23 25544. 95 25517. 47	$\left. egin{array}{lll} 5s\ ^4 ext{P}_{5/2}6p'\ ext{S}_{3/2}' \ 5s'\ ^2 ext{P}_{3/2}7p'\ ^2 ext{S}_{1/2}' \ 5s'\ ^2 ext{P}_{3/2}7p'\ ^2 ext{S}_{1/2}' \end{array} ight.$
3917. 819 3929. 196 3934. 068 3934. 084 3934. 119	$ \begin{array}{c c} & 150 \\ & 5 \\ & 60 \\ & 80 \\ & 100 \end{array} $	25517. 18 25443. 30 25411. 79 25411. 68 25411. 46	$ \left. \begin{array}{c} 58 \ ^{1} \ ^{5/2} \ ^{2} \ ^{2} \ ^{3/2} \\ 58' \ ^{2} \ ^{2} \ ^{3/2} - 9p \ ^{4} \ ^{6} \ ^{6} \ ^{2} \\ 58 \ ^{4} \ ^{2} \ ^{5/2} - 6p' \ ^{2} \ ^{2} \ ^{2} \ ^{3/2} \end{array} \right. $
3942. 026 3943. 776 3943. 835 3945. 565 3969. 416	$egin{array}{c} 2h \\ 2 \\ 8 \\ 4 \\ 1 \end{array}$	25360, 49 25349, 23 25348, 86 25337, 74 25185, 50	$\begin{array}{c} 5s' \ ^4\mathrm{P}_{1/2} -\!$
3969. 434 3973. 967 3991. 363 3992. 363 3993. 177	1 5 300 1500 5	25185, 38 25156, 66 25047, 01 25040, 74 25035, 64	$rac{5s}{5s}rac{4 ext{P}_{5/2} ext{}(^1 ext{D}_2)}{5s}rac{2 ext{D}_{5/2}^2 ext{}6p^{\prime\prime}}{2 ext{P}_{3/2}^2 ext{}6p^{\prime\prime}}$
3998, 492 3999, 070 4000, 599 4009, 296 4009, 325	$ \begin{array}{c} 35 \\ 200 \\ 2w \\ 2w \\ 2 \end{array} $	25002. 36 24998. 75 24989. 19 24934. 99 24934. 81	$\left.\begin{array}{c} 58\ ^{4}\mathrm{P}_{5/2}(^{1}\mathrm{D}_{2})5p\ ^{2}\mathrm{D}_{3/2}^{\circ}\\ 58\ ^{4}\mathrm{P}_{3/2}6p''\ ^{2}\mathrm{P}_{3/2}^{\circ}\\ 58\ ^{4}\mathrm{P}_{5/2}4f\ [1]_{3/2}^{\circ}\\ \end{array}\right\}$
4012. 548 4012. 580 4016. 453 4018. 316 4021. 772	$ \begin{cases} 60s \\ 50 \\ 1h \\ 200 \\ 250 \end{cases} $	24914. 78 24914. 58 24890. 55 24879. 01 24857. 64	$\left.\begin{array}{l} 5s\ ^{4}\mathrm{P}_{5/2}4f\ [3]_{7/2}^{\circ} \\ \\ 5s\ ^{4}\mathrm{P}_{3/2}7p\ ^{4}\mathrm{P}_{3/2}^{\circ} \\ 5s\ ^{4}\mathrm{P}_{3/2}7p\ ^{4}\mathrm{P}_{1/2}^{\circ} \end{array}\right.$
4033, 762 4037, 324 4056, 404 4056, 425 4057, 020	$ \begin{array}{c} 1h \\ 500 \\ 30s \\ 25 \\ 30 \end{array} $	24783. 75 24761. 89 24645. 42 24645. 29 24641. 67	$\left.\begin{array}{c} 5s' \ ^4\mathrm{P}_{1/2} - 8p \ ^4\mathrm{D}_{3/2} \\ 5s \ ^4\mathrm{P}_{3/2} - 7p \ ^4\mathrm{D}_{3/2} \\ 5s' \ ^4\mathrm{P}_{1/2} - 8p \ ^4\mathrm{P}_{3/2} \\ 5s \ ^4\mathrm{P}_{3/2} - 7p \ ^4\mathrm{P}_{5/2} \\ \end{array}\right\}$
4058, 406 4064, 150 4075, 503 4076, 956 4083, 149	75 4 50 75 50	24633. 26 24598. 44 24529. 92 24521. 18 24483. 99	$\begin{array}{c} 58\ ^4\mathrm{P}_{3/2}7\ p\ ^4\mathrm{P}_{3/2}^{\circ} \\ 58'\ ^4\mathrm{P}_{1/2}4f'\ [2]_{3/2}^{\circ} \\ 58''\ ^2\mathrm{P}_{1/2}7p''\ ^2\mathrm{P}_{3/2}^{\circ} \\ 58'\ ^2\mathrm{P}_{1/2}7p''\ ^2\mathrm{P}_{3/2}^{\circ} \\ 58'\ ^2\mathrm{P}_{3/2}8p\ ^4\mathrm{D}_{3/2}^{\circ} \end{array}$

Table A2. Observed lines of Br I—Continued

	TABLE AZ.	Observed times of .	Bi i—Continued
Wavelength	Intensity	Wave number	Classification
Å		K	
4094, 643 4106, 352 4113, 132 4114, 165 4124, 233	$ \begin{array}{r} 35 \\ 25 \\ 4 \\ 25 \\ 75 \end{array} $	24415. 26 24345. 65 24305. 52 24299. 41 24240. 09	$egin{array}{cccccccccccccccccccccccccccccccccccc$
4143. 974 4157. 425 4164. 158 4164. 270 4168. 115	$ \begin{array}{r} 600 \\ 250 \\ 4 \\ 15 \\ 10 \end{array} $	24124. 62 24046. 57 24007. 69 24007. 04 23984. 90	$\begin{array}{c} 5s\ ^4\mathrm{P}_{3/2} -\!$
4175. 786 4179. 322 4191. 935 4196. 468 4199. 473	50 100 50 50 35	23940. 84 23920. 58 23848. 61 23822. 85 23805. 80	$\begin{array}{c} 5s\ ^4\mathrm{P}_{3/2} -\!$
4202. 487 4220. 701 4220. 859 4231. 643 4240. 366	$ \begin{array}{c} 450 \\ 5h \\ 2 \\ 4 \\ 40 \end{array} $	23788. 73 23686. 08 23685. 19 23624. 83 23576. 23	$\begin{array}{c} 5s\ ^4\mathrm{P}_{5/2}(^1\mathrm{D}_2)\ 5p\ ^2\mathrm{F}_{7/2}^2\\ 5s'\ ^2\mathrm{P}_{3/2}5f\ [2]_{5/2}^2\\ 5s'\ ^2\mathrm{P}_{3/2}5f\ [2]_{3/2}^2\\ 5s\ ^4\mathrm{P}_{5/2}(^1\mathrm{D}_2)\ 5p\ ^2\mathrm{F}_{5/2}^2\\ 5s\ ^4\mathrm{P}_{3/2}(^1\mathrm{D}_2)\ 5p\ ^2\mathrm{D}_{5/2}^2\\ \end{array}$
4248. 414 4250. 818 4250. 924 4260. 625 4261. 964	50 40 15 40 3	23531, 57 23518, 26 23517, 68 23464, 13 23456, 76	$\begin{array}{c} 5s\ ^4\mathrm{P}_{3/2}(^1\mathrm{D}_2)5p\ ^2\mathrm{D}^3_{3/2} \\ 5s\ ^4\mathrm{P}_{3/2}4f\ [1]^3_{3/2} \\ 5s\ ^4\mathrm{P}_{3/2}4f\ [1]^1_{1/2} \\ 5s\ ^4\mathrm{P}_{3/2}4f\ [2]^2_{3/2} \\ 5s\ ^4\mathrm{P}_{3/2}4f\ [2]^3_{3/2} \end{array}$
4334. 54 4342. 45 4365. 14 4369. 22 4391. 60	$20 \\ 2000 \\ 75 \\ 800$	23064. 01 23022. 00 22902. 34 22880. 95 22764. 35	$egin{array}{cccccccccccccccccccccccccccccccccccc$
4399. 73 4404. 57 4412. 49 4423. 03 4425. 14	450 10 100 200 1500	22722. 28 22697. 32 22656. 58 22602. 59 22591. 81	$\begin{array}{c} 5s'\ ^2\mathrm{P}_{3/2} -\!$
4427. 20 4431. 59 4441. 74 b4446. 08 4466. 21	$ \begin{array}{r} 40 \\ 1 \\ 10000 \\ 15 \\ 200 \end{array} $	22581, 30 22558, 93 22507, 38 22485, 41 22384, 07	$\begin{array}{c} 5s'\ ^2\mathrm{P}_{3/2} \!$
4469. 98 4471. 67 4472. 61 4477. 72 4490. 42	$ \begin{array}{c} 150 \\ 80 \\ 10000 \\ 20000 \\ 1000 \end{array} $	22365. 19 22356. 74 22352. 04 22326. 53 22263. 39	$\begin{array}{c} 5s'\ ^2\mathrm{P}_{3/2} -\!$
4513. 44 4513. 82 4525. 59 4529. 79 4575. 74	$ \begin{array}{r} 3000 \\ 150 \\ 15000 \\ 600 \\ 3000 \end{array} $	22149. 84 22147. 97 22090. 37 22069. 89 21848. 27	$\begin{array}{c} 5s\ ^4\mathrm{P}_{5/2}6p\ ^4\mathrm{P}_{3/2}^{\circ} \\ 5s'\ ^4\mathrm{P}_{1/2}6p'\ ^4\mathrm{S}_{3/2}^{\circ} \\ 5s\ ^4\mathrm{P}_{5/2}6p\ ^4\mathrm{P}_{5/2}^{\circ} \\ 5s'\ ^4\mathrm{P}_{1/2}6p'\ ^2\mathrm{D}_{3/2}^{\circ} \\ 5s'\ ^2\mathrm{P}_{3/2}6p'\ ^4\mathrm{S}_{3/2}^{\circ} \end{array}$
4592. 16 4614. 58 4640. 92 4641. 02 4643. 52	$200 \\ 2500 \\ 40 \\ 70 \\ 900$	21770. 15 21664. 38 21541. 42 21540. 96 21529. 36	$egin{array}{cccccccccccccccccccccccccccccccccccc$

Table A2. Observed lines of Br 1—Continued

	TABLE A2.	Observed lines of 1	Br 1—Continued
Wavelength	Intensity	Wave number	Classification
Å		K	
b4654. 18 4693. 56 4703. 42 4706. 36 4720. 03	$ \begin{array}{r} 25 \\ 300 \\ 50 \\ 12h \\ 100 \end{array} $	21480, 05 21299, 83 21255, 18 21241, 90 21180, 38	$\begin{array}{c} 5s' \ ^{4}\mathrm{P}_{1/2} -\!$
4722. 97 4752. 28 4765. 63 4775. 20 4780. 31	$\begin{array}{c} 9 \\ 2500 \\ 250 \\ 750 \\ 4000 \end{array}$	21167. 20 21036. 65 20977. 72 20935. 68 20913. 30	$\begin{array}{c} 5s'\ ^2\mathrm{P}_{3/2} -\!$
4785. 19 4802. 67 4807. 61 4834. 46 4848. 39	$ \begin{array}{c} 1600 \\ 250 \\ 350 \\ 500 \\ 25 \end{array} $	20891. 97 20815. 93 20794. 54 20679. 06 20619. 64	$\begin{array}{c} 58{}^{4}\mathrm{P}_{3/2}6p{}^{4}\mathrm{P}_{1/2}^{7} \\ 58''{}^{2}\mathrm{P}_{1/2}7p{}^{4}\mathrm{D}_{3/2}^{2} \\ 58''{}^{2}\mathrm{P}_{1/2}7p{}^{4}\mathrm{P}_{1/2}^{2} \\ 58{}^{4}\mathrm{P}_{3/2}6p{}^{4}\mathrm{P}_{3/2}^{2} \\ 58{}^{4}\mathrm{P}_{3/2}6p{}^{4}\mathrm{P}_{3/2}^{2} \end{array}$
4860. 04 4906. 49 4920. 98 4921. 80 4954. 73	250 10 300 $10w$ 300	20570. 22 20375. 48 20315. 48 20312. 10 20177. 10	$egin{array}{cccccccccccccccccccccccccccccccccccc$
4979. 76 4983. 25 5002. 72 5029. 37 5063. 74	4000 75 500 200 40	20075. 69 20061. 63 19983. 55 19877. 66 19742. 75	$\begin{array}{c} 5s' \ ^2\mathrm{P}_{3/2} -\!$
5092, 60 5138, 55 5139, 47 5148, 78 5150, 47	$egin{array}{c} 4 \\ 10 \\ 10w \\ 10 \\ 6h \end{array}$	19630. 86 19455. 32 19451. 84 19416. 67 19410. 30	$egin{array}{cccccccccccccccccccccccccccccccccccc$
5220. 71 5222. 32 5226. 91 5241. 43 5245. 12	$ \begin{array}{r} 8 \\ 12 \\ 20 \\ 5 \\ 350 \end{array} $	19149. 15 19143. 25 19126. 44 19073. 45 19060. 04	$\begin{array}{c} 5p\ ^4\mathrm{P}_{5/2}^{\circ}-10d\ ^4\mathrm{D}_{3/2} \\ 5p\ ^4\mathrm{P}_{5/2}^{\circ}-10d\ ^4\mathrm{D}_{5/2} \\ 5p\ ^4\mathrm{P}_{5/2}^{\circ}-10d\ ^4\mathrm{D}_{7/2} \\ 5p\ ^4\mathrm{P}_{5/2}^{\circ}-6d''\ ^2\mathrm{D}_{3/2} \\ 5s'\ ^4\mathrm{P}_{1/2}-6p\ ^4\mathrm{D}_{3/2}^{\circ} \end{array}$
5261. 80 5285. 23 5297. 32 5314. 18 5315. 85	5 50 15 10 8	18999. 62 18915. 39 18872. 22 18812. 34 18806. 44	$\begin{array}{c} 5p\ ^4\mathrm{P}_{5/2}^{2}-\!$
5317. 25 5318. 83 5323. 20 5328. 92 5329. 30	$ \begin{array}{c} 20 \\ 7 \\ 12 \\ 200 \\ 12 \end{array} $	18801. 48 18795. 90 18780. 47 18760. 31 18758. 97	$\begin{array}{c} 5p\ ^4\mathrm{P}_{5/2}^{8}-6d'\ ^4\mathrm{P}_{5/2} \\ 5p\ ^4\mathrm{D}_{1/2}^{9}-12d\ ^4\mathrm{F}_{9/2} \\ 5p\ ^4\mathrm{P}_{5/2}^{8}-6d'\ ^4\mathrm{P}_{3/2} \\ 5s'\ ^2\mathrm{P}_{3/2}-6p\ ^4\mathrm{D}_{3/2}^{8} \\ 5p\ ^4\mathrm{P}_{5/2}^{8}-9d\ ^4\mathrm{F}_{7/2} \end{array}$
5337. 10 5345. 42 5348. 29 5354. 67 5356. 74	$40 \\ 600 \\ 10 \\ 120 \\ 12w$	18731, 56 18702, 40 18692, 37 18670, 09 18662, 88	$\begin{array}{c} 5p\ ^4\mathrm{P}_{5/2} - 9d\ ^4\mathrm{D}_{7/2} \\ 5s'\ ^4\mathrm{P}_{1/2} - 6p\ ^4\mathrm{P}_{3/2}^* \\ 5p\ ^4\mathrm{P}_{3/2}^* - (^1\mathrm{D}_2)\ 4d\ ^2\mathrm{S}_{1/2} \\ 5p\ ^4\mathrm{P}_{5/2}^* - 6d'\ ^2\mathrm{F}_{5/2} \\ 5p\ ^4\mathrm{P}_{3/2}^* - (^1\mathrm{D}_2)\ 4d\ ^2\mathrm{F}_{5/2} \end{array}$
5364, 19 5370, 34 5379, 13 5382, 96 5384, 27	300 300 15 120 8	18636. 96 18615. 62 18585. 20 18571. 98 18567. 46	$egin{array}{cccccccccccccccccccccccccccccccccccc$

Table A2. Observed lines of Br I—Continued

Wavelength	Intensity	Wave number	Classification
\mathring{A}		K	
5393. 58 5395. 48 5395. 55 5414. 26 5420. 43	7 1200 900 4 13	18535. 41 18528. 88 18528. 64 18464. 61 18443. 59	$ \begin{cases} 5p \ ^4\mathrm{P}_{3/2}^{\circ} - 11s \ ^4\mathrm{P}_{5/2} \\ 5s'' \ ^2\mathrm{P}_{1/2} - (^1\mathrm{D}_2) 5p \ ^2\mathrm{P}_{1/2}^{\circ} \\ 5p \ ^4\mathrm{P}_{3/2}^{\circ} - 6d' \ ^4\mathrm{P}_{5/2} \\ 5p \ ^4\mathrm{P}_{3/2}^{\circ} - 6d' \ ^4\mathrm{P}_{3/2} \end{cases} $
5420. 80 5424. 61 5432. 47 5434. 22 5450. 09	$ \begin{array}{r} 18 \\ 8h \\ 60 \\ 7 \\ 550 \end{array} $	18442, 33 18429, 38 18402, 72 18396, 79 18343, 22	$\begin{array}{c} 5p\ ^4\mathrm{P}_{3/2}^{3}-9d\ ^4\mathrm{P}_{1/2} \\ 5p\ ^4\mathrm{D}_{5/2}^{5}-11d\ ^4\mathrm{F}_{7/2} \\ 5s'\ ^2\mathrm{P}_{3/2}-6p\ ^4\mathrm{P}_{3/2}^{5} \\ 5p\ ^4\mathrm{P}_{3/2}^{3}-9d\ ^4\mathrm{D}_{5/2} \\ 5s'\ ^2\mathrm{P}_{3/2}-6p\ ^4\mathrm{P}_{5/2}^{5} \end{array}$
5453. 03 5455. 16 5455. 96 5463. 72 5466. 22	90 60 70 80 1200	18333. 33 18326. 17 18323. 49 18297. 46 18289. 10	$\begin{array}{c} 5p\ ^4\mathrm{P}_{3/2}^{*}-6d'\ ^2\mathrm{F}_{5/2} \\ 5p\ ^4\mathrm{P}_{3/2}^{*}-9d\ ^4\mathrm{D}_{3/2} \\ 5p\ ^4\mathrm{P}_{5/2}^{*}-10s\ ^4\mathrm{P}_{5/2} \\ 5p\ ^4\mathrm{D}_{7/2}^{*}-10d\ ^4\mathrm{F}_{9/2} \\ 5s''\ ^2\mathrm{P}_{1/2}^{*}-(^1\mathrm{D}_2)5p\ ^2\mathrm{P}_{3/2}^{*} \end{array}$
5469. 76 5470. 90 5504. 46 5520. 86 5522. 53	$15 \\ 5 \\ 35h \\ 20 \\ 90$	18277. 26 18273. 45 18162. 04 18108. 09 18102. 61	$\begin{array}{c} 5p \ ^4\mathrm{D}_{7/2}^{ 2} - 10d \ ^4\mathrm{D}_{7/2} \\ 5p \ ^4\mathrm{P}_{3/2}^{ 2} - 9d \ ^4\mathrm{D}_{1/2} \\ 5p \ ^4\mathrm{D}_{5/2}^{ 2} - 10d \ ^4\mathrm{F}_{7/2} \\ 5p \ ^4\mathrm{P}_{3/2}^{ 2} - 6d' \ ^2\mathrm{P}_{3/2} \\ 5p \ ^4\mathrm{P}_{5/2}^{ 2} - 8d \ ^4\mathrm{D}_{3/2} \end{array}$
5529. 01 5532. 22 5536. 37 5544. 72 5546. 90	175 100 300 6 20	18081. 40 18070. 91 18057. 36 18030. 17 18023. 08	$\begin{array}{c} 5p\ ^4\mathrm{P}_{5/2}^{5}-8d\ ^4\mathrm{D}_{5/2} \\ 5p\ ^4\mathrm{P}_{5/2}^{5}-8d\ ^4\mathrm{F}_{5/2} \\ 5p\ ^4\mathrm{P}_{5/2}^{5}-8d\ ^4\mathrm{D}_{7/2} \\ 5p\ ^4\mathrm{P}_{3/2}^{5}-10s\ ^4\mathrm{P}_{3/2} \\ 5p\ ^4\mathrm{D}_{7/2}^{5}-11s\ ^4\mathrm{P}_{5/2} \end{array}$
5558. 13 5558. 40 5588. 17 5590. 56 5597. 26	40 10 160 50 15	17986. 67 17985. 79 17889. 98 17882. 33 17860. 93	$\begin{array}{c} 5p\ ^4\mathrm{P}_{3/2}^{}-10s\ ^4\mathrm{P}_{5/2}^{}\\ 5p\ ^4\mathrm{D}_{5/2}^{}-(^1\mathrm{D}_2)\ 4d\ ^2\mathrm{F}_{7/2}^{}\\ 5p\ ^4\mathrm{D}_{7/2}^{}-9d\ ^4\mathrm{F}_{9/2}^{}\\ 5p\ ^4\mathrm{D}_{7/2}^{}-9d\ ^4\mathrm{D}_{7/2}^{}\\ 5p\ ^4\mathrm{D}_{5/2}^{}-11s\ ^4\mathrm{P}_{3/2}^{}\\ \end{array}$
5602. 41 5604. 96 5609. 83 5611. 62 5627. 24	$ \begin{array}{r} 8h \\ 4 \\ 10h \\ 125 \\ 200 \end{array} $	17844, 51 17836, 39 17820, 91 17815, 22 17765, 77	$ \left\{ \begin{array}{c} 5p\ ^4\mathrm{P}_{3/2}^{}{-}8d\ ^4\mathrm{F}_{3/2} \\ 5p\ ^4\mathrm{P}_{5/2}^{}{-}(^1\mathrm{D}_2)4d\ ^2\mathrm{D}_{3/2} \\ 5p\ ^4\mathrm{P}_{1/2}^{}{-}(^1\mathrm{D}_2)4d\ ^2\mathrm{P}_{1/2} \\ 5p\ ^4\mathrm{D}_{7/2}^{}{-}6d'\ ^2\mathrm{F}_{5/2} \\ 5p\ ^4\mathrm{P}_{3/2}^{}{-}8d\ ^4\mathrm{P}_{1/2} \\ 5p\ ^4\mathrm{P}_{3/2}^{}{-}8d\ ^4\mathrm{D}_{3/2} \end{array} \right. $
5633. 97 5637. 27 5640. 86 5645. 31 5645. 97	140 130 35 30 45	17744. 55 17734. 16 17722. 87 17708. 90 17706. 83	$egin{array}{c} 5p\ ^4\mathrm{P}_{3/2}^{3}-8d\ ^4\mathrm{D}_{5/2}\ 5p\ ^4\mathrm{D}_{5/2}^{2}-9d\ ^4\mathrm{F}_{7/2}\ 5p\ ^4\mathrm{D}_{5/2}^{2}-6d'\ ^2\mathrm{F}_{7/2}\ 5p\ ^4\mathrm{D}_{5/2}^{3}-9d\ ^4\mathrm{D}_{5/2}\ 5p\ ^4\mathrm{D}_{5/2}^{5}-9d\ ^4\mathrm{D}_{7/2} \end{array}$
5667. 75 5667. 92 5685. 77 5697. 29 5705. 74	$9 \\ 8 \\ 15W \\ 18 \\ 85$	17638. 79 17638. 26 17582. 89 17547. 34 17521. 35	$\begin{array}{c} 5p\ ^4\mathrm{P}_{1/2}^{\circ}-6d'\ ^4\mathrm{P}_{3/2}^{\circ}\\ 5p\ ^4\mathrm{D}_{5/2}^{\circ}-9d\ ^4\mathrm{D}_{3/2}^{\circ}\\ (^1\mathrm{D}_2)5s\ ^2\mathrm{D}_{3/2}^{\circ}-7p''\ ^2\mathrm{P}_{1/2}^{\circ}\\ 5p\ ^4\mathrm{D}_{5/2}^{\circ}-6d'\ ^2\mathrm{F}_{7/2}\\ 5p\ ^4\mathrm{P}_{1/2}^{\circ}-9d\ ^4\mathrm{D}_{3/2}^{\circ}\end{array}$
5709. 44 5716. 26 5721. 11 5722. 97 5764. 66	$\begin{array}{c} 9\\225\\100\\40h\\45 \end{array}$	17509. 99 17489. 10 17474. 28 17468. 60 17342. 27	$\begin{array}{c} 5p\ ^4\mathrm{P}_{5/2}^{5}-9s\ ^4\mathrm{P}_{3/2}^{2} \\ 5p\ ^4\mathrm{P}_{5/2}^{6}-9s\ ^4\mathrm{P}_{5/2}^{6} \\ 5p\ ^4\mathrm{D}_{7/2}^{6}-10s\ ^4\mathrm{P}_{5/2}^{6} \\ 5p\ ^4\mathrm{P}_{1/2}^{6}-9d\ ^4\mathrm{D}_{1/2}^{1/2} \\ 5p\ ^4\mathrm{D}_{5/2}^{6}-10s\ ^4\mathrm{P}_{3/2}^{6} \end{array}$
5771. 30 5777. 69 5779. 97 5783. 32 5784. 02	5 10 4 500 30	17322, 31 17303, 16 17296, 33 17286, 31 17284, 22	$egin{array}{l} 5p\ ^4\mathrm{D}_{7/2}^{2}-8d\ ^4\mathrm{F}_{7/2}\ 5p\ ^4\mathrm{P}_{1/2}^{\circ}-6d'\ ^2\mathrm{P}_{3/2} \end{array} \ \ egin{array}{l} 5p\ ^4\mathrm{D}_{7/2}^{\circ}-8d\ ^4\mathrm{F}_{9/2}\ 5p\ ^4\mathrm{P}_{5/2}^{\circ}-7d\ ^4\mathrm{F}_{5/2} \end{array}$

Table A2. Observed lines of Br I—Continued

Wavelength	Intensity	Wave number	Classification
Å		K	
5793. 98 5801. 50 5803. 82 5805. 04 5809. 59	100 40 30 30 250	17254. 51 17232. 14 17225. 26 17221. 64 17208. 15	$\begin{array}{c} 5p\ ^4\mathrm{P}_{5/2}^{\circ}-7d\ ^4\mathrm{F}_{3/2} \ 5p\ ^4\mathrm{D}_{7/2}^{\circ}-8d\ ^4\mathrm{D}_{5/2} \ 5p\ ^4\mathrm{P}_{1/2}^{\circ}-10s\ ^4\mathrm{P}_{3/2} \ 5p\ ^4\mathrm{D}_{7/2}^{\circ}-8d\ ^4\mathrm{F}_{5/2} \ 5p\ ^4\mathrm{D}_{2/2}^{\circ}-8d\ ^4\mathrm{D}_{7/2} \end{array}$
5819. 56 5821. 45 5826. 07 5827. 08 5828. 51	$ \begin{array}{c} 80 \\ 12 \\ 5 \\ 10 \\ 150 \end{array} $	17178. 67 17173. 09 17159. 47 17156. 50 17152. 29	$egin{array}{l} 5p\ ^4 ext{P}_{1/2}^{\circ}-6d'\ ^2 ext{P}_{1/2} \ 5p\ ^4 ext{P}_{3/2}^{\circ}-9s\ ^4 ext{P}_{3/2} \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \$
5828. 89 5830. 39 5833. 39 5836. 84 5852. 08	7 400 900 150 1800	17151. 17 17146. 76 17137. 94 17127. 81 17083. 21	$5p\ ^4\mathrm{D}_{5/2}^{5}-8d\ ^4\mathrm{F}_{7/2} \ 5p\ ^4\mathrm{P}_{5/2}^{5}-7d\ ^4\mathrm{D}_{5/2} \ 5p\ ^4\mathrm{P}_{5/2}^{5}-7d\ ^4\mathrm{D}_{3/2} \ 5p\ ^4\mathrm{P}_{5/2}^{5}-7d\ ^4\mathrm{D}_{7/2}$
5861. 20 5864. 82 5867. 05 5869. 47 5877. 15	$ \begin{array}{c} 120 \\ 90 \\ 60 \\ 18 \\ 50w \end{array} $	17056. 62 17046. 10 17039. 62 17032. 59 17010. 34	$5p\ ^4D_{5/2}^2-8d\ ^4D_{5/2}$ $5p\ ^4D_{5/2}^3-8d\ ^4F_{5/2}$ $5p\ ^4P_{1/2}^2-8a\ ^4F_{3/2}$ $5p\ ^4P_{5/2}^3-8d\ ^4D_{7/2}$ $5p\ ^4P_{1/2}^2-8d\ ^4P_{1/2}$
5886. 87 5889. 87 5894. 28 5898. 32 5898. 51	$ \begin{cases} 60cw \\ 60 \\ 18 \\ 8 \\ 8d \end{cases} $	16982, 25 16973, 60 16960, 90 16949, 28 16948, 74	$\left. \begin{array}{c} ^{(1}\mathrm{D}_{2})5s~^{2}\mathrm{D}_{3/2} -\!$
5898. 96	90	16947. 44	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$
5905. 45 5905. 66 5909. 31 5937. 85	900 80 250 90	16928. 82 16928. 22 16917. 76 16836. 45	$ \left\{ \begin{array}{c} 5p \ ^4\mathrm{P}_{3/2}^* - (^1\mathrm{S}_0) 5s \ ^2\mathrm{S}_{1/2} \\ 5p \ ^4\mathrm{D}_{3/2}^* - (^1\mathrm{D}_2) 4d \ ^2\mathrm{F}_{5/2} \\ (^1\mathrm{D}_2) 5s \ ^2\mathrm{D}_{5/2} - 7p' \ ^2\mathrm{D}_{3/2}^* \\ 5p \ ^4\mathrm{P}_{3/2}^* - 7d \ ^4\mathrm{F}_{3/2} \\ 5p \ ^4\mathrm{P}_{1/2} - 8d \ ^4\mathrm{D}_{1/2} \end{array} \right. $
5940. 48 5943. 71 5945. 50 5950. 32 5953. 92	1600 140 600 750 750	16828. 99 16819. 85 16814. 78 16801. 16 16791. 01	$\begin{array}{c} 5s^{\prime\prime} ^2\mathrm{P}_{1/2} \!$
5982, 90 5983, 35 5985, 32 5992, 89 5999, 73	$ \begin{array}{c} 12 \\ 5 \\ 125 \\ 100 \\ 95 \end{array} $	16709. 67 16708. 42 16702. 92 16681. 82 16662. 80	$\begin{array}{c} 5p\ ^4\mathrm{D}_{3/2}^{8}-6d'\ ^4\mathrm{P}_{3/2}\\ 5p\ ^4\mathrm{D}_{3/2}^{8}-9a\ ^4\mathrm{P}_{1/2}\\ 5p\ ^4\mathrm{P}_{1/2}^{9}-(^1\mathrm{D}_2)4d\ ^2\mathrm{D}_{3/2}\\ 5p\ ^4\mathrm{P}_{3/2}^{8}-7d\ ^4\mathrm{D}_{1/2}\\ 5p\ ^4\mathrm{D}_{3/2}^{8}-9d\ ^4\mathrm{D}_{5/2} \end{array}$
6007. 96 6018. 17 6022. 66 6025. 24 6057. 87	$ \begin{array}{r} 200 \\ 200 \\ 30 \\ 10h \\ 200 \end{array} $	16639, 98 16611, 75 16599, 36 16592, 25 16502, 88	$\begin{array}{c} 5p\ ^4\mathrm{D}_{3/2}^{9}-9s\ ^4\mathrm{P}_{5/2} \ 5p\ ^4\mathrm{P}_{3/2}^{9}-7s'\ ^2\mathrm{P}_{3/2} \ 5p\ ^4\mathrm{D}_{3/2}^{9}-6d'\ ^2\mathrm{F}_{5/2} \ 5p\ ^4\mathrm{D}_{3/2}^{9}-9d\ ^4\mathrm{D}_{3/2} \ 5p\ ^4\mathrm{P}_{5/2}^{9}-5d''\ ^2\mathrm{D}_{5/2} \end{array}$
6064. 35 6067. 36 6095. 74 6107. 69 6111. 10	200 5 700 125 80	16485. 25 16477. 07 16400. 36 16368. 27 16359. 14	$\begin{array}{c} 5p\ ^{4}\mathrm{D}_{3/2}^{8}-9s\ ^{4}\mathrm{P}_{3/2}^{2}\\ 5p\ ^{4}\mathrm{D}_{7/2}^{2}-(^{1}\mathrm{D}_{2})4d\ ^{2}\mathrm{G}_{7/2}\\ 5p\ ^{4}\mathrm{D}_{7/2}^{2}-(^{1}\mathrm{D}_{2})4d\ ^{2}\mathrm{G}_{9/2}\\ 5p\ ^{4}\mathrm{P}_{1/2}^{2}-9s\ ^{4}\mathrm{P}_{3/2}\\ 5p\ ^{4}\mathrm{D}_{7/2}^{2}-7d\ ^{4}\mathrm{F}_{7/2} \end{array}$
6116. 19 6118. 80 6122. 14 6122. 35 6126. 53	$\begin{bmatrix} 60 \\ 300 \\ 2400 \\ 100 \\ 30W \end{bmatrix}$	16345, 52 16338, 55 16329, 64 16329, 08 16317, 94	$\begin{array}{c} 5p\ ^4\mathrm{D}_{3/2}^3-\!$

Table A2. Observed lines of Br I—Continued

Wavelength	Intensity	Wave number	Classification
Å		K	
6132. 71 6134. 71 6137. 49 6141. 04 6142. 73	$ \begin{array}{c} 800 \\ 75 \\ 150 \\ 100w \\ 100 \end{array} $	16301, 49 16296, 18 16288, 80 16279, 38 16274, 90	$\begin{array}{c} 5p\ ^4\mathrm{D}^{\circ}_{5/2} - (^{1}\mathrm{D}_{2})\ 4d\ ^2\mathrm{G}_{7/2} \\ 5p\ ^4\mathrm{D}^{\circ}_{3/2} - 10s\ ^4\mathrm{D}_{3/2} \\ 5p\ ^4\mathrm{D}^{\circ}_{7/2} - 7d\ ^4\mathrm{D}_{5/2} \\ (^{1}\mathrm{D}_{2})\ 5s\ ^2\mathrm{D}_{3/2} - 6f\ [3]^{\circ}_{3/2} \\ 5p\ ^4\mathrm{P}^{\circ}_{3/2} - 7s'\ ^2\mathrm{P}_{3/2} \end{array}$
6148, 60 d 6151, 10 6158, 19 6177, 39 6184, 09	$ \begin{array}{r} 40000 \\ 50d \\ 750 \\ 2000 \\ 200 \end{array} $	16259. 37 16252. 76 16234. 05 16183. 59 16166. 05	$\begin{array}{c} 5s\ ^4\mathrm{P}_{5/2} \!\!-\!\! 5p''\ ^2\mathrm{P}_{3/2}^2 \\ 5p\ ^4\mathrm{D}_{3/2}^2 \!\!-\!\! 10s\ ^4\mathrm{P}_{5/2} \\ 5p\ ^4\mathrm{D}_{7/2}^2 \!\!-\!\! 7d\ ^4\mathrm{D}_{7/2} \\ 5p\ ^4\mathrm{P}_{5/2}^2 \!\!-\!\! 7d\ ^4\mathrm{F}_{7/2} \\ 5p\ ^4\mathrm{P}_{3/2}^2 \!\!-\!\! 5d''\ ^2\mathrm{P}_{5/2} \end{array}$
6199. 74 6203. 08 6204. 35 6204. 49 6205. 40	50 900 90 90 90 120	16125. 25 16116. 56 16113. 27 16112. 90 16110. 54	$\begin{array}{c} 5p\ ^4\mathrm{P}^{5}_{5/2} -\!\!-\! 8s\ ^4\mathrm{P}_{3/2} \\ 5p\ ^4\mathrm{P}^{2}_{5/2} -\!\!-\! 8s\ ^4\mathrm{P}_{5/2} \\ 5p\ ^4\mathrm{P}^{5}_{5/2} -\!\!-\! 7d\ ^4\mathrm{D}_{5/2} \\ 5p\ ^4\mathrm{P}^{5}_{1/2} -\!\!-\! 7d\ ^4\mathrm{F}_{3/2} \\ 5p\ ^4\mathrm{D}^{3}_{3/2} -\!\!-\! 8d\ ^4\mathrm{F}_{3/2} \end{array}$
6208. 26 6216. 71 6225. 51 6235. 87 6244. 39	$ \begin{array}{r} 100 \\ 5 \\ 65 \\ 10 \\ 400 \end{array} $	16103. 12 16081. 23 16058. 50 16031. 82 16009. 95	$\begin{array}{c} 5p~^4\mathrm{D}_{5/2}^*-\!\!-\!7d~^4\mathrm{D}_{3/2} \ 5p~^4\mathrm{D}_{3/2}^*-\!\!-\!8d~^4\mathrm{P}_{1/2} \ 5p~^4\mathrm{D}_{5/2}^*-\!\!-\!7d~^4\mathrm{D}_{7/2} \ 5p~^4\mathrm{D}_{3/2}^*-\!\!-\!8d~^4\mathrm{D}_{3/2} \ 5p~^4\mathrm{P}_{1/2}^*-\!\!-\!7d~^4\mathrm{P}_{1/2} \end{array}$
6248. 24 6251. 32 6253. 69 6282. 46 6284. 69	200 300 400 600 8	16000. 08 15992. 20 15986. 14 15912. 93 15907. 28	$\begin{cases} 5p \ ^4\mathrm{D}_{3/2} -\!$
6290. 13 6296. 71 6301. 36 6315. 12 6331. 99	550 700 275 5 300	15893. 53 15876. 92 15865. 20 15830. 63 15788. 46	$egin{array}{l} 5p\ ^4 ext{P}_{5/2}^*-5d'\ ^2 ext{F}_{5/2}\ 5p\ ^4 ext{P}_{1/2}^*-7d\ ^4 ext{D}_{1/2}\ 5p\ ^4 ext{P}_{3/2}^*-5d'\ ^4 ext{P}_{3/2}\ 5p'\ ^2 ext{S}_{1/2}^*-8s'\ ^2 ext{P}_{3/2}\ 5p\ ^4 ext{P}_{3/2}^*-8s\ ^4 ext{P}_{3/2} \end{array}$
6335. 48 6336. 39 6336. 56 6337. 85 6343. 79	$ \begin{cases} 1500 \\ 300 \\ 500 \\ 60 \\ 35w \end{cases} $	15779. 76 15777. 49 15777. 07 15773. 86 15759. 09	$\left.\begin{array}{c} 5p\ ^{4}\mathrm{P}_{3/2}^{\circ}-8s\ ^{4}\mathrm{P}_{5/2}\\ (^{1}\mathrm{D}_{2})5s\ ^{2}\mathrm{D}_{5/2}-8p\ ^{4}\mathrm{D}_{3/2}^{\circ}\\ 5p\ ^{4}\mathrm{D}_{3/2}^{\circ}-(^{1}\mathrm{D}_{2})4d\ ^{2}\mathrm{D}_{3/2}\\ (^{1}\mathrm{D}_{2})5s\ ^{2}\mathrm{D}_{3/2}-8p\ ^{4}\mathrm{D}_{3/2}^{\circ}\end{array}\right.$
6345. 30 6349. 82 6350. 73 6371. 60 6392. 57	$\begin{array}{c} 500 \\ 500 \\ 60000 \\ 200 W \\ 35 W \end{array}$	15755. 34 15744. 12 15741. 87 15690. 31 15638. 84	$\begin{array}{c} 5p\ ^4\mathrm{P}_{5/2}^{\circ}-5d'\ ^2\mathrm{F}_{7/2} \\ 5p\ ^4\mathrm{P}_{5/2}^{\circ}-6d\ ^4\mathrm{F}_{3/2} \\ 5s\ ^4\mathrm{P}_{5/2}-5p'\ ^4\mathrm{S}_{3/2}^{\circ} \\ (^1\mathrm{D}_2)\ 5s\ ^2\mathrm{D}_{3/2}-8p\ ^4\mathrm{P}_{1/2}^{\circ} \\ (^1\mathrm{D}_2)\ 5s\ ^2\mathrm{D}_{5/2}-8p\ ^4\mathrm{P}_{3/2}^{\circ} \end{array}$
6398. 03 6399. 99 6405. 52 6405. 58 6405. 68	$ \begin{array}{c} 30 \\ 10w \\ 3 \\ 4 \\ 6 \end{array} $	15625. 49 15620. 71 15607. 22 15607. 07 15606. 83	$ \left. \begin{array}{c} 5p' ^2\mathrm{S}_{1/2}^\circ - (^1\mathrm{D}_2) 4d ^2\mathrm{S}_{1/2} \\ (^1\mathrm{D}_2) 5s ^2\mathrm{D}_{3/2} - 8p ^4\mathrm{P}_{3/2}^\circ \\ \\ \left. \begin{array}{c} (^1\mathrm{D}_2) 5s ^2\mathrm{D}_{5/2} - 4f' [4]_{7/2}^\circ \end{array} \right. \end{array} \right. $
6405. 80 6410. 32 6418. 29 6418. 96 6419. 03	$ \begin{array}{c} 8 \\ 2500 \\ 45 \\ 15s \\ 15 \end{array} $	15606. 54 15595. 53 15576. 17 15574. 54 15574. 37	$\begin{array}{c} \begin{array}{c} \\ 5p~^{4}\mathrm{P}_{5/2}^{\circ}-6d~^{4}\mathrm{F}_{7/2} \\ 5p~^{4}\mathrm{P}_{3/2}^{\circ}-5d'~^{4}\mathrm{P}_{5/2} \\ \end{array} \\ \left. \begin{array}{c} (^{1}\mathrm{D}_{2})5s~^{2}\mathrm{D}_{3/2}-4f'~[2]_{5/2}^{\circ} \end{array} \right. \end{array}$
6426. 30 6435. 81 6436. 60 6438. 02 6458. 92	$ \begin{array}{c c} 500 \\ 15 \\ 10 \\ 600 \\ 25h \end{array} $	15556. 75 15533. 77 15531. 86 15528. 43 15478. 19	$egin{array}{l} 5p\ ^4 ext{P}_{3/2}^*-5d'\ ^2 ext{F}_{5/2}\ 5p\ ^4 ext{P}_{1/2}^*-7s'\ ^4 ext{P}_{1/2}\ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ $

Table A2. Observed lines of Br 1—Continued

	TABLE A2. Observed times of Br 1—Continued				
Wavelength	Intensity	Wave number	Classification		
\mathring{A}		K			
6462, 32 6470, 41 6475, 23 6483, 56 6483, 96	500 200 $100h$ 1800 35	15470. 04 15450. 70 15439. 20 15419. 36 15418. 41	$egin{array}{l} 5p\ ^4\mathrm{P}_{1/2}^{\circ}-7s'\ ^2\mathrm{P}_{3/2}\ 5p\ ^4\mathrm{P}_{3/2}^{\circ}-5d'\ ^2\mathrm{P}_{3/2}\ 5p\ ^4\mathrm{D}_{3/2}^{\circ}-9s\ ^4\mathrm{P}_{3/2}\ 5p\ ^4\mathrm{P}_{3/2}^{\circ}-6d\ ^4\mathrm{P}_{1/2}\ 5p\ ^4\mathrm{D}_{3/2}^{\circ}-9s\ ^4\mathrm{P}_{5/2} \end{array}$		
6488. 62 6493. 80 6501. 51 6514. 32 6514. 62	$ \begin{array}{r} 800 \\ 20h \\ 4 \\ 12 \\ 1000 \end{array} $	15407. 34 15395. 05 15376. 79 15346. 56 15345. 85	$egin{array}{cccccccccccccccccccccccccccccccccccc$		
b 6531, 39 6532, 29 6541, 30 6544, 57 6548, 09	70 600 600 20000 1500	15306. 45 15304. 34 15283. 26 15275. 62 15267. 41	$egin{array}{c} 5p'\ ^2\mathrm{D}_{3/2}^*-8s'\ ^4\mathrm{P}_{1/2}\ 5p\ ^4\mathrm{D}_{5/2}^*-5d'\ ^2\mathrm{D}_{3/2}\ 5p\ ^4\mathrm{P}_{3/2}^*-5d'\ ^2\mathrm{P}_{1/2}\ 5p\ ^4\mathrm{P}_{5/2}^*-6d\ ^4\mathrm{D}_{5/2}\ 5p\ ^4\mathrm{D}_{7/2}^*-8s\ ^4\mathrm{P}_{5/2} \end{array}$		
6551. 57 6559. 80 6571. 31 6574. 29 6576. 24	$ \begin{array}{c} 12 \\ 50000cw \\ 1000 \\ 15 \\ 10W \end{array} $	15259. 30 15240. 16 15213. 46 15206. 57 15202. 06	$rac{5p'\ ^2 ext{S}_{1/2}^{\circ}-9d\ ^4 ext{D}_{3/2}}{5s\ ^4 ext{P}_{5/2}-5p'\ ^2 ext{D}_{3/2}^{\circ}} \ rac{5p\ ^4 ext{D}_{3/2}^{\circ}-7d\ ^4 ext{F}_{5/2}}{5p'\ ^2 ext{S}_{1/2}^{\circ}-9d\ ^4 ext{D}_{1/2}}$		
6579. 14 6579. 36 6582. 17 6582. 62 6584. 14	1800 300 20000 35 600	15195, 36 15194, 85 15188, 36 15187, 32 15183, 82	$\begin{array}{c} 5p\ ^4\mathrm{P}_{5/2}^{\rm s}-\!\!-\!\!6d\ ^4\mathrm{D}_{3/2} \\ 5p\ ^4\mathrm{D}_{3/2}^{\rm s}-\!\!-\!\!(^1\mathrm{S}_0).5s\ ^2\mathrm{S}_{1/2} \\ 5p\ ^4\mathrm{P}_{5/2}^{\rm s}-\!\!6d\ ^4\mathrm{D}_{7/2} \\ 5p\ ^4\mathrm{P}_{1/2}^{\rm s}-\!\!5d^{\prime\prime}\ ^2\mathrm{D}_{3/2} \\ 5p\ ^4\mathrm{D}_{3/2}^{\rm s}-\!\!7d\ ^4\mathrm{F}_{3/2} \end{array}$		
6589. 62 6604. 80 6613. 05 6620. 47 6621. 44	$ \begin{array}{r} 10w \\ 15 \\ 4 \\ 1500 \\ 8 \end{array} $	15171. 19 15136. 32 15117. 44 15100. 50 15098. 29	$egin{array}{c} 5p'\ ^2\mathrm{D}^5_{5/2} -\!-\!-\!(^1\mathrm{D}_2)4d\ ^2\mathrm{F}_{7/2} \ 4d\ ^4\mathrm{D}_{7/2} -\!-\!10f\ [4]^5_{9/2} \ 5p'\ ^4\mathrm{D}^5_{1/2} -\!8s'\ ^4\mathrm{P}_{1/2} \ 5p\ ^4\mathrm{D}^5_{5/2} -\!8s\ ^4\mathrm{P}_{3/2} \ 5p'\ ^2\mathrm{D}^5_{5/2} -\!-\!(^1\mathrm{D}_2)4d\ ^2\mathrm{P}_{3/2} \end{array}$		
6624, 26 6629, 09 6631, 62 6635, 10 6636, 62	$ \begin{array}{c} 20 \\ 20 \\ 50000cw \\ 60 \\ 150 \end{array} $	15091, 86 15080, 86 15075, 11 15067, 20 15063, 75	$\begin{array}{c} 5p\ ^4\mathrm{D}_{5/2}^{8}-8s\ ^4\mathrm{P}_{5/2} \\ 5p\ ^4\mathrm{D}_{3/2}^{8}-7d\ ^4\mathrm{P}_{1/2} \\ 5s\ ^4\mathrm{P}_{5/2}-5p'\ ^2\mathrm{D}_{5/2}^{s} \\ 5p\ ^4\mathrm{D}_{3/2}^{8}-7d\ ^4\mathrm{P}_{5/2} \\ 5p\ ^4\mathrm{D}_{7/2}^{8}-5d'\ ^4\mathrm{P}_{5/2} \end{array}$		
6639. 57 6645. 17 6646. 59 6653. 82 6666. 93	$ \begin{array}{r} 40 \\ 40 \\ 120 \\ 15 \\ 150 \end{array} $	15057. 06 15044. 37 15041. 16 15024. 81 14995. 27	$egin{array}{l} 5p\ ^4\mathrm{D}_{3/2}^3-7d\ ^4\mathrm{D}_{3/2}\ 5p\ ^4\mathrm{D}_{7/2}^2-5d'\ ^2\mathrm{F}_{5/2}\ 5p'\ ^2\mathrm{S}_{1/2}^2-6d'\ ^2\mathrm{P}_{3/2}\ -5p'\ ^2\mathrm{D}_{3/2}^3-(^1\mathrm{D}_2) 4d\ ^2\mathrm{S}_{1/2}\ 5p'\ ^2\mathrm{D}_{3/2}^3-(^1\mathrm{D}_2) 4d\ ^2\mathrm{F}_{5/2} \end{array}$		
6672. 15	600	14983. 54	5p 4P _{1/2} —8s 4P _{3/2}		
6676. 54	15	14973. 68	$\left\{egin{array}{ccccc} 4d\ ^4{ m D}_{7/2}5f'\ [3]_{5/2}^{st}\ 5p'\ ^2{ m D}_{3/2}^{st}({}^{ m i}{ m D}_2)4d\ ^2{ m P}_{1/2} \end{array} ight.$		
6676. 72 6681. 71 6682. 28	$ \begin{array}{c} 8 \\ 15 \\ 20000 \end{array} $	14973. 28 14962. 10 14960. 82	$egin{array}{cccccccccccccccccccccccccccccccccccc$		
6684, 17 6684, 23 6684, 33 6687, 33 6688, 08	50 50 70 90 90	14956, 59 14956, 46 14956, 23 14949, 52 14947, 85	$ \begin{cases} & (^{1}\mathrm{D}_{2})5s\ ^{2}\mathrm{D}_{5/2}5f\ [4]_{5/2}^{\circ} \\ & 4d\ ^{4}\mathrm{D}_{7/2}5f'\ [4]_{9/2}^{\circ} \\ & 5p\ ^{4}\mathrm{D}_{3/2}^{\circ}-7d\ ^{4}\mathrm{D}_{1/2} \end{cases} $		
6690. 35 6692. 13 6694. 62 6700. 71 6702. 07	$ \begin{array}{c} 120cw \\ 10000 \\ 10 \\ 60h \\ 110 \end{array} $	14942. 78 14938. 80 14933. 24 14919. 67 14916. 65	$\begin{array}{c} (^{1}\mathrm{D}_{2})5s^{2}\mathrm{D}_{3/2} - 5f[3]_{7/2}^{5} \\ 5p^{4}\mathrm{P}_{3/2}^{5} - 6d^{4}\mathrm{D}_{5/2}^{5} \\ 5p'^{2}\mathrm{D}_{3/2}^{5} - (^{1}\mathrm{D}_{2})4d^{2}\mathrm{P}_{3/2} \\ 5p'^{2}\mathrm{D}_{5/2}^{5} - 9d^{4}\mathrm{F}_{7/2} \\ 5p'^{2}\mathrm{S}_{1/2}^{5} - 6d'^{2}\mathrm{P}_{1/2} \end{array}$		

Table A2. Observed lines of Br I—Continued

Wavelength	Intensity	Wave number	Classification
Å		K	
6706, 79 6712, 12 6713, 06 6714, 86 6719, 97	60 30 125 90 8	14906. 15 14894. 31 14892. 23 14888. 23 14876. 91	$egin{array}{cccccccccccccccccccccccccccccccccccc$
6720. 68 6723. 65 6728. 28 6738. 61 6739. 66	75h 400 8000 100 8	14875. 34 14868. 77 14858. 54 14835. 76 14833. 45	$egin{array}{l} 4d\ ^4\mathrm{D}_{7/2}\!\!-\!9f\ [4]_{3/2}^3 \ 5p\ ^4\mathrm{D}_{5/2}^3\!\!-\!5d'\ ^2\mathrm{F}_{5/2} \ 5p\ ^4\mathrm{P}_{3/2}^3\!\!-\!6d\ ^4\mathrm{D}_{3/2} \ 5p'\ ^4\mathrm{D}_{1/2}^3\!\!-\!(^1\mathrm{D}_2)\ 4d\ ^2\mathrm{S}_{1/2} \ 5p\ ^4\mathrm{D}_{7/2}^2\!\!-\!6d\ ^4\mathrm{F}_{5/2} \ \end{array}$
6740. 83 6752. 67 6760. 06 6761. 92 6765. 12	5662000 258	14830. 87 14804. 87 14788. 69 14784. 62 14777. 62	$egin{array}{cccccccccccccccccccccccccccccccccccc$
6771. 95 6774. 63 6778. 57 6779. 48 6785. 23	$175 \\ 30h \\ 60 \\ 2000 \\ 10$	14762. 72 14756. 88 14748. 30 14746. 32 14733. 83	$egin{array}{cccccccccccccccccccccccccccccccccccc$
6785. 74 6786. 74 6787. 34 6787. 77 6790. 04	$900 \\ 2200 \\ 175 \\ 8 \\ 6500$	14732. 72 14730. 55 14729. 25 14728. 31 14723. 39	$egin{array}{cccccccccccccccccccccccccccccccccccc$
6791. 48 6801. 35 6816. 72 6820. 04 6820. 39	$1600c \\ 60 \\ 50 \\ 6h \\ 800$	14720. 27 14698. 91 14665. 76 14658. 63 14657. 87	$egin{array}{cccccccccccccccccccccccccccccccccccc$
6826. 02 6840. 62 6844. 82 6845. 27 6846. 27	400 $ 175 $ $ 5 $ $ 40c $ $ 150$	14645. 78 14614. 53 14605. 56 14604. 60 14602. 46	$egin{array}{cccccccccccccccccccccccccccccccccccc$
6853. 51 6858. 22 6859. 43 6861. 15 6875. 22	$\begin{array}{c} 3\\ 45\\ 20\\ 1800\\ 20w \end{array}$	14587. 04 14577. 02 14574. 45 14570. 80 14540. 98	$egin{array}{cccccccccccccccccccccccccccccccccccc$
6887, 99 6888, 73 6904, 95 6922, 86 6923, 04	$ \begin{array}{r} 12h \\ 80h \\ 400 \\ 30 \\ 1 \end{array} $	14514. 02 14512. 46 14478. 37 14440. 91 14440. 54	$\begin{array}{c} 4d\ ^4\mathrm{D}_{7/2} -\!\!-\! 8f\ [3]_{7/2}^2,\ _{5/2}^2 \\ 4d\ ^4\mathrm{D}_{7/2} -\!\!-\! 8f\ [4]_{8/2}^6 \\ 5p\ ^4\mathrm{P}_{1/2}^* -\!\!-\! 5d'\ ^2\mathrm{P}_{1/2}^2 \\ 5p'\ ^2\mathrm{S}_{1/2}^3 -\!\!-\! (^1\mathrm{D}_2)4d\ ^2\mathrm{D}_{3/2}^2 \\ 5p'\ ^2\mathrm{D}_{3/2}^3 -\!\!-\! 6d'\ ^2\mathrm{P}_{3/2}^2 \end{array}$
6927. 10 6927. 34 6929. 78 6936. 76 6962. 99	$ \begin{array}{c} 2 \\ 8 \\ 400 \\ 20 \\ 10 \end{array} $	14432. 07 14431. 57 14426. 49 14411. 98 14357. 69	$egin{array}{cccccccccccccccccccccccccccccccccccc$
6971. 96 6992. 25 6992. 83 6993. 31 7005. 19	$ \begin{array}{r} 250 \\ 5 \\ 5 \\ 20 \\ 10000 \end{array} $	14339. 21 14297. 60 14296. 42 14295. 44 14271. 19	$egin{array}{l} 5p\ ^4\mathrm{D}_{7/2}^\circ-6d\ ^4\mathrm{D}_{7/2} \ 4d\ ^4\mathrm{D}_{5/2}-8f\ [3]_{7/2}^\circ,\ _{5/2} \ 4d\ ^4\mathrm{D}_{5/2}-8f\ [4]_{7/2} \ 5p'\ ^4\mathrm{S}_{3/2}^\circ-6d'\ ^4\mathrm{P}_{5/2} \ 5s\ ^4\mathrm{P}_{3/2}-5p'\ ^4\mathrm{S}_{3/2}^\circ. \end{array}$

Table A2. Observed lines of Br I—Continued

	TABLE A2. (or i—Continued
Wavelength	Intensity	Wave number	Classification
\mathring{A}		K	
7011. 53 7015. 15 7024. 70 7026. 53 7031. 36	50 75 20 2 15	14258. 29 14250. 93 14231. 56 14227. 85 14218. 08	$egin{array}{cccccccccccccccccccccccccccccccccccc$
7054. 88 7058. 38 7061. 71 7066. 33 7076. 71	50 200 1 4 5	14170. 68 14163. 65 14156. 97 14147. 72 14126. 97	$\begin{array}{c} 5p\ ^4\mathrm{D}^8_{5/2} ext{}6d\ ^4\mathrm{D}_{3/2} \ 5p\ ^4\mathrm{D}^8_{5/2} ext{}6d\ ^4\mathrm{D}_{7/2} \ 5p'\ ^4\mathrm{S}^8_{3/2} ext{}9d\ ^4\mathrm{D}_{3/2} \ 5p'\ ^2\mathrm{D}^8_{3/2} ext{}8d\ ^4\mathrm{P}_{1/2} \ 5p'\ ^4\mathrm{D}^8_{1/2} ext{}6d'\ ^2\mathrm{P}_{1/2} \end{array}$
7082. 63 7091. 12 7101. 80 7107. 14 7111. 50	$\begin{bmatrix} 3\\4\\5h\\1\\100 \end{bmatrix}$	14115, 16 14098, 26 14077, 06 14066, 48 14057, 86	$\begin{array}{c} 5p''\ ^{2}\mathrm{P}_{1/2}^{\circ}-8s'\ ^{4}\mathrm{P}_{1/2} \\ 5p'\ ^{2}\mathrm{D}_{3/2}^{\circ}-8d\ ^{4}\mathrm{D}_{3/2} \\ 5p'\ ^{2}\mathrm{D}_{3/2}^{\circ}-8d\ ^{4}\mathrm{D}_{5/2} \\ 5p'\ ^{2}\mathrm{D}_{3/2}^{\circ}-8d\ ^{4}\mathrm{F}_{5/2} \\ \end{array}$
7111. 62 7111. 73 7113. 22 7113. 60 7117. 59	100 150 50 100 15	14057. 62 14057. 40 14054. 46 14053. 71 14045. 83	$ \begin{cases} (^{1}\mathrm{D}_{2})5s\ ^{2}\mathrm{D}_{5/2} -\!$
7120, 78 7120, 88 7133, 95 7138, 05 7138, 19		14039, 54 14039, 34 14013, 62 14005, 57 14005, 29	$ \begin{cases} & (^{1}\mathrm{D}_{2})5s\ ^{2}\mathrm{D}_{3/2} -\!$
7142. 17 7142. 30 7145. 56 7147. 07 7149. 08	300 400 2 4 15	13997. 49 13997. 23 13990. 85 13987. 89 13983. 96	$ \begin{cases} & (^{1}\mathrm{D}_{2})5s \ ^{2}\mathrm{D}_{3/2} -\!$
7150. 00 7150. 30 7153. 13 7160. 74 7162. 10	$ \begin{array}{c} 8 \\ 75 \\ 5h \\ 15 \\ 750 \end{array} $	13982, 16 13981, 57 13976, 04 13961, 19 13958, 54	$egin{array}{c} 4d\ ^4\mathrm{D}_{7/2}7f\ [4]_{5/2}^2\ 4d\ ^4\mathrm{D}_{7/2}7f\ [4]_{5/2}^2\ 5p''\ ^2\mathrm{P}_{3/2}^3(^1\mathrm{D}_2)\ 4d\ ^2\mathrm{F}_{5/2}\ 4d\ ^4\mathrm{D}_{1/2}5f'\ [2]_{3/2}^3\ 5s\ ^4\mathrm{P}_{3/2}5p'\ ^4\mathrm{D}_{1/2}^5 \end{array}$
7172. 22 7175. 74 7177. 89 7184. 30 7194. 24	$ \begin{array}{c} 25 \\ 5H \\ 5h \\ 300 \\ 3 \end{array} $	13938. 84 13932. 00 13927. 83 13915. 41 13896. 18	$egin{array}{cccccccccccccccccccccccccccccccccccc$
7194, 30 7194, 41 7194, 53 7214, 95 7217, 78	$ \begin{cases} 5 \\ 7 \\ 15 \\ 5c \\ 20 \end{cases} $	13896. 06 13895. 85 13895. 62 13856. 29 13850. 86	$ \begin{cases} & (^{1}\mathrm{D}_{2})5s\ ^{2}\mathrm{D}_{5/2}7p\ ^{4}\mathrm{D}_{3/2}^{\circ} \\ \\ & (^{1}\mathrm{D}_{2})5s\ ^{2}\mathrm{D}_{3/2}7p\ ^{4}\mathrm{P}_{1/2}^{\circ} \\ & 5p'\ ^{2}\mathrm{S}_{1/2}^{\circ}7d\ ^{4}\mathrm{F}_{3/2} \end{cases} $
7222. 31 7232. 45 7236. 86 7247. 29 7255. 56	$50 \\ 100 \\ 7 \\ 5h \\ 15cw$	13842. 17 13822. 76 13814. 34 13794. 46 13778. 74	$\begin{array}{c} 5p \ ^4\mathrm{D}_{3/2} - 5d' \ ^4\mathrm{P}_{5/2} \\ 5p \ ^4\mathrm{D}_{3/2} - 5d' \ ^2\mathrm{F}_{5/2} \\ 5p' \ ^4\mathrm{S}_{3/2} - 6d' \ ^2\mathrm{P}_{1/2} \\ 5p \ ^4\mathrm{D}_{3/2}^3 - 5d' \ ^4\mathrm{P}_{3/2} \\ (^1\mathrm{D}_2) 5s \ ^2\mathrm{D}_{5/2} - 7p \ ^4\mathrm{D}_{5/2}^5 \end{array}$
7257. 82 7260. 45 7261. 46 7262. 40 7265. 16	$ \begin{array}{r} 10h \\ 2000 \\ 20 \\ 25 \\ 5 \end{array} $	13774, 45 13769, 46 13767, 54 13765, 76 13760, 53	$\begin{array}{c} 4d\ ^4\mathrm{D}_{5/2}7f\ [2]_{5/2}^{8/2} \\ 5s\ ^4\mathrm{P}_{3/2}5p'\ ^2\mathrm{D}_{3/2}^{8/2} \\ 4d\ ^4\mathrm{D}_{5/2}7f\ [3]_{7/2}^{8/2},\ 5/2 \\ 4d\ ^4\mathrm{D}_{5/2}7f\ [4]_{7/2}^{8/2} \\ (^1\mathrm{D}_2)5s\ ^2\mathrm{D}_{3/2}^{8/2}7p\ ^4\mathrm{D}_{5/2}^{8/2} \end{array}$

Table A2. Observed lines of Br I—Continued

Wavelength	Intensity	Wave number	Classification
Ă		K	
7272. 58 7284. 41 7288. 40 7291, 92 7305. 03	$ \begin{array}{r} 3w \\ 5 \\ 150 \\ \hline 1 \\ 75 \end{array} $	13746. 49 13724. 17 13716. 65 13710. 03 13685. 43	$\substack{(^{1}\mathrm{D}_{2})5s\ ^{2}\mathrm{D}_{5/2}-7p\ ^{4}\mathrm{D}_{7/2}^{2}\\5p'\ ^{2}\mathrm{S}_{1/2}^{2}-7d\ ^{4}\mathrm{D}_{3/2}\\5p\ ^{4}\mathrm{D}_{3/2}^{2}-5d'\ ^{2}\mathrm{P}_{3/2}\\5p''\ ^{2}\mathrm{P}_{3/2}^{3}-9d\ ^{4}\mathrm{D}_{5/2}\\5p\ ^{4}\mathrm{D}_{3/2}^{3}-6d\ ^{4}\mathrm{P}_{1/2}}$
7310. 45 7311. 48 7319. 31 7319. 41 7319. 53	$ \left. \begin{array}{c} 4 \\ 100d \\ 20 \\ 25 \\ 30 \end{array} \right. $	13675. 28 13673. 35 13658. 73 13658. 54 13658. 32	$ \left. \begin{array}{c} 5p' {}^{4}\mathrm{S}_{3/2}^{*} - 8d {}^{4}\mathrm{F}_{3/2} \\ 5p {}^{4}\mathrm{D}_{3/2}^{*} - 6d {}^{4}\mathrm{F}_{3/2} \\ \end{array} \right\} \\ ({}^{1}\mathrm{D}_{2}) 5s {}^{2}\mathrm{D}_{5/2} - 7p {}^{4}\mathrm{P}_{5/2}^{*} \end{array} \right\} $
7323, 35 7329, 20 7329, 68 7333, 65 7333, 77	$ \begin{array}{c} 10 \\ 10hw \\ 1 \\ 20s \\ 25 \end{array} $	13651. 19 13640. 30 13639. 40 13632. 02 13631. 80	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$
7342, 85 7344, 53 7348, 51 7378, 43 7383, 69	$ \begin{array}{r} 10 \\ 100 \\ 10000 \\ 20d \\ 5 \end{array} $	13614. 94 13611. 83 13604. 45 13549. 29 13539. 63	$egin{array}{cccccccccccccccccccccccccccccccccccc$
7385, 54 7420, 69 7425, 85 7452, 13 7453, 43	$ \begin{array}{c} 4h \\ 2 \\ 750 \\ 60 \\ 70 \end{array} $	13536. 24 13472. 13 13462. 76 13415. 29 13412. 95	$4d\ ^4\mathrm{D}_{1/2}$ —8 $f\ [2]_{3/2}$ $5p'\ ^4\mathrm{S}_{3/2}$ —8 $d\ ^4\mathrm{D}_{1/2}$ $5p\ ^4\mathrm{P}_{5/2}$ —78 $^4\mathrm{P}_{5/2}$ $5p'\ ^2\mathrm{D}_{3/2}$ —7 $d\ ^4\mathrm{F}_{3/2}$ $5p'\ ^2\mathrm{D}_{3/2}$ —7 $s''\ ^2\mathrm{P}_{3/2}$
7458. 23 7461. 86 7464. 68 7495. 46 7497. 81	$50h \\ 70 \\ 8 \\ 15hw \\ 5$	13404. 32 13397. 79 13392. 73 13337. 74 13333. 56	$\begin{array}{c} 4d\ ^4\mathrm{D}_{3/2}7f\ [2]_{3/2}^3\\ 4d\ ^4\mathrm{D}_{3/2}7f\ [3]_{5/2}^5\\ 5p''\ ^2\mathrm{P}_{3/2}^3(^1\mathrm{D}_2)4d\ ^2\mathrm{D}_{5/2}\\ 4d\ ^4\mathrm{F}_{9/2}9f\ [5]_{1}_{1/2}^1\\ 4d\ ^4\mathrm{D}_{5/2}9p\ ^4\mathrm{P}_{3/2}^3\end{array}$
7512. 96 7532. 77 7535. 79 7545. 60 7551. 40	40000 50 400 1 100	13306. 67 13271. 67 13266. 36 13249. 11 13238. 93	$\begin{array}{c} 5s{}^{4}\mathrm{P}_{5/2} -\!$
7551, 48 7551, 60 7559, 99 7569, 08 7570, 87	125 175 70 500 550	13238. 79 13238. 58 13223. 89 13208. 01 13204. 89	$ \begin{cases} (^{1}\mathrm{D}_{2})5s\ ^{2}\mathrm{D}_{3/2}6p'\ ^{4}\mathrm{D}_{1/2}^{\circ} \\ 5p'\ ^{4}\mathrm{D}_{1/2}^{\circ}7s''\ ^{2}\mathrm{P}_{34} \\ 5p'\ ^{2}\mathrm{S}_{1/2}^{\circ}7s'\ ^{2}\mathrm{P}_{3/2} \\ 5p\ ^{4}\mathrm{D}_{3/2}^{\circ}6d\ ^{4}\mathrm{D}_{5/2} \end{cases} $
7575. 23 7582. 56 7582. 98 7586. 60 7591. 61	15w $10h$ $100h$ 25 1600	13197. 29 13184. 53 13183. 80 13177. 51 13168. 81	$rac{4d\ ^4\mathrm{D}_{7/2}4f^{\prime\prime}\ [3]_{^7/2}}{4d\ ^4\mathrm{D}_{7/2}6f\ [2]_{^5/2}^\circ} \ _{5s\ ^4\mathrm{P}_{3/2}5p^\prime\ ^2\mathrm{S}_{1/2}^\circ}$
7594. 52 7595. 07 7598. 00 7604. 03 7606. 28	$ \begin{array}{c} 25 \\ 1800 \\ 4 \\ 70 \\ 120 \end{array} $	13163. 77 13162. 81 13157. 74 13147. 30 13143. 41	$\begin{array}{c} 4d\ ^{4}\mathrm{D}_{7/2}6f\ [4]_{7/2}^{2}\\ 4d\ ^{4}\mathrm{D}_{7/2}6f\ [4]_{8/2}^{6/2}\\ 5p''\ ^{2}\mathrm{P}_{3/2}^{3}8d\ ^{4}\mathrm{F}_{3/2}\\ 5p'\ ^{2}\mathrm{D}_{3/2}^{3}7d\ ^{4}\mathrm{P}_{1/2}\\ 4d\ ^{4}\mathrm{D}_{7/2}6f\ [3]_{7/2}^{3/2},\ _{5/2}^{6/2} \end{array}$
7607. 15 7607. 25 7607. 38 7607. 54 7616. 41	$ \left\{ \begin{array}{c} 100 \\ 150 \\ 200 \\ 300 \\ 2000 \end{array} \right. $	13141. 91 13141. 74 13141. 51 13141. 24 13125. 93	$ \left. \begin{array}{c} (^{1}{\rm D}_{2})5s\ ^{2}{\rm D}_{5/2}6p'\ ^{4}{\rm S}_{3/2}^{*} \\ \\ 5p\ ^{4}{\rm P}_{3/2}^{*}7s\ ^{4}{\rm P}_{5/2} \end{array} \right. $

Table A2. Observed lines of Br I—Continued

Wavelength	Intensity	Wave number	Classification
Å		K	
7617, 82 7617, 90 7618, 03 7632, 87 7641, 64	$ \begin{cases} 3 \\ 5 \\ 10 \\ 50 \\ 600 \end{cases} $	13123. 50 13123. 36 13123. 14 13097. 63 13082. 59	$ \left. \begin{array}{c} \left\{ \begin{array}{cccccccccccccccccccccccccccccccccccc$
7647. 70 7652. 60 7652. 70 7652. 85 7653. 02	$ \left. \begin{array}{c} 40 \\ 30 \\ 40 \\ 60 \\ 120 \end{array} \right. $	13072. 23 13063. 86 13063. 69 13063. 43 13063. 14	$\left.\begin{array}{c} 5p'\ ^4\mathrm{D}_{1/2}^{\circ}-\!\!-\!\!(^1\mathrm{S}_0)5s\ ^2\!\mathrm{S}_{1/2}\\ \\ (^1\mathrm{D}_2)5s\ ^2\mathrm{D}_{5/2}-\!\!-\!\!6p'\ ^2\mathrm{D}_{3/2}^{\circ} \end{array}\right.$
7663, 39 7663, 50 7663, 63 7680, 56 7680, 74		13045. 46 13045. 28 13045. 06 13016. 30 13016. 00	$ \left. \begin{array}{c} \left\{ \begin{array}{cccccccccccccccccccccccccccccccccccc$
7681. 75 7685. 11 7698. 72 7699. 58 7703. 17	20 50 35 80 2	13014, 29 13008, 60 12985, 60 12984, 15 12978, 10	$\begin{array}{c} 5p'\ ^2\mathrm{D}_{3/2}^{\ast}-7d\ ^4\mathrm{D}_{1/2}\\ 4d\ ^4\mathrm{D}_{1/2}-7f\ [2]_{3/2}^{\ast}\\ 5p''\ ^2\mathrm{P}_{1/2}^{\ast}-8d\ ^4\mathrm{F}_{3/2}\\ 5s'\ ^4\mathrm{P}_{1/2}-5p''\ ^2\mathrm{P}_{1/2}^{\ast}\\ 4d\ ^4\mathrm{D}_{3/2}-9p\ ^4\mathrm{P}_{1/2}^{\ast}\end{array}$
7704. 32 7706. 69 7708. 39 7709. 35 7709. 55	$60h \\ 15hw \\ 8 \\ 15h \\ 15h$	12976. 16 12972. 17 12969. 31 12967. 69 12967. 36	$\begin{array}{c} 4d\ ^4\mathrm{F}_{9/2} -\!\!-\! 8f\ [5]_{1^5/2}^{\bullet} \\ 4d\ ^4\mathrm{F}_{5/2} -\!\!-\! 6f\ [1]_{3^{\circ}/2}^{\bullet} \\ 4d\ ^4\mathrm{F}_{9/2} -\!\! 8f\ [4]_{9/2}^{\circ} \\ 4d\ ^4\mathrm{D}_{5/2} -\!\! 4f''\ [3]_{3^{\circ}/2}^{\circ} \\ 4d\ ^4\mathrm{D}_{5/2} -\!\! 4f''\ [3]_{7/2}^{\circ} \end{array}$
7711. 68 7713. 27 7713. 54 7715. 06 7715. 16	$\begin{bmatrix} 5 \\ 200 \\ 45 \\ 100 \\ 150 \end{bmatrix}$	12963. 78 12961. 10 12960. 65 12958. 10 12957. 93	$\begin{array}{c} 4d\ ^{4}\mathrm{D}_{3/2}-9p\ ^{4}\mathrm{P}_{3/2}^{2}\\ 4d\ ^{4}\mathrm{D}_{5/2}-6f\ [2]_{5/2}^{6}\\ 4d\ ^{4}\mathrm{D}_{5/2}-6f\ [2]_{3/2}^{6}\\ \end{array}$
7715. 31 7715. 48 7717. 18 7721. 45 7726. 02	200 400 2 300 60	12957. 68 12957. 39 12954. 54 12947. 37 12939. 71	$ \begin{vmatrix} (^{1}D_{2}) 5s \ ^{2}D_{5/2} - 6p' \ ^{2}D_{5/2}^{*} \\ 5p'' \ ^{2}P_{3/2}^{*} - 8d \ ^{4}D_{1/2} \\ 4d \ ^{4}D_{5/2} - 6f \ [4]_{7/2}^{?} \\ \end{vmatrix} $
7726, 13 7726, 26 7729, 19 7733, 61 7734, 61	80 110 1 900 25	12939. 53 12939. 31 12934. 41 12927. 01 12925. 34	$ \begin{cases} (^{1}\mathrm{D}_{2})5s~^{2}\mathrm{D}_{3/2} - 6p'~^{2}\mathrm{D}_{5/2}^{\circ} \\ 5p'~^{4}\mathrm{D}_{1/2} - 7d^{-4}\mathrm{D}_{3/2} \\ 4d~^{4}\mathrm{D}_{5/2} - 6f~[3]_{7/2}^{\circ},~_{5/2} \\ 5p'~^{2}\mathrm{S}_{1/2}^{\circ} - 5d''~^{2}\mathrm{D}_{3/2}^{\circ} \end{cases} $
7742, 05 7783, 44 7795, 01 7803, 02 7807, 52	$ \begin{array}{c} 1\\ 9w\\ 15\\ 30000\\ \end{array} $	12912. 92 12844. 26 12825. 19 12812. 03 12804. 64	$ \begin{array}{c} 4d\ ^4\mathrm{F}_{7/2} \!$
7807, 64 7807, 78 7821, 10 7827, 23 7835, 08	60 100 30 1200 30	12804. 44 12804. 21 12782. 41 12772. 40 12759. 60	$ \begin{vmatrix} (^{1}\mathrm{D}_{2})5s\ ^{2}\mathrm{D}_{3/2} - 6p'\ ^{2}\mathrm{S}_{1/2}^{\circ} \\ 5p''\ ^{2}\mathrm{P}_{1/2}^{\circ} - 8d\ ^{4}\mathrm{D}_{1/2} \\ 5p'\ ^{2}\mathrm{D}_{5/2}^{\circ} - 7s'\ ^{2}\mathrm{P}_{3/2} \\ 5p'\ ^{4}\mathrm{S}_{3/2}^{\circ} - (^{1}\mathrm{S}_{0})5s\ ^{2}\mathrm{S}_{1/2} \end{vmatrix} $
7841. 87 7843. 58 7844. 01 7844. 10 7869. 53	30 600 80 80 10	12748. 55 12745. 77 12745. 07 12744. 93 12703. 74	$\left.\begin{array}{c} 5p' \ ^4\mathrm{S}_{3/2}^{\circ}-7d \ ^4\mathrm{F}_{3/2} \\ 5p \ ^4\mathrm{P}_{3/2}^{\circ}-5d \ ^4\mathrm{F}_{5/2} \\ 5p \ ^4\mathrm{P}_{3/2}^{\circ}-5d \ ^4\mathrm{P}_{1/2} \\ 4d \ ^4\mathrm{F}_{9/2}-10p \ ^4\mathrm{D}_{7/2}^{\circ} \end{array}\right.$

See footnotes at end of table.

Table A2. Observed lines of Br I—Continued

	Tubile III		
Wavelength	Intensity	Wave number	Classification
Å		K	
7881. 45 7881. 57 7889. 85 7903. 66 7905. 70	$ \begin{array}{c} 2500s \\ 2500 \\ 600 \\ 50 \\ 50 \end{array} $	12684, 53 12684, 34 12671, 03 12648, 89 12645, 62	$ \left. \begin{array}{l} 5s'{}^{2}\mathrm{P}_{3/2} \!\!-\!\! 5p''{}^{2}\mathrm{P}_{1/2}^{\alpha} \\ 5p'{}^{2}\mathrm{D}_{3/2}^{\alpha} \!\!-\!\! 7s'{}^{4}\mathrm{P}_{1/2} \\ 5p''{}^{2}\mathrm{P}_{1/2}^{\alpha} \!\!-\!\! (^{1}\mathrm{D}_{2})4d{}^{2}\mathrm{D}_{3/2} \\ 5p'{}^{4}\mathrm{S}_{3/2}^{\alpha} \!\!-\!\! 7d{}^{4}\mathrm{P}_{1/2} \end{array} \right. $
7925. 81 7929. 69 7932. 99 7938. 68 7940. 07	2500 250 - 30W 30000c 60	12613. 54 12607. 37 12602. 12 12593. 09 12590. 88	$ \left\{ \begin{array}{c} 5p\ ^4\mathrm{D}_{7/2}^{\circ}{-}7s\ ^4\mathrm{P}_{5/2} \\ 5p'\ ^2\mathrm{D}_{3/2}^{\circ}{-}7s'\ ^2\mathrm{P}_{3/2} \\ 4d\ ^4\mathrm{D}_{3/2}{-}6f\ [1]_{3/2}^{\circ} \\ 4d\ ^4\mathrm{D}_{3/2}{-}6f\ [1]_{1/2}^{\circ} \\ (^1\mathrm{D}_2)5s\ ^2\mathrm{D}_{5/2}^{\circ}{-}(^1\mathrm{D}_2)5p\ ^2\mathrm{D}_{5/2}^{\circ} \\ 4d\ ^4\mathrm{D}_{3/2}{-}6f\ [2]_{3/2}^{\circ} \end{array} \right. $
7947. 94 7950. 18 7961. 33 7966. 84 7967. 03	3000 3000 350 350 600 750	12578. 42 12574. 87 12557. 26 12548. 58 12548. 28	$ \begin{array}{c} 5p\ ^4\mathrm{D}_{5/2}^{8}-7s\ ^4\mathrm{P}_{3/2} \\ (^1\mathrm{D}_{2})5s\ ^2\mathrm{D}_{3/2}-(^1\mathrm{D}_{2})5p\ ^2\mathrm{D}_{5/2}^{8} \\ 4d\ ^4\mathrm{D}_{3/2}-6f\ [3]_{5/2}^{8} \end{array} \\ \left. \begin{array}{c} (^1\mathrm{D}_{2})5s\ ^2\mathrm{D}_{5/2}-(^1\mathrm{D}_{2})5p\ ^2\mathrm{D}_{3/2}^{8} \end{array} \right. \end{array}$
7975. 11 7975. 23 7975. 39 7975. 62 7978. 44	$ \begin{cases} 10 \\ 20 \\ 15 \\ 35 \\ 8000 \end{cases} $	12535. 56 12535. 38 12535. 12 12534. 76 12530. 33	$ \begin{cases} (^{1}D_{2}) 5s \ ^{2}D_{5/2} - 4f \ [1]_{3/2} \\ (^{1}D_{2}) 5s \ ^{2}D_{5/2} - 4f \ [1]_{3/2} \end{cases} $
7978. 57 7989. 94 7997. 02 7998. 75 8004. 36	10000 30000 80 175 25	12530. 13 12512. 30 12501. 22 12498. 52 12489. 76	$ \begin{cases} (^1\mathrm{D}_2) 58 ^2\mathrm{D}_{3/2} - (^1\mathrm{D}_2) 5p ^2\mathrm{D}_{3/2}^2 \\ 58' ^2\mathrm{P}_{3/2} - 5p'' ^2\mathrm{P}_{3/2}^3 \\ 4d ^4\mathrm{D}_{7/2} - 4f' [3]_{7/2}^{7/2} \\ 5p' ^2\mathrm{D}_{3/2}^3 - 5d'' ^2\mathrm{D}_{5/2} \\ 5p' ^2\mathrm{D}_{5/2}^3 - 5d'' ^2\mathrm{D}_{3/2} \end{cases} $
8008. 87 8009. 36 8009. 77 8009. 86 8010. 00	$ \begin{array}{c} 30h \\ 75 \\ 100 \\ 150 \\ 200 \end{array} $	12482. 72 12481. 96 12481. 32 12481. 18 12480. 96	$ \begin{cases} 4d {}^{4}\mathrm{F}_{7/2} \!\!-\!\! 8f [5]_{3/2}^{8/2} \\ 5p' {}^{4}\mathrm{D}_{1/2}^{\circ} \!\!-\!\! 7s' {}^{4}\mathrm{P}_{1/2} \\ \\ ({}^{1}\mathrm{D}_{2}) 5s {}^{2}\mathrm{D}_{5/2} \!\!-\!\! 4f [2]_{5/2}^{8/2} \end{cases} $
8010. 17 8013. 03 8014. 51 8014. 60 8014. 74	$ \begin{array}{c} $	12480. 70 12476. 24 12473. 94 12473. 80 12473. 58	$ \begin{array}{c} $
8014. 89 8021. 57 8021. 67 8021. 80 8022. 52	$ \begin{array}{c} 200 \\ 100 \\ 150 \\ 200 \\ 800 \end{array} $	12473. 35 12462. 96 12462. 80 12462. 60 12461. 48	$ \left. \begin{array}{l} \\ \\ \\ \\ \\ \\ \\ \end{array} \right\} = (^{1}\mathrm{D}_{2})5s^{2}\mathrm{D}_{3/2} -\!$
8022. 83 8022. 98 8023. 14 8023. 91 8026. 35	$ \begin{cases} 15 \\ 15 \\ 20 \\ 900 \\ 2000 \end{cases} $	12461. 00 12460. 77 12460. 52 12459. 33 12455. 54	$ \left. \begin{array}{c} \left\{ \begin{array}{c} (^{1}\mathrm{D}_{2})5s^{2}\mathrm{D}_{5/2} - 4f[3]_{7/2}^{2} \\ \\ 5p^{4}\mathrm{P}_{5/2}^{2} - 5d^{4}\mathrm{D}_{3/2} \\ \\ \left\{ \begin{array}{c} (^{1}\mathrm{D}_{2})5s^{2}\mathrm{D}_{3/2} - 4f[2]_{3/2}^{3} \end{array} \right. \end{array} \right. $
8026. 54 8028. 22 8028. 80 8029. 48 8029. 53	$ \begin{array}{c c} $	12455. 24 12452. 64 12451. 74 12450. 68 12450. 60	$ \begin{cases} Ad^{4}D_{7/2} - 4f' [4]_{7/2}^{8/2} \\ Ad^{4}D_{7/2} - 4f' [4]_{8/2}^{8/2} \\ Ad^{4}D_{7/2} - 4f' [4]_{7/2}^{8/2} \end{cases} $
8029. 67 8029. 81 8030. 91 8033. 40	$ \begin{vmatrix} 200 \\ 250 \\ 250h \\ 40 \end{vmatrix} $	12450. 39 12450. 17 12448. 47 12444. 61	$ \left\{ \begin{array}{c} (^{4}D_{2})^{\frac{1}{2}} 58 ^{4}D_{5/2} - 4f [^{4}]_{7/2} \\ 4d ^{4}F_{9/2} - 7f [^{5}]_{1.72}^{\circ} \\ 4d ^{4}D_{7/2} - 8p ^{4}P_{5/2}^{\circ} \end{array} \right. $

Table A2. Observed lines of Br I—Continued

	TABLE A2.	br i—Continued	
Wavelength	Intensity	Wave number	Classification
$reve{A}$		K	
8034. 83	75	12442. 39	
8034. 91 8035. 03	100 150	12442. 27 12442. 08	$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$
8037. 34	40	12438. 51	$\left\{egin{array}{cccccccccccccccccccccccccccccccccccc$
8037. 65 8050. 41	60 15	12438. 03 12418. 31	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$
8066. 38 8072. 89	200 300	12393. 73 12383. 73	$5p^{\prime\prime}{}^{2}\mathrm{P}_{3/2}^{\circ}$ $-7s^{\prime\prime}{}^{2}\mathrm{P}_{1/2}$ $5p^{\prime}{}^{2}\mathrm{S}_{1/2}^{\circ}$ $-5d^{\prime}{}^{2}\mathrm{P}_{3/2}$
8101. 21 8111. 56	150 50	12340. 44 12324. 70	$5p'\ {}^{2}\mathrm{S}_{1/2}^{6}-6d\ {}^{4}\mathrm{F}_{3/2}^{3/2} \ 5p'\ {}^{2}\mathrm{D}_{3/2}^{3/2}-5d''\ {}^{2}\mathrm{D}_{3/2}^{3/2}$
8113. 04 8131. 52	$\frac{20h}{30000}$	12322. 45 12294. 44	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$
8137. 89	150	12284. 82	$4d\ ^4\mathrm{D}_{5/2}$ $-4f'\ [3]_{7/2}^6$
8142. 80	4 0h	12277. 41	$\left\{egin{array}{cccccccccccccccccccccccccccccccccccc$
8148. 84 8152. 65	$\frac{20}{1000c}$	12268. 31 12262. 58	$\begin{array}{c} 4d \ ^{4}\mathrm{D}_{5/2} - 8p \ ^{4}\mathrm{P}_{3/2}^{3/2} \\ 5p \ ^{4}\mathrm{P}_{3/2}^{3} - 5d \ ^{4}\mathrm{D}_{1/2} \end{array}$
8153. 75	10000	12260. 92	$5p^{-4}\mathrm{P}_{5/2}^{\circ}$ $-5d^{-4}\mathrm{D}_{5/2}^{1/2}$
8154. 00 8166. 32	25000 175	12260. 55 12242. 05	$rac{5s\ ^4 ext{P}_{5/2} ext{}5p\ ^4 ext{D}_{5/2}^{\circ}}{5p^{\prime\prime}\ ^2 ext{P}_{3/2}^{\circ} ext{}(^1 ext{S}_0)5s\ ^2 ext{S}_{1/2}}$
8170. 21 8172. 09	$\frac{200d}{200}$	12236. 22 12233. 41	$\begin{array}{c} 4d\ ^{4}\mathrm{D}_{5/2}^{5/2}-4f'\ [4]_{7/2}^{9/2}\\ 5p\ ^{4}\mathrm{D}_{7/2}^{9/2}-5d\ ^{4}\mathrm{F}_{5/2} \end{array}$
8173. 71	20	12230. 98	$5p^{\prime\prime}{}_{^{2}}\!$
8175. 58 8179. 65	25 40	12228. 19 12222. 10	$rac{4d\ ^4\mathrm{D}_{5/2} ext{}8p\ ^4\mathrm{P}_{5/2}^{\circ}}{4d\ ^4\mathrm{D}_{5/2} ext{}4f'} rac{[2]_{5/2}^{\circ}}{[2]_{5/2}^{\circ}}$
8180. 01 8183. 52	100 450	12221. 56 12216. 32	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$
8189. 93	1hw	12206. 76	$\left. igg egin{array}{cccccccccccccccccccccccccccccccccccc$
8190. 08	$1hw$	12206. 54	J 51/2 Of [1]3/2
8190. 28 8197. 73	$\frac{40}{175h}$	12206. 24 12195. 15	$4d\ ^4\mathrm{D}_{1/2}$ $-6f\ [1]_{1/2}^{\circ}$ $4d\ ^4\mathrm{D}_{1/2}$ $-6f\ [2]_{3/2}^{\circ}$
8210. 49 8215. 12	$\frac{3h}{200}$	12176. 19 12169. 33	$\begin{array}{c} 4d\ ^{4}\mathrm{D}_{1/2}6f\ [2]_{3/2}^{3/2} \\ 4d\ ^{4}\mathrm{F}_{3/2}5f'\ [3]_{5/2}^{5} \\ 5p'\ ^{4}\mathrm{S}_{3/2}^{3}7s'\ ^{4}\mathrm{P}_{1/2} \end{array}$
8237. 96	500	12135. 59	$5p' {}^4\mathrm{D}_{1/2}^{\circ} - 5d'' {}^2\mathrm{D}_{3/2}$
8246. 86 8247. 97	5000 30	12122. 50 12120. 86	$5p\ ^4\mathrm{P}_{3/2}^{\circ}-5d\ ^4\mathrm{D}_{3/2} \ 5p'\ ^2\mathrm{D}_{3/2}^{\circ}-8s\ ^4\mathrm{P}_{3/2}$
$8252.\ 38$ $8253.\ 87$	30 35	12114. 39 12112. 20	$5p^{\prime\prime}{}^{_{2}}\mathrm{P}^{_{3/2}}_{_{3/2}}$ $-7d{}^{_{4}}\mathrm{D}^{_{5/2}}_{_{5/2}}$ $5p^{\prime}{}^{_{2}}\mathrm{D}^{_{3/2}}_{_{3/2}}$ $-8s{}^{_{4}}\mathrm{P}^{_{5/2}}_{_{5/2}}$
8258. 32	300	12105. 67	$5p'$ ${}^4\mathrm{S}_{3/2}^{\circ}$ $-7s'$ ${}^2\mathrm{P}_{3/2}$
8264. 96 8272. 44	$15000 \\ 75000c$	12095. 95 12085. 01	$egin{array}{cccccccccccccccccccccccccccccccccccc$
8280. 24 8280. 75	$100 \\ 100w$	12073. 63 12072. 88	$rac{5p'^2\mathrm{D}_{5/2}^2-5d'^4\mathrm{P}_{5/2}}{5p^4\mathrm{D}_{5/2}^2-5d^4\mathrm{F}_{3/2}}$
8291. 06	900	12057. 87	$5p~^4{ m D}_{5/2}^2-5d~^4{ m F}_{5/2}$
8293. 57 8308. 49	$\frac{100}{20}$	12054. 22 12032. 57	$5p^{\prime}{}^{2}{ m D}_{5/2}^{\circ}\!\!\!-\!\!\!\!-\!\!\!\!-\!\!\!\!\!-\!\!\!\!\!-\!\!\!\!\!\!\!\!\!\!\!$
8313. 09 8334. 70	$\frac{10}{20000}$	12025. 92 11994. 74	$5p'\ ^{2}{ m D}_{5/2}^{6/2} - 5d'\ ^{4}{ m P}_{3/2}^{3/2} \ 5s'\ ^{2}{ m P}_{3/2} - 5p'\ ^{4}{ m S}_{3/2}^{8/2}$
8343.70	10000	11981. 80	$5s' {}^4P_{1/2} - 5p' {}^4D_{1/2}^{\circ}$
8361. 71 8362. 38	300 100h	11955. 99 11955. 03	$egin{array}{cccccccccccccccccccccccccccccccccccc$
8367. 20 8369. 00	$\frac{15}{20}$	11948. 15 11945. 58	$5p'{}^{2}\mathrm{D}_{5/2}^{\circ}-5d'{}^{2}\mathrm{P}_{3/2}\ 4d{}^{4}\mathrm{F}_{7/2}-7f[4]_{7/2}^{\circ}$
8372. 79 8378. 65	150	11940. 17	5p 4P _{1/2} —5d 4P _{1/2}
8382. 42	$100w \\ 5 \\ 1200$	11931. 82 11926. 45	$5p^7 {}^4\mathrm{D}_{1/2}^{\circ} - 8s {}^4\mathrm{P}_{3/2} \ 4d {}^4\mathrm{D}_{3/2}^{\circ} - 8p {}^4\mathrm{D}_{5/2}^{\circ}$
8384. 04 8387. 99	$\begin{array}{c} 1200 \\ 100 \end{array}$	11924. 15 11918. 53	$\begin{array}{cccccccccccccccccccccccccccccccccccc$

Table A2. Observed lines of Br I—Continued

Wavelength	Intensity	Wave number	Classification
\mathring{A}		K	
8389. 75 8392. 76 8397. 62 8402. 08 8408. 68	300 75 50 20 250	11916. 03 11911. 76 11904. 86 11898. 55 11889. 21	$egin{array}{cccccccccccccccccccccccccccccccccccc$
8410. 79 8411. 61 8421. 88 8428. 77 8430. 51	$ \begin{array}{r} 10w \\ 5w \\ 4 \\ 75 \\ 50 \end{array} $	11886. 22 11885. 06 11870. 57 11860. 87 11858. 42	$rac{4d}{4} rac{4}{5_{5/2}} - 8f [3]_{7/2,5/2}^{\circ} \ 4d \ ^4F_{5/2} - 8f [4]_{7/2}^{\circ} \ 5p' \ ^2D_{3/2}^{\circ} - 5d' \ ^4P_{3/2} \ 4d \ ^4D_{3/2} - 8p \ ^4P_{5/2}^{\circ}$
8434. 83 8435. 34 8441. 26 8446. 55 8454. 25	150 75 100 40000 75	11852. 35 11851. 63 11843. 32 11835. 90 11825. 12	$egin{array}{c} 4d\ ^4\mathrm{D}_{3/2} \!\!-\! 4f'\ [2]_{5/2}^3 \ 4d\ ^4\mathrm{D}_{3/2} \!\!-\! 4f'\ [2]_{3/2}^3 \ 5p'\ ^2\mathrm{D}_{5/2}^3 \!\!-\! 6d\ ^4\mathrm{F}_{5/2} \ 5s\ ^4\mathrm{P}_{3/2} \!\!-\! 5p\ ^4\mathrm{D}_{3/2}^3 \ 4d\ ^4\mathrm{D}_{7/2} \!\!-\! 5f\ [2]_{5/2}^3 \ \end{array}$
8467. 38 8470. 62 8471. 51 8477. 45 8484. 40	300 100 500 4000 30	11806. 78 11802. 27 11801. 03 11792. 76 11783. 10	$egin{array}{l} 4d\ ^4\mathrm{D}_{7/2}5f\ [3]_{7/2,5/2}^2 \ 4d\ ^4\mathrm{D}_{7/2}5f\ [4]_{7/2}^2 \ 4d\ ^4\mathrm{D}_{7/2}5f\ [4]_{3/2}^2 \ 5s'\ ^4\mathrm{P}_{1/2}5p'\ ^2\mathrm{D}_{3/2}^3 \ 5p'\ ^2\mathrm{D}_{3/2}^35d'\ ^2\mathrm{P}_{3/2} \end{array}$
8503. 78 8513. 38 8515. 69 8529. 38 8557. 73	$ \begin{array}{r} 400 \\ 1500 \\ 100 \\ 8h \\ 1000 \end{array} $	11756. 25 11742. 99 11739. 80 11720. 96 11682. 13	$egin{array}{cccccccccccccccccccccccccccccccccccc$
8560, 57 8565, 30 8566, 28 8578, 84 8580, 00	300 200 1000 400 20	11678. 26 11671. 81 11670. 47 11653. 39 11651. 81	$egin{array}{cccccccccccccccccccccccccccccccccccc$
8592, 30 8595, 50 8595, 67 8603, 00 8603, 69	250 50 50 2 100	11635. 13 11630. 80 11630. 57 11620. 66 11619. 73	$4d\ ^4\mathrm{F}_{9/2}$ — $6f\ [5]_{11/2}^{\circ}$ $4d\ ^4\mathrm{F}_{5/2}$ — $5f\ [1]_{3/2}$ $4d\ ^4\mathrm{F}_{9/2}$ — $6f\ [4]_{7/2}^{\circ}$ $4d\ ^4\mathrm{F}_{9/2}$ — $6f\ [4]_{9/2}^{\circ}$
8604. 06 8606. 68 8608. 98 8610. 49 8611. 85	20hw 20 $30hw$ 100 250	11619. 23 11615. 69 11612. 59 11610. 55 11608. 72	$\left\{ \begin{array}{c} 5p' \ ^{4}\mathrm{S}_{3/2}^{\circ} - \mathrm{Ss} \ ^{4}\mathrm{P}_{3/2} \\ 5p' \ ^{2}\mathrm{D}_{3/2}^{\circ} - 5d' \ ^{2}\mathrm{P}_{1/2} \\ 4d \ ^{4}\mathrm{P}_{1/2} - 7f \ [1]_{3/2}^{\circ} \\ 4d \ ^{4}\mathrm{P}_{1/2} - 7f \ [1]_{1/2}^{\circ} \\ 5p' \ ^{4}\mathrm{S}_{3/2}^{\circ} - \mathrm{Ss} \ ^{4}\mathrm{P}_{5/2} \\ 4d \ ^{4}\mathrm{D}_{5/2} - 5f \ [2]_{5/2}^{\circ} \end{array} \right.$
8612. 48 8625. 38 8625. 47 8628. 84 8638. 66	75 750c 100 200 20000	11607. 87 11590. 51 11590. 39 11585. 86 11572. 69	$\begin{array}{c} 4d\ ^4\mathrm{D}_{5/2} \!$
8655. 11 8658. 84 8668. 84 8698. 53 8708. 68	100 10 300 4000 75	11550. 70 11545. 72 11532. 40 11493. 04 11479. 64	$egin{array}{cccccccccccccccccccccccccccccccccccc$
8708. 96 8725. 33 8726. 75 8741. 64 8742. 99	$ \begin{array}{r} 100 \\ 500 \\ 100c \\ \hline 50 \\ 75 \end{array} $	11479. 27 11457. 74 11455. 87 11436. 36 11434. 59	$egin{array}{cccccccccccccccccccccccccccccccccccc$

Table A2. Observed lines of Br I—Continued

Wavelength	Intensity	Wave number	Classification
\overline{A}		K	
8749. 10 8757. 24 8760. 46 8764. 20 8775. 71	10 10 700 500 250	11426. 61 11415. 99 11411. 79 11406. 92 11391. 96	$egin{array}{cccccccccccccccccccccccccccccccccccc$
8779. 15 8793. 47 8798. 04 8801. 03 8803. 41 8804. 76 8807. 44 8807. 68 8808. 85 8819. 96	$ \begin{array}{c} 50 \\ 10000c \\ 6 \\ 150 \\ 100 \\ 20 \\ 600 \\ 800 \\ 500 \\ 15000 \end{array} $	11387. 50 11368. 95 11363. 05 11359. 19 11356. 12 11354. 38 11350. 92 11350. 61 11349. 10 11334. 81	$ \begin{cases} 5p' & ^4\mathrm{S}_{3/2}^\circ - 5d' & ^2\mathrm{F}_{5/2} \\ (^1\mathrm{D}_2) 5s & ^2\mathrm{D}_{5/2} - (^1\mathrm{D}_2) 5p & ^2\mathrm{P}_{3/2}^\circ \\ & 4d & ^4\mathrm{F}_{5/2} - ^7f & [2]_5^\circ / 2 \\ & 5p' & ^4\mathrm{S}_{3/2}^\circ - 5d' & ^4\mathrm{P}_{3/2} \\ & 4d & ^4\mathrm{F}_{5/2} - ^7f & [3]_{7/2}^\circ , _{5/2} \\ & 5p' & ^2\mathrm{D}_{5/2}^\circ - 6d & ^4\mathrm{D}_{3/2} \\ & 4d & ^4\mathrm{F}_{5/2} - ^7f & [4]_{7/2}^\circ \\ & \\ & (^1\mathrm{D}_2) 5s & ^2\mathrm{D}_{3/2} - (^1\mathrm{D}_2) 5p & ^2\mathrm{P}_{3/2}^\circ \\ & & 5p' & ^2\mathrm{D}_{5/2}^\circ - 6d & ^4\mathrm{D}_{7/2} \\ & (^1\mathrm{D}_2) 5s & ^2\mathrm{D}_{5/2} - (^1\mathrm{D}_2) 5p & ^2\mathrm{F}_{7/2}^\circ \end{cases} $
8825. 22 8833. 33 8842. 85 8861. 69 8869. 64	25000 175 125 85 300	11328, 05 11317, 65 11305, 47 11281, 43 11271, 32	$\begin{array}{c} 5s'\ ^2\mathrm{P}_{3/2} \!$
8870. 46 8875. 43 8877. 89 8878. 40 8886. 34	200 3 90 90 40	11270. 28 11263. 97 11260. 85 11260. 20 11250. 14	$\begin{array}{c} 5p\ ^4\mathrm{P}_{3/2}^{8}-6s''\ ^2\mathrm{P}_{1/2} \\ 6s\ ^4\mathrm{P}_{5/2}-7p''\ ^2\mathrm{P}_{3/2}^{8} \\ 4d\ ^4\mathrm{D}_{3/2}-5f\ [1]_{3/2}^{8} \\ 4d\ ^4\mathrm{D}_{3/2}-5f\ [1]_{1/2}^{8} \\ 5p'\ ^4\mathrm{S}_{3/2}^{8}-6d\ ^4\mathrm{P}_{1/2} \end{array}$
8888. 98 8895. 21 8895. 87 8897. 62 8909. 73	4000 20 150 30000 300	11246. 80 11238. 92 11238. 09 11235. 88 11220. 60	$ \begin{cases} 5p \ ^4\mathrm{D}_{7/2} - 5d \ ^4\mathrm{D}_{7/2} \\ 4d \ ^4\mathrm{D}_{3/2} - 5f \ [2]_{5/2}^2 \\ 4d \ ^4\mathrm{D}_{3/2} - 5f \ [2]_{5/2}^2 \\ 5p' \ ^4\mathrm{S}_{3/2} - 6d \ ^4\mathrm{F}_{3/2} \\ 5s \ ^4\mathrm{P}_{5/2} - 5p \ ^4\mathrm{P}_{5/2}^2 \\ 4d \ ^4\mathrm{D}_{3/2} - 5f \ [3]_{5/2}^4 \end{cases} $
8925. 64 8932. 40 8933. 29 8936. 96 8944. 83	$ \begin{array}{c} 15 \\ 6000 \\ 25 \\ 15 \\ 50 \end{array} $	11200. 60 11192. 13 11191. 01 11186. 42 11176. 57	$egin{array}{l} 4d\ ^4{ m F}_{3/2}7f\ [1]_{3/2}^2\ 5s'\ ^4{ m P}_{1/2}-5p'\ ^2{ m S}_{1/2}^3\ 5p'\ ^2{ m D}_{3/2}^3-6d\ ^4{ m D}_{3/2}\ 4d\ ^4{ m F}_{3/2}^3-7f\ [3]_{5/2}^4\ 5p'\ ^4{ m S}_{3/2}^3-6d\ ^4{ m F}_{5/2} \end{array}$
8949. 39 8964. 00 8972. 83 8979. 59 8984. 56	1800 9000 350 200 100	11170. 88 11152. 67 11141. 70 11133. 31 11127. 15	$\substack{\substack{(^{1}\mathrm{D}_{2})5s\ ^{2}\mathrm{D}_{5/2}-(^{1}\mathrm{D}_{2})5p\ ^{2}\mathrm{F}_{5/2}^{\circ}\\ (^{1}\mathrm{D}_{2})5s\ ^{2}\mathrm{D}_{3/2}-(^{1}\mathrm{D}_{2})5p\ ^{2}\mathrm{F}_{5/2}^{\circ}\\ 4d\ ^{4}\mathrm{F}_{7/2}-6f\ [5]_{5/2}^{\circ}\\ 5p^{\prime\prime}\ ^{2}\mathrm{P}_{1/2}^{\circ}-5d^{\prime\prime}\ ^{2}\mathrm{D}_{3/2}\\ 4d\ ^{4}\mathrm{F}_{7/2}-6f\ [4]_{7/2}^{\circ}\\ }$
8985. 32 8995. 21 9000. 99 9005. 20 9012. 23	$ \begin{array}{c} 1 \\ 50 \\ 2h \\ 50h \\ 20 \end{array} $	11126. 21 11113. 98 11106. 84 11101. 65 11092. 99	$\begin{array}{c} 4d\ ^4\mathrm{F}_{7/2} \!$
9029. 94 9045. 06 9066. 28 9078. 64 9079. 29	300 40 400 700 125	11071, 23 11052, 72 11026, 86 11011, 84 11011, 05	$egin{array}{cccccccccccccccccccccccccccccccccccc$
9084, 29 9086, 80 9107, 37 9123, 17 9126, 14	$\begin{array}{c} 25 \\ 100 \\ 10 \\ 50 \\ 4 \end{array}$	11004. 99 11001. 95 10977. 11 10958. 09 10954. 53	$\begin{array}{c} 4d\ ^4\mathrm{F}_{5/2} \!$

Table A2. Observed lines of Br I—Continued

	1110000 1121		
Wavelength	Intensity	Wave number	Classification
Å		K	
9129. 31 9147. 04 9166. 06 9173. 63 9178. 16	70 75h 30000 15000 20000	10950, 72 10929, 50 10906, 82 10897, 82 10892, 44	$egin{array}{c} 4d\ ^4\mathrm{F}_{9/2}-8p\ ^4\mathrm{D}_{7/2}^{\circ}\ 5p''\ ^2\mathrm{P}_{1/2}^{\circ}-8s\ ^4\mathrm{P}_{3/2}\ 5s\ ^4\mathrm{P}_{3/2}-5p\ ^4\mathrm{P}_{1/2}^{\circ}\ 5s''\ ^2\mathrm{P}_{1/2}-5p''\ ^2\mathrm{P}_{1/2}^{\circ}\ 5s''\ ^2\mathrm{P}_{3/2}-5p'\ ^2\mathrm{S}_{1/2}^{\circ} \end{array}$
9183. 14	30	10886. 53	$\left\{egin{array}{cccccccccccccccccccccccccccccccccccc$
9197. 16 9201. 21 9201. 78 9202. 47	50 80w 80 80	10869. 94 10865. 15 10864. 48 10863. 67	$\begin{array}{c} 4d^{-1}_{1/2} - jj & ^{[2]_{3/2}}_{5p'} \\ 5p'' ^2\mathrm{P}_{3/2} - 5d' ^2\mathrm{F}_{5/2} \\ 4d ^4\mathrm{D}_{1/2} - 5f [1]_{3/2}^2 \\ 4d ^4\mathrm{D}_{1/2} - 5f [1]_{1/2}^2 \\ 5p' ^4\mathrm{D}_{1/2}^2 - 6d ^4\mathrm{D}_{1/2} \end{array}$
9220, 56 9221, 19 9226, 57 9254, 06 9254, 39	200 35 15 70 30	10842. 35 10841. 61 10835. 29 10803. 10 10802. 72	$4d\ ^4\mathrm{D}_{1/2}$ — $5f\ [2]_{3/2}^{8/2}$ $5p''\ ^2\mathrm{P}_{3/2}^{8/2}$ — $5d'\ ^4\mathrm{P}_{3/2}$ $4d\ ^4\mathrm{P}_{3/2}$ — $9p\ ^4\mathrm{D}_{3/2}^{8/2}$ $4d\ ^4\mathrm{P}_{1/2}$ — $6f\ [1]_{3/2}^{8/2}$ $4d\ ^4\mathrm{P}_{1/2}$ — $6f\ [1]_{1/2}^{8/2}$
9265. 42 9282. 80 9287. 82 9288. 52 9314. 85	$40000 \\ 100 \\ 2 \\ 200 \\ 30$	10789. 86 10769. 66 10763. 84 10763. 02 10732. 60	$\begin{array}{c} 58\ ^4\mathrm{P}_{3/2} - 5p\ ^4\mathrm{D}_{5/2}^\circ \\ 5p'\ ^4\mathrm{S}_{3/2} - 6d\ ^4\mathrm{D}_{5/2} \\ 5p''\ ^2\mathrm{P}_{3/2} - 5d'\ ^2\mathrm{P}_{3/2} \\ 5p\ ^4\mathrm{P}_{5/2}^\circ - 6s'\ ^2\mathrm{P}_{3/2} \\ 5p''\ ^2\mathrm{P}_{3/2}^\circ - 6d\ ^4\mathrm{P}_{1/2} \end{array}$
9320. 86 9325. 32 9352. 55 9354. 56 9369. 94	15000 20 175 15 100	10725. 68 10720. 55 10689. 34 10687. 04 10669. 50	$egin{array}{cccccccccccccccccccccccccccccccccccc$
9378. 84 9379. 21 9399. 55 9400. 19 9409. 56	1 30 5 90 20	10659. 37 10658. 95 10635. 89 10635. 17 10624. 57	$\begin{array}{c} 4d'\ ^2\mathrm{F}_{7/2}5f'\ [3]_{5/2}^{\circ} \ 4d'\ ^2\mathrm{F}_{7/2}5f'\ [3]_{7/2}^{\circ} \ 4d'\ ^2\mathrm{F}_{7/2}5f'\ [4]_{7/2}^{\circ} \ 4d'\ ^2\mathrm{F}_{7/2}5f'\ [4]_{9/2}^{\circ} \ 4d\ ^4\mathrm{D}_{7/2}7\ p\ ^4\mathrm{D}_{5/2}^{\circ} \end{array}$
9428. 58 9434. 51 9438. 20 9460. 14 9466. 45	$\begin{array}{c} 3 \\ 3 \\ 80 \\ 500d \\ 2w \end{array}$	10603. 14 10596. 48 10592. 33 10567. 77 10560. 73	$\begin{array}{c} 68\ ^4\mathrm{P}_{5/2}\text{—-}7p'\ ^4\mathrm{S}_{3/2}^{\circ} \\ 5p''\ ^2\mathrm{P}_{3/2}^{\circ}\text{—-}5d'\ ^2\mathrm{P}_{1/2}^{\circ} \\ 4d\ ^4\mathrm{D}_{7/2}\text{—-}7p\ ^4\mathrm{D}_{7/2}^{\circ} \\ 5p\ ^4\mathrm{P}_{3/2}^{\circ}\text{—-}6s'\ ^4\mathrm{P}_{1/2}^{\circ} \\ 4d\ ^4\mathrm{F}_{5/2}\text{—-}6f\ [1]_{3/2}^{\circ} \end{array}$
9470. 67 9475. 11 9476. 32 9477. 50 9488. 66	20h 80 30 60 120	10556. 02 10551. 07 10549. 73 10548. 41 10536. 01	$egin{array}{l} 4d\ ^4{ m F}_{5/2}-4f''\ [3]_{7/2}^{\circ} \ 5p''\ ^4{ m S}_{3/2}-6d\ ^4{ m D}_{1/2} \ 4d\ ^4{ m F}_{5/2}-6f\ [2]_{5/2}^{\circ} \ 5p''\ ^2{ m P}_{1/2}^{\circ}-6d\ ^4{ m F}_{3/2} \ 4d\ ^4{ m F}_{5/2}-6f\ [4]_{7/2}^{\circ} \ \end{array}$
9495. 30 9498. 32 9503. 78 9507. 03 9517. 27	$ \begin{array}{r} 50 \\ 20 \\ 6w \\ 75 \\ 100 \end{array} $	10528. 64 10525. 29 10519. 24 10515. 65 10504. 33	$\begin{array}{c} 5p\ ^4\mathrm{D}_{3/2}^*-5d\ ^4\mathrm{D}_{1/2} \\ 4d\ ^4\mathrm{D}_{5/2}^*-7p\ ^4\mathrm{D}_{3/2}^* \\ 5p\ ^4\mathrm{P}_{5/2}^*-4d''\ ^2\mathrm{D}_{3/2} \\ 4d\ ^4\mathrm{F}_{5/2}^*-6f\ [3]_{7/2}^*,\ _{5/2}^* \\ 4d\ ^4\mathrm{D}_{7/2}^*-7p\ ^4\mathrm{P}_{5/2}^* \end{array}$
9543. 03 9552. 67 9556. 43 9560. 17 9576. 17	$ \begin{array}{r} 40 \\ 175 \\ 2h \\ 6 \\ 1 \end{array} $	10475. 98 10465. 41 10461. 29 10457. 20 10439. 72	$egin{array}{l} 4d\ ^4\mathbf{F}_{7/2}8p\ ^4\mathbf{D}_{5/2}^{\circ}\ 5p\ ^4\mathbf{P}_{7/2}^{\circ}6s''\ ^2\mathbf{P}_{1/2}\ 4d\ ^4\mathbf{F}_{7/2}4f'\ [3]_{5/2}^{\circ}\ 4d\ ^4\mathbf{F}_{7/2}8p\ ^4\mathbf{D}_{7/2}^{\circ}\ 4d'\ ^2\mathbf{P}_{3/2}5f'\ [3]_{5/2}^{\circ} \end{array}$
9588. 61 9590. 33 9597. 98 9598. 79 9603. 62	$ \begin{array}{c} 300 \\ 20 \\ \hline 2 \\ 60 \\ 35 \end{array} $	10426, 18 10424, 31 10416, 00 10415, 12 10409, 89	$\begin{array}{c} 5p\ ^4\mathrm{P}_{3/2}^{2}-6s'\ ^2\mathrm{P}_{3/2} \\ 5p''\ ^2\mathrm{P}_{7/2}^{2}-5d'\ ^2\mathrm{P}_{1/2} \\ 4d\ ^4\mathrm{F}_{7/2}-4f'\ [4]_{7/2}^{2} \\ 4d\ ^4\mathrm{F}_{7/2}-4f'\ [4]_{5/2}^{5} \\ 4d'\ ^2\mathrm{P}_{3/2}^{2}-5f'\ [2]_{5/2}^{5} \end{array}$

Table A2. Observed lines of Br I—Continued

TABLE A2. Coserved times of B1 1—Continued				
Wavelength	Intensity	Wave number	Classification	
Å		K		
9604. 30 9605. 34 9621. 02 9623. 33 9625. 17	$\begin{array}{c} 6 \\ 1 \\ 35h \\ 15 \\ 8 \end{array}$	10409. 15 10408. 02 10391. 06 10388. 56 10386. 58	$4d'\ ^2\mathrm{P}_{3/2}$ — $5f'\ [2]_{3/2}$ $4d\ ^4\mathrm{F}_{7/2}$ — $8p\ ^4\mathrm{P}_{5/2}$ $4d\ ^4\mathrm{F}_{3/2}$ — $6f\ [1]_{3/2}$ $5p\ ^4\mathrm{D}_{3/2}$ — $5d\ ^4\mathrm{D}_{3/2}$ $4d\ ^4\mathrm{F}_{3/2}$ — $4f''\ [3]_{5/2}^2$	
9627. 54 9631. 30 9635. 05 9661. 08 9662. 99	$\begin{array}{c} 4 \\ 60 \\ 20 \\ 1 \\ 20 \end{array}$	10384. 02 10379. 97 10375. 93 10347. 97 10345. 93	$\begin{array}{c} 6s\ ^4\mathrm{P}_{5/2} - 9p\ ^4\mathrm{D}_{7/2} \\ 4d\ ^4\mathrm{F}_{3/2} - 6f\ [2]_{5/2}^* \\ 4d\ ^4\mathrm{D}_{5/2} - 7p\ ^4\mathrm{D}_{7/2}^* \\ 6s\ ^4\mathrm{P}_{5/2} - 9p\ ^4\mathrm{P}_{5/2}^* \\ 4d\ ^4\mathrm{F}_{3/2} - 6f\ [3]_{5/2}^* \end{array}$	
9710. 68 9717. 48 9719. 20 9725. 42 9731. 76	12 40 600 40 350	10295. 12 10287. 91 10286. 09 10279. 51 10272. 82	$4d'\ ^2{ m F}_{5/2}$ — $5f'\ [3]_{5/2}$ $4d\ ^4{ m D}_{5/2}$ — $7p\ ^4{ m P}_{5/2}^2$ $4d\ ^4{ m F}_{9/2}$ — $5f\ [5]_{11/2}^2$ $4d\ ^4{ m D}_{5/2}$ — $7p\ ^4{ m P}_{3/2}^2$ $5p\ ^4{ m P}_{5/2}^2$ — $4d''\ ^2{ m D}_{5/2}$	
9732. 90 9740. 41 9745. 86 9751. 45 9793. 48	50 15 100 90 6000	10271. 61 10263. 69 10257. 95 10252. 07 10208. 08	$\begin{array}{c} 4d' \ ^2\mathrm{F}_{5/2} \!$	
9810. 67 9818. 04 9818. 21 9828. 40 9831. 25	$ \begin{cases} 60 \\ 60 \\ 70 \\ 15 \\ 2 \end{cases} $	10190, 19 10182, 54 10182, 36 10171, 81 10168, 86	$\left.\begin{array}{c} 5p\ ^{4}\mathrm{D}_{3/2}5d\ ^{4}\mathrm{D}_{5/2} \\ 5p\ ^{4}\mathrm{P}_{3/2}^{*}4d^{\prime\prime}\ ^{2}\mathrm{D}_{3/2} \\ 5p^{\prime\prime}\ ^{2}\mathrm{P}_{3/2}^{3/2}6d\ ^{4}\mathrm{D}_{3/2} \\ 4d\ ^{4}\mathrm{P}_{1/2}^{*}-8p\ ^{4}\mathrm{P}_{1/2}^{*} \end{array}\right.$	
9844. 17 9864. 95 9896. 40 9943. 85 9944. 18	$ \begin{array}{c} 7 \\ 2 \\ 10000 \\ 250 \\ 250 \end{array} $	10155. 51 10134. 12 10101. 91 10053. 71 10053. 38	$ \begin{array}{c} 4d\ ^{4}\mathrm{D}_{3/2} -\!$	
9945. 22 9959. 04 9961. 92 9962. 13 9963. 89		10052. 33 10038. 38 10035. 47 10035. 26 10033. 49	$ \left. \begin{array}{c} 4d\ ^4\mathrm{P}_{1/2} \!$	
10001. 96 10025. 23 10045. 71 10056. 87 10061. 53	$ \begin{array}{c} 10 \\ 4h \\ 3h \\ 15 \\ 25 \end{array} $	9995. 30 9972. 10 9951. 77 9940. 73 9936. 12	$ \begin{cases} 4d \ ^4F_{5/2} - 8p \ ^4D_{3/2} \\ 6s \ ^4P_{5/2} - 6f \ [4]_{7/2}^2 \\ 6s \ ^4P_{5/2} - 6f \ [3]_{7/2}^2, \ _{5/2} \\ 4d' \ ^4P_{3/2} - 5f' \ [3]_{5/2}^2 \\ \end{cases} $	
10061. 68 10079. 75 10085. 99 10087. 94 10088. 31	25 10 35 1 10	9935. 97 9918. 16 9912. 03 9910. 11 9909. 75	$ \begin{cases} 4d\ ^4\mathrm{D}_{3/2} - 7p\ ^4\mathrm{P}_{5/2}^2 \\ (^1\mathrm{D}_2)5s\ ^2\mathrm{D}_{3/2} - 6p\ ^4\mathrm{D}_{5/2}^2 \\ 4d'\ ^4\mathrm{P}_{3/2} - 5f'\ [2]_{3/2}^2 \\ 4d\ ^4\mathrm{D}_{3/2} - 7p\ ^4\mathrm{P}_{3/2}^2 \end{cases} $	
10102. 82 10107. 60 10107. 86 10113. 74 10119. 19	$ \left. \begin{array}{c} 12 \\ 100 \\ 150 \\ 2 \\ 3 \end{array} \right. $	9895. 51 9890. 83 9890. 58 9884. 83 9879. 51	$ \left. \begin{array}{c} 58'' \ ^2\mathrm{P}_{1/2} \!$	
10128. 79 10133. 00 10140. 08 10174. 75 10175. 47	4 5 3000 12 20	9870. 14 9866. 04 9859. 15 9825. 56 9824. 86	$\begin{array}{c} 4d\ ^4\mathrm{F}_{5/2}-4f'\ [3]_{5/2}^{5/2} \\ 4d\ ^4\mathrm{F}_{5/2}-8p\ ^4\mathrm{D}_{7/2}^{6} \\ 5s'\ ^4\mathrm{P}_{1/2}-5p\ ^4\mathrm{D}_{3/2}^{3/2} \\ 4d\ ^4\mathrm{F}_{3/2}-8p\ ^4\mathrm{D}_{3/2}^{3/2} \\ 4d\ ^4\mathrm{F}_{5/2}-4f'\ [4]_{7/2}^{6/2} \end{array}$	

Table A2. Observed lines of Br I—Continued

Wavelength	Intensity	Wave number	Classification
A		K	
10184. 02 10184. 49 10197. 61 10208. 91 10225. 23	300 200 18 300 1	9816. 61 9816. 16 9803. 53 9792. 68 9777. 05	$\left.\begin{array}{l} 5p\ ^4\mathrm{P}_{3/2}^*-4s4p^6\ ^2\mathrm{S}_{1/2} \\ 4d\ ^4\mathrm{D}_{7/2}-6p'\ ^2\mathrm{D}_{5/2}^* \\ 4d\ ^4\mathrm{F}_{7/2}-5f\ [5]_{5/2}^6 \\ 4d'\ ^2\mathrm{F}_{5/2}-7p''\ ^2\mathrm{P}_{3/2}^* \end{array}\right.$
10231. 61 10232. 39 10237. 14 10237. 74 10243. 31	8 15 75 6000 3	9770. 95 9770. 21 9765. 68 9765. 10 9759. 79	$\begin{array}{c} 4d\ ^{4}\mathrm{D}_{5/2}6p'\ ^{4}\mathrm{S}_{3/2}^{\circ} \\ 4d\ ^{4}\mathrm{F}_{7/2}5f\ [3]_{7/2}^{\circ},\ _{5/2}^{\circ} \\ 4d\ ^{4}\mathrm{F}_{7/2}5f\ [4]_{7/2}^{\circ} \\ 5s\ ^{4}\mathrm{P}_{3/2}5p\ ^{4}\mathrm{P}_{3/2}^{\circ} \\ 4d\ ^{4}\mathrm{D}_{1/2}7p\ ^{4}\mathrm{D}_{3/2}^{\circ} \end{array}$
10248. 27 10265. 89 10293. 23 10293. 65 10299. 62	$ \begin{array}{c} 3 \\ 15h \\ 1 \\ 50 \\ 1000 \end{array} $	9755. 07 9738. 33 9712. 46 9712. 07 9706. 44	$4d^{\prime\prime}^2\mathrm{D}_{5/2}$ — $5f^{\prime\prime}[3]_{1/2}^2$ $5p^4\mathrm{D}_{5/2}^2$ — $68^{\prime}^2\mathrm{P}_{3/2}$ $4d^{\prime}^4\mathrm{P}_{5/2}$ — $5f^{\prime}[3]_{5/2}^3$ $4d^{\prime}^4\mathrm{P}_{5/2}$ — $5f^{\prime}[3]_{7/2}^7$ $58^{\prime\prime}^2\mathrm{P}_{1/2}$ — $5p^{\prime}^2\mathrm{D}_{3/2}^3$
10305. 99 10310. 62 10310. 92 10312. 90 10314. 03	$ \left. \begin{array}{c} 10 \\ 600s \\ 700 \\ 40 \\ 8 \end{array} \right. $	9700. 44 9696. 08 9695. 80 9693. 94 9692. 87	$\left.\begin{array}{c} 4d\ {}^{4}\mathrm{F}_{3/2}\!\!-\!\!4f'\ [3]_{5/2}^{\circ}\\ ({}^{1}\mathrm{D}_{2})5s\ {}^{2}\mathrm{D}_{5/2}\!\!-\!\!6p\ {}^{4}\mathrm{P}_{3/2}^{\circ}\\ 5p'\ {}^{2}\mathrm{S}_{1/2}^{\circ}\!\!-\!\!5d\ {}^{4}\mathrm{F}_{3/2}\\ 4d\ {}^{4}\mathrm{D}_{5/2}\!\!-\!\!6p'\ {}^{2}\mathrm{D}_{3/2}^{\circ}\end{array}\right.$
10320. 09 10324. 96 10329. 74 10330. 17 10370. 29	$ \left. \begin{array}{c} 15 \\ 3 \\ 80s \\ 100 \\ 1 \end{array} \right. $	9687. 18 9682. 61 9678. 13 9677. 73 9640. 29	$\left.\begin{array}{c} 4d\ ^{4}\mathrm{F}_{3/2}8p\ ^{4}\mathrm{P}_{3/2}^{3/2}\\ 4d'\ ^{4}\mathrm{P}_{5/2}^{3}-5f'\ [2]_{5/2}^{2}\\ \end{array}\right\} \\ -(^{1}\mathrm{D}_{2})5\ ^{2}\mathrm{D}_{3/2}6p\ ^{4}\mathrm{P}_{3/2}^{3}\\ -4d\ ^{4}\mathrm{F}_{3/2}4f'\ [2]_{3/2}^{3}\end{array}$
10374. 37 10377. 65 10384. 15 10388. 40 10390. 74	$12cw \\ 1500 \\ 2 \\ 1 \\ 175$	9636. 50 9633. 45 9627. 42 9623. 48 9621. 32	$ \begin{array}{c} (^{1}\mathrm{D}_{2})5s\ ^{2}\mathrm{D}_{5/2} -\!$
10414. 97 10415. 10 10457. 96 10483. 35 10505. 05	\begin{cases} 8 & 8 & 8 & 8 & 8 & 8 & 8 & 8 & 8 & 8 &	9598. 93 9598. 81 9559. 47 9536. 32 9516. 62	$5p'\ ^2\mathrm{D}_{3/2}^{\circ}$ — $7s\ ^4\mathrm{P}_{3/2}$ $5s'\ ^2\mathrm{P}_{3/2}$ — $5p\ ^4\mathrm{D}_{3/2}^{\circ}$ $5p\ ^4\mathrm{D}_{3/2}^{\circ}$ — $6s''\ ^2\mathrm{P}_{3/2}$ $4d\ ^4\mathrm{D}_{3/2}$ — $6p'\ ^4\mathrm{D}_{1/2}^{\circ}$
$\begin{array}{c} 10507.\ 91 \\ 10513.\ 67 \\ 10529.\ 40 \\ 10529.\ 58 \\ 10554.\ 72 \end{array}$	12 1 3 3 1	9514. 03 9508. 82 9494. 62 9494. 45 9471. 84	$\left.\begin{array}{c} 4d\ ^4D_{1/2}7p\ ^4\mathrm{P}_{3/2}^2\\ 4d^{\prime\prime\prime}\ ^2\mathrm{D}_{3/2}5f^{\prime\prime\prime}\ [3]_{5/2}^2\\ 5p\ ^4\mathrm{D}_{5/2}^24d^{\prime\prime\prime}\ ^2\mathrm{D}_{3/2} \end{array}\right\}$
10566. 15 10566. 87 10569. 67 10591. 51 10599. 94	$ \begin{array}{c} 250 \\ 100 \\ 2d \\ 10w \\ 1 \end{array} $	9461, 59 9460, 95 9458, 44 9438, 94 9431, 43	$\begin{array}{c} 4d\ ^4\mathrm{P}_{1/2}\!\!-\!5f\ [1]_{3/2} \\ 4d\ ^4\mathrm{P}_{1/2}\!\!-\!5f\ [1]_{1/2} \\ 5p'\ ^2\mathrm{D}_{3/2}^2\!\!-\!7s\ ^4\mathrm{P}_{5/2} \\ 4d\ ^4\mathrm{D}_{7/2}\!\!-\!(^1\mathrm{D}_2)5p\ ^5\mathrm{D}_{5/2}^2 \\ 6s\ ^4\mathrm{P}_{5/2}\!\!-\!8p\ ^4\mathrm{D}_{3/2}^2 \end{array}$
10608. 66 10616. 31 10619. 65 10624. 34 10629. 45	40 1 1 1 18	9423. 68 9416. 89 9413. 93 9409. 77 9405. 25	$egin{array}{l} 5p\ ^4{ m D}_{7/2}^2-4d''\ ^2{ m D}_{5/2} \ &4d'\ ^4{ m P}_{3/2}-7p''\ ^2{ m P}_{1/2}^2 \ &5p'\ ^4{ m D}_{1/2}^2-7s\ ^4{ m P}_{3/2} \ &5p\ ^4{ m P}_{5/2}^2-4d'\ ^4{ m P}_{3/2} \end{array}$
10634. 08 10638. 82 10660. 80 10694. 91 10718. 84	1 15 3 12 100	9401. 15 9396. 96 9377. 59 9347. 68 9326. 81	$\begin{array}{c} 4d\ ^4\mathrm{D}_{3/2} \!$

Table A2. Observed lines of Br I—Continued

	TABLE A2.	Ooservea lines of .	Br 1—Continued
Wavelength	Intensity	Wave number	Classification
Ă		K	
10723. 10 10742. 14	15 1000	9323. 11 9306. 58 9306. 29	$ \begin{array}{c} 4d\ ^4\mathrm{D}_{3/2} -\!$
10742. 48 10747. 20 10753. 58	100 20 150	9300. 29 9302. 20 9296. 68	$ \left\{ \begin{array}{c} 6s ^4P_{5/2} \! - \! 4J' [3] _{5/2}^8 \\ 4d ^4D_{7/2} \! - \! 4f [3] _{5/2}^8 \\ 6s ^4P_{5/2} \! - \! 8p ^4D_{7/2}^9 \\ 5p ^4P_{3/2} \! - \! 4d' ^4P_{5/2} \end{array} \right. $
10754. 13 10755. 92 10757. 77 10794. 92 10795. 10	$ \begin{bmatrix} 200 \\ 3000 \\ 6 \\ 5 \\ 10 \end{bmatrix} $	9296. 21 9294. 66 9293. 06 9261. 08 9260. 93	$ \left. \begin{array}{c} 4d\ ^4\mathrm{D}_{7/2} - 4f\ [4]_{7/2}^{\circ} \\ 4d\ ^4\mathrm{D}_{7/2} - 4f\ [4]_{5/2}^{\circ} \\ 6s\ ^4\mathrm{P}_{5/2} - 8p\ ^4\mathrm{P}_{3/2}^{\circ} \\ \end{array} \right\} $
10798. 09 10804. 40 10810. 05 10811. 47 10815. 65	$ \begin{array}{c} 25 \\ 3 \\ 300 \\ 1w \\ 6 \end{array} $	9258. 36 9252. 95 9248. 12 9246. 90 9243. 33	$egin{array}{c} 5p'\ ^2\mathrm{D}_{5/2}^2-5d\ ^4\mathrm{F}_{3/2} \ 6s\ ^4\mathrm{P}_{5/2}-8p\ ^4\mathrm{P}_{5/2}^2 \ 5p\ ^4\mathrm{D}_{5/2}^2-4d''\ ^2\mathrm{D}_{5/2} \ 6s\ ^4\mathrm{P}_{5/2}-4f'\ [2]_{5/2}^3 \ 5p'\ ^2\mathrm{D}_{5/2}^2-5d\ ^4\mathrm{F}_{5/2} \end{array}$
10840. 05 10843. 86 10846. 12 10869. 70 10871. 62	500d $2w$ 10 100 90	9222. 52 9219. 28 9217. 36 9197. 37 9195. 74	$egin{array}{l} 4d\ ^4\mathrm{D}_{5/2}-\!-\!(^1\mathrm{D}_2)5p\ ^2\mathrm{D}_{5/2}^{5}\ 4d\ ^4\mathrm{F}_{5/2}-\!-\!5f[1]_{3/2}^{3}\ 4d\ ^4\mathrm{D}_{3/2}-\!6p'\ ^2\mathrm{D}_{5/2}^{5}\ 4d\ ^4\mathrm{F}_{5/2}-\!5f[2]_{5/2}^{5}\ 5p'\ ^2\mathrm{S}_{1/2}^{6}-\!5d\ ^4\mathrm{D}_{1/2} \end{array}$
10891. 43 10892. 79 10896. 79 10908. 62 10947. 50	250 100 200 60d 1	9179. 02 9177. 87 9174. 50 9164. 55 9132. 00	$\begin{array}{c} 4d\ ^4\mathrm{F}_{5/2}-5f\ [3]_{7/2,5/2}^3\\ 4d\ ^4\mathrm{D}_{5/2}-(^1\mathrm{D}_2)\ 5p\ ^2\mathrm{D}_{3f2}^o\\ 4d\ ^4\mathrm{F}_{5/2}-5f\ [4]_{7/2}^2\\ 4d\ ^4\mathrm{D}_{5/2}-4f\ [1]_{3/2}^3\\ 4d'\ ^2\mathrm{P}_{1/2}-6f\ [1]_{3/2}^o\end{array}$
10947. 89 10960. 83 10961. 30 10973. 48 10979. 00	$ \begin{array}{c} 1\\2w\\3\\500\\300 \end{array} $	9131. 68 9120. 90 9120. 51 9110. 38 9105. 80	$rac{4d'}{4}^2 ext{P}_{1/2}6f\ [1]_{1/2}^{\circ} \ 4d^4 ext{D}_{1/2}6p'^4 ext{D}_{1/2}^{\circ} \ 4d'^2 ext{P}_{1/2}6f\ [2]_{3/2}^{\circ} \ 4d^4 ext{D}_{5/2}4f\ [2]_{5/2}^{\circ} \ 5s''^2 ext{P}_{1/2}5p'^2 ext{S}_{1/2}^{\circ}$
10982. 30 10989. 51 10991. 12 10997. 85 10998. 28	100 1 2 600 400	9103. 07 9097. 10 9095. 76 9090. 20 9089. 84	$egin{array}{l} 4d\ ^4\mathrm{D}_{5/2}4f\ [2]_{3/2}\ 5p'\ ^4\mathrm{S}_{3/2}^*7s\ ^4\mathrm{P}_{3/2}\ 4d'\ ^2\mathrm{F}_{5/2}7p'\ ^2\mathrm{D}_{3/2}\ 4d\ ^4\mathrm{D}_{5/2}4f\ [3]_{7/2}\ 4d\ ^4\mathrm{D}_{5/2}4f\ [3]_{5/2}^* \end{array}$
11007. 42 11010. 45 11012. 31 11013. 22 11018. 55	10 600 5 20 1	9082, 29 9079, 79 9078, 26 9077, 51 9073, 12	$egin{array}{l} 4d\ ^4{ m D}_{3/2}6p'\ ^2{ m S}_{1/2}^{\circ} \ 4d\ ^4{ m D}_{5/2}4f\ [4]_{7/2}^{\circ} \ 5p'\ ^2{ m D}_{3/2}^{\circ}-5d\ ^4{ m F}_{5/2} \ 5p'\ ^2{ m D}_{3/2}^{\circ}-5d\ ^4{ m P}_{1/2} \ 4d''\ ^2{ m D}_{5/2}^{\circ}-5f'\ [3]_{3/2}^{\circ} \end{array}$
11024. 25 11039. 80 11045. 74 11047. 21 11047. 62	10 60 800 <i>d</i> 10 17	9068. 43 9055. 66 9050. 79 9049. 58 9049. 25	$ \left\{ \begin{array}{c} 5p\ ^4\mathrm{P}_{3/2} - 4d'\ ^4\mathrm{P}_{3/2} \\ 5p\ ^2\mathrm{S}_{1/2}^2 - 5d\ ^4\mathrm{D}_{3/2} \\ 5p\ ^4\mathrm{P}_{3/2}^2 - 4d'\ ^2\mathrm{F}_{5/2} \\ 4d\ ^4\mathrm{F}_{3/2} - 5f\ [1]_{3/2}^2 \\ 4d''\ ^2\mathrm{D}_{5/2} - 5f'\ [4]_{7/2}^2 \\ 4d\ ^4\mathrm{F}_{9/2} - 7p\ ^4\mathrm{D}_{5/2}^2 \end{array} \right.$
11074. 08 11093. 51 11094. 24 11096. 59 11100. 65	\$\\ \begin{pmatrix} 40 \\ 250 \\ 100 \\ 20 \\ 1	9027. 62 9011. 81 9011. 22 9009. 31 9006. 02	$\left. \begin{array}{c} 4d\ ^4\mathrm{F}_{3/2}5f\ [2]_{5/2}^{\circ} \\ 5p\ ^4\mathrm{P}_{1/2}^{\circ}4s4p^6\ ^2\mathrm{S}_{1/2} \\ 4d\ ^4\mathrm{F}_{3/2}5f\ [3]_{5/2}^{\circ} \\ 6s\ ^4\mathrm{P}_{3/2}^{\circ}8p\ ^4\mathrm{D}_{3/2}^{\circ} \end{array} \right.$

Table A2. Observed lines of Br I—Continued

		-		
Wavelength	Intensity	Wave number	Classification	
\check{A}		K		
11101. 36 11161. 72 11194. 99 11197. 14 11225. 08	$1 \\ 1w \\ 100 \\ 3 \\ 250$	9005. 44 8956. 74 8930. 12 8928. 41 8906. 18	$egin{array}{l} 4d\ ^4\mathrm{D}_{1/2}6p'\ ^4\mathrm{S}_{3/2}^\circ\ 5p'\ ^4\mathrm{S}_{3/2}^\circ7s\ ^4\mathrm{P}_{5/2}^\circ\ 5s'\ ^4\mathrm{P}_{1/2}5p\ ^4\mathrm{P}_{1/2}^\circ\ 5p'\ ^2\mathrm{D}_{5/2}^\circ5d\ ^4\mathrm{F}_{7/2}^\circ\ 5p\ ^4\mathrm{P}_{3/2}^\circ-4d'\ ^2\mathrm{P}_{3/2}^\circ \end{array}$	
11247. 50 11292. 69 11316. 93 11325. 97 11350. 04	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	8888. 43 8852. 86 8833. 90 8826. 85 8808. 13	$egin{array}{c} 5p' \ ^4\mathrm{D}_{1/2}^\circ - 5d \ ^4\mathrm{P}_{1/2} \ 4d \ ^4\mathrm{D}_{3/2} - (^1\mathrm{D}_2)5p \ ^2\mathrm{D}_{5/2}^\circ \ 5p \ ^4\mathrm{D}_{3/2}^\circ - 6s' \ ^4\mathrm{P}_{1/2} \ \end{array} \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ $	
11367. 25 11367. 97 11437. 61 11447. 27 11464. 58	$\begin{array}{ccc} 4w & (7) \\ 4w & (7) \\ 9 & (11) \\ 18 & (36) \\ 35 & (100) \end{array}$	8794, 79 8794, 24 8740, 69 8733, 32 8720, 13	$egin{array}{l} 4d\ ^4\mathrm{D}_{3/2}-\!\!\!-4f\ [1]_{3/2}^2 \ 4d\ ^4\mathrm{D}_{3/2}-\!\!\!-4f\ [2]_{5/2}^2 \ 4d\ ^4\mathrm{D}_{3/2}-\!\!\!-4f\ [2]_{5/2}^2 \ 4d\ ^4\mathrm{D}_{3/2}-\!\!\!\!-4f\ [3]_{3/2}^3 \end{array}$	
11501. 31 11508. 77 11583. 71 11597. 64 11631. 33	$\begin{array}{ccc} 1 & (m) \\ 30 & (85) \\ 10 & (8) \\ 1 & (4) \\ 2 & (2) \end{array}$	8692. 28 8686. 65 8630. 45 8620. 08 8595. 12	$rac{5p^4\mathrm{D}_{3/2}^*-6s^\prime{}^2\mathrm{P}_{3/2}}{4d^4\mathrm{D}_{1/2}-6p^\prime{}^2\mathrm{S}_{1/2}^*} \ 5s^\prime{}^2\mathrm{P}_{3/2}-5p^4\mathrm{P}_{3/2}^* \ 5p^\prime{}^2\mathrm{D}_{5/2}^*-5d^4\mathrm{D}_{3/2} \ 5p^\prime{}^2\mathrm{D}_{3/2}^*-5d^4\mathrm{D}_{1/2}$	
11636. 08 11640. 94 11666. 20 11742. 85 11810. 29	$\begin{array}{ccc} 1 & (-) \\ 1 & (10) \\ 20d & (169) \\ 400 & (219) \\ 1 & (1) \end{array}$	8591. 61 8588. 02 8569. 43 8513. 49 8464. 88	$egin{array}{l} 5p' \ ^4 ext{S}_{3/2}^{\circ} - 5d \ ^4 ext{F}_{3/2} \ 4d \ ^4 ext{F}_{7/2} - 7p \ ^4 ext{D}_{5/2}^{\circ} \ 5p \ ^4 ext{P}_{3/2} - 4d' \ ^2 ext{P}_{3/2} \ 5s' \ ^2 ext{P}_{3/2} - 5p \ ^4 ext{D}_{5/2}^{\circ} \ 4d' \ ^2 ext{F}_{5/2} - 6f \ [3]_{7/2}^{\circ}, \ _{5/2} \ \end{array}$	
11819, 40 11833, 14 11870, 77 11883, 95 11902, 79	$\begin{array}{cccc} 2 & (-) \\ 1 & (-) \\ 4 & (5) \\ 4 & (6) \\ 5 & (19) \end{array}$	8458. 35 8448. 53 8421. 75 8412. 41 8399. 09	$5p\ ^4\mathrm{D}_{3/2}^{\circ}-4d^{\prime\prime}\ ^2\mathrm{D}_{3/2}^{\circ} \ 5p^{\prime}\ ^2\mathrm{D}_{5/2}^{\circ}-5d\ ^4\mathrm{D}_{5/2}^{\circ} \ 4d\ ^4\mathrm{D}_{1/2}^{\circ}-(^1\mathrm{D}_2)5p\ ^2\mathrm{D}_{3/2}^{\circ} \ 4d\ ^4\mathrm{D}_{1/2}-4f\ [1]_{3/2}^{\circ}$	
11903. 61 11928. 11 11990. 57 12088. 15 12285. 57	8 (16) 8 (13) 15 (38) 2 (-) 2 (m)	8398. 51 8381. 26 8337. 61 8270. 30 8137. 40	$egin{array}{l} 4d\ ^4{ m D}_{1/2}4f\ [1]_{1/2}^{\circ}\ 4d'\ ^2{ m P}_{1/2}4f'\ [2]_{3/2}^{\circ}\ 4d\ ^4{ m D}_{1/2}4f\ [2]_{3/2}^{\circ}\ \end{array} \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \$	
12303. 89 12354. 43 12368. 95 12809. 50	$\begin{array}{ccc} 35 & (275) \\ 1 & (40) \\ 1w & (170) \\ 1 & (400) \end{array}$	8125. 29 8092. 05 8082. 55 7804. 57	$rac{5s'}{4}rac{4P_{1/2}}{4P_{3/2}} - 5p \ ^4P_{3/2}^3 - 4d' \ ^2P_{1/2} \ 5p \ ^4D_{3/2}^3 - 4s4p^6 \ ^2S_{1/2} \ 4d \ ^4F_{9/2} - 4f \ [5]_{11/2}^n$	

 $^{^{\}rm a}\,\lambda4000.599$ If this line itself is not of molecular origin, it is probably the blended hfs components of shorter wavelength for this transition, the long wavelength component having been lost in the molecular background. $^{\rm b}\,\lambda\lambda4446.08,4654.18,6531.39.$ There are Cl lines of almost

these same wavelengths.

° $\lambda5905.66$ This line appears to be the long wavelength hfs component of this transition, the other components having been obscured by the wing of the adjacent line.

d $\lambda6151.10$ Blended with ghost of adjacent strong line.
e $\lambda7617.82$ The transition 5p' $^2\mathrm{D}_{3/2}^2$ —7d $^4\mathrm{D}_{3/2}$ probably contributes to this component.

tributes to this component.

			•				
Com- puted wave- length	Intensity	Com- puted wave number	Classification	Computed wavelength	Intensity	Computed wave number	Classification
Ă		K		Å		K	
11806, 31 12094, 54 12108, 09 12108, 16 12303, 93	$ \left. \begin{array}{c} 1 \\ 3 \\ 14 \\ 275 \end{array} \right. $	8467. 73 8265. 93 {8256. 68 8256. 63 8125. 26	$\begin{array}{c} 4d\ ^4\mathrm{F}_{7/2} \!\!-\! 7p\ ^4\mathrm{P}_{5/2}^{\circ} \\ 5p'\ ^4\mathrm{D}_{1/2}^{\circ} \!\!-\! 5d\ ^4\mathrm{D}_{3/2} \\ 5p'\ ^2\mathrm{D}_{5/2}^{\circ} \!\!-\! 5d\ ^4\mathrm{D}_{7/2} \\ 5p'\ ^2\mathrm{D}_{3/2}^{\circ} \!\!-\! 5d\ ^4\mathrm{D}_{5/2} \\ 5s'\ ^4\mathrm{P}_{1/2} \!\!-\! 5p\ ^4\mathrm{P}_{3/2}^{\circ} \end{array}$	15593. 01 15599. 33 15624. 15 15624. 47 15804. 98	$ \left. \begin{array}{c} 37 \\ 18 \\ 24 \\ 7 \end{array} \right. $	$\begin{array}{c} 6411.\ 38 \\ 6408.\ 78 \\ 6398.\ 60 \\ 6398.\ 47 \\ 6325.\ 39 \end{array}$	$egin{array}{cccccccccccccccccccccccccccccccccccc$
12354. 56 12368. 99 12513. 40 12647. 55 12755. 94	$ \begin{array}{c} 40 \\ 170 \\ 20 \\ 10 \\ 60 \end{array} $	8091, 96 8082, 52 7989, 25 7904, 51 7837, 34	$egin{array}{l} 5p\ ^4\mathrm{P}_{3/2}^{\circ}-4d'\ ^2\mathrm{P}_{1/2}\ 5p\ ^4\mathrm{D}_{3/2}^{\circ}-4s4p^6\ ^2\mathrm{S}_{1/2}\ 5p\ ^4\mathrm{P}_{5/2}^{\circ}-6s\ ^4\mathrm{P}_{3/2}\ 4d'\ ^2\mathrm{P}_{3/2}-4f'\ [2]_{5/2}^{\circ}\ 5p\ ^4\mathrm{D}_{7/2}^{\circ}-4d'\ ^2\mathrm{F}_{7/2} \end{array}$	15890. 32 15948. 40 15955. 01 15977. 73 16253. 66	2 20 8 47 90	6291. 42 6268. 51 6265. 91 6257. 00 6150. 78	$4d^{\prime\prime}{}^{2}\mathrm{D}_{3/2}$ — $4f^{\prime}[2]_{5/2}^{8}$ $5p^{\prime}{}^{2}\mathrm{D}_{3/2}^{3}$ — $4d^{\prime\prime}{}^{2}\mathrm{D}_{5/2}$ $4d{}^{4}\mathrm{D}_{5/2}$ — $6p{}^{4}\mathrm{P}_{5/2}^{8}$ $5p^{\prime}{}^{4}\mathrm{S}_{3/2}^{8}$ — $6s^{\prime}{}^{2}\mathrm{P}_{3/2}$ $5p{}^{4}\mathrm{D}_{7/2}^{2}$ — $4d{}^{4}\mathrm{F}_{5/2}$
12809. 50 12859. 77 12861. 74 12897. 12 13048. 21	400 30 30 48 38	7804. 57 7774. 06 7772. 87 7751. 55 7661. 79	$\begin{array}{c} 4d\ ^4\mathrm{F}_{9/2} - 4f\ [5]_{1/2}^{\circ} \\ 4d'\ ^2\mathrm{F}_{5/2} - 4f'\ [4]_{7/2}^{\circ} \\ 5s''\ ^2\mathrm{P}_{1/2} - 5p\ ^4\mathrm{D}_{3/2}^{\circ} \\ 4d\ ^4\mathrm{F}_{9/2} - 4f\ [4]_{9/2}^{\circ} \\ 5p\ ^4\mathrm{D}_{3/2}^{\circ} - 4d'\ ^2\mathrm{F}_{7/2}^{\circ} \end{array}$	16584. 66 16625. 37 16657. 89 16731. 19 17254. 05	250 10 10 1800 105	6028. 02 6013. 26 6001. 52 5975. 23 5794. 16	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
13064. 15 13217. 17 13328. 04 13349. 73 13392. 33	$ \begin{array}{r} 110 \\ 1700 \\ 10 \\ 36 \\ 20 \end{array} $	7652. 44 7563. 85 7500. 93 7488. 74 7464. 92	$\begin{array}{c} 5p\ ^4\mathrm{P}_{3/2}^{\circ}-6s\ ^4\mathrm{P}_{3/2}^{\circ}\\ 5p\ ^4\mathrm{P}_{5/2}^{\circ}-6s\ ^4\mathrm{P}_{5/2}^{\circ}\\ 5p'\ ^2\mathrm{S}_{1/2}^{\circ}-6s'\ ^4\mathrm{P}_{1/2}^{\circ}\\ 5s'\ ^2\mathrm{P}_{3/2}^{\circ}-5p\ ^4\mathrm{P}_{5/2}^{\circ}\\ 4d'\ ^4\mathrm{P}_{3/2}^{\circ}-4f'\ [3]_{5/2}^{\circ} \end{array}$	17335. 81 17801. 50 17982. 00 18191. 08 18212. 72	25 330 40 60 10	5766. 83 5615. 97 5559. 60 5495. 70 5489. 17	$egin{array}{cccccccccccccccccccccccccccccccccccc$
13484, 87 13584, 48 13674, 13 13808, 19 13808, 45	$ \begin{cases} 9 \\ 30 \\ 300 \end{cases} $ $ \begin{cases} 23 \end{cases} $	7413. 69 7359: 33 7311. 08 {7240. 10 7239. 96	$\begin{array}{c} 5p' \ ^4\mathrm{D}_{1/2}^{\circ} - 6s'' \ ^2\mathrm{P}_{1/2} \\ 5p' \ ^2\mathrm{S}_{1/2}^{\circ} - 6s' \ ^2\mathrm{P}_{3/2} \\ 4d \ ^4\mathrm{F}_{7/2} - 4f \ [5]_{9/2}^{\circ} \\ 6p \ ^4\mathrm{P}_{1/2}^{\circ} - 10s \ ^4\mathrm{P}_{3/2} \\ 4d' \ ^4\mathrm{P}_{5/2} - 4f' \ [3]_{7/2}^{\circ} \end{array}$	18510. 48 18568. 31 18779. 29 19045. 07 19181. 99	180 500 313 547 266	5400. 87 5384. 05 5323. 56 5249. 27 5211. 80	$egin{array}{c} 5p'\ ^2\mathrm{D}_{3/2}^2-4d'\ ^4\mathrm{P}_{3/2} \ 5p'\ ^4\mathrm{D}_{5/2}^2-4d\ ^4\mathrm{F}_{7/2} \ 5p''\ ^2\mathrm{P}_{1/2}^2-4d''\ ^2\mathrm{D}_{3/2} \ 5p''\ ^2\mathrm{P}_{3/2}^2-4d''\ ^2\mathrm{D}_{5/2} \ 5p'\ ^4\mathrm{D}_{1/2}^2-4d'\ ^4\mathrm{P}_{3/2} \ \end{array}$
13821. 95 13822. 07 13822. 56 13833. 14 14078. 51		$ \begin{cases} 7232.89 \\ 7232.83 \\ 7232.57 \\ 7227.04 \\ 7101.08 \end{cases} $	$\begin{array}{c} 4d\ ^4\mathrm{D}_{1/2} -\!\!-\!\! (^1\mathrm{D}_2)5p\ ^2\mathrm{P}_{3/2}^\circ \\ 6p\ ^4\mathrm{D}_{7/2}^\circ -\!\!\!-\!\!\! 10s\ ^4\mathrm{P}_{5/2}^\circ \\ 4d'\ ^4\mathrm{P}_{5/2} -\!\!\!-\!\!\! 8p\ ^4\mathrm{D}_{7/2}^\circ \\ 5p\ ^4\mathrm{P}_{3/2}^\circ -\!\!\!-\!\!\! 6s\ ^4\mathrm{P}_{5/2}^\circ \\ 5p'\ ^4\mathrm{S}_{3/2}^\circ -\!\!\!\!-\!\!\!\! 6s''\ ^2\mathrm{P}_{1/2}^\circ \end{array}$	19183. 17 19317. 02 19497. 62 19606. 57 19733. 62	$ \begin{array}{c c} & 120 \\ & 313 \\ & 263 \\ & 3450 \end{array} $	5211. 48 5175. 37 5127. 43 5098. 94 5066. 11	$egin{array}{cccccccccccccccccccccccccccccccccccc$
14290. 98 14292. 20 14354. 57 14439. 14 14488. 21	} 68 1800 344 105	\$\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\	$egin{array}{l} 4d\ ^4 ext{P}_{1/2}-4f\ [1]_{3/2}^3\ 4d\ ^4 ext{P}_{1/2}-4f\ [1]_{1/2}^6\ 5p\ ^4 ext{D}_{5/2}^5-6s\ ^4 ext{P}_{3/2}\ 5p\ ^2 ext{D}_{5/2}^5-6s\ ^2 ext{P}_{3/2}\ 5p\ ^2 ext{D}_{3/2}^3-6s\ ^4 ext{P}_{1/2} \end{array}$	19810. 58 19894. 71 19925. 72 20073. 55 20162. 50	452 68 138 7 138	5046. 43 5025. 09 5017. 27 4980. 32 4958. 35	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
14599. 73 14607. 80 14888. 70 14896. 37 14923. 43	$ \begin{array}{r} 338 \\ 21 \\ 1250 \\ 55 \\ 12 \end{array} $	6847. 57 6843. 79 6714. 67 6711. 21 6699. 04	$\begin{array}{c} 5p\ ^4\mathrm{P}_{1/2}^{\circ}{-}6s\ ^4\mathrm{P}_{3/2}^{\circ}\\ 5s^{\prime\prime}\ ^2\mathrm{P}_{1/2}^{\circ}{-}5p\ ^4\mathrm{P}_{1/2}^{\circ}\\ 5p\ ^4\mathrm{D}_{7/2}^{\circ}{-}6s\ ^4\mathrm{P}_{5/2}^{\circ}\\ 5p^{\prime}\ ^4\mathrm{D}_{1/2}^{\circ}{-}6s^{\prime}\ ^4\mathrm{P}_{1/2}\\ 4d\ ^4\mathrm{F}_{5/2}^{\circ}{-}4f\ [2]_{5/2}^{\circ}\end{array}$	20248. 38 20281. 73 20405. 97 20624. 67 21093. 83	500 7 547 43	4937. 32 4929. 20 4899. 19 4847. 24 4739. 43	$\begin{array}{c} (^{1}\mathrm{D}_{2})5p \ ^{2}\mathrm{F}_{5/2}^{5}-(^{1}\mathrm{D}_{2})4d \ ^{2}\mathrm{G}_{7/2} \\ 5p \ ^{4}\mathrm{D}_{3/2}^{3}-4d \ ^{4}\mathrm{F}_{5/2} \\ 5p' \ ^{4}\mathrm{S}_{3/2}^{3}-4d' \ ^{4}\mathrm{P}_{3/2} \\ 5p' \ ^{2}\mathrm{D}_{5/2}^{5}-4d' \ ^{2}\mathrm{F}_{7/2} \\ 4d \ ^{4}\mathrm{F}_{7/2}-6p \ ^{4}\mathrm{D}_{5/2}^{5/2} \end{array}$
14968. 61 14969. 33 14991. 94 15003. 82 15052. 95	\right\} \ 85 \ 89 \ 35 \ 10	\$\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\	$\begin{array}{c} 4d\ ^4\mathrm{F}_{5/2}-4f\ [3]_{7/2}^{\circ}\\ 4d\ ^4\mathrm{F}_{5/2}-4f\ [3]_{5/2}^{\circ}\\ 4d\ ^4\mathrm{F}_{5/2}-4f\ [4]_{7/2}^{\circ}\\ 5p\ ^4\mathrm{P}_{3/2}^{\circ}-4d\ ^4\mathrm{F}_{5/2}^{\circ}\\ 4d\ ^4\mathrm{F}_{3/2}-(^1\mathrm{D}_2)5p\ ^2\mathrm{D}_{5/2}^{\circ} \end{array}$	21213. 11 21285. 74 21330. 29 21631. 91 21743. 98	14 17 37 359 14	4712. 78 4696. 70 4686. 89 4621. 54 4597. 72	$\begin{array}{c} 5p' \ ^4\mathrm{D}_{1/2}^{\circ} - 4d' \ ^2\mathrm{P}_{3/2} \\ (^1\mathrm{D}_2)5p \ ^2\mathrm{F}_{7/2}^{\circ} - (^1\mathrm{D}_2)4d \ ^2\mathrm{G}_{9/2} \\ 5p \ ^4\mathrm{D}_{3/2}^{\circ} - 4d \ ^4\mathrm{P}_{1/2} \\ 5p \ ^4\mathrm{P}_{3/2}^{\circ} - 4d \ ^4\mathrm{D}_{3/2} \\ 6s \ ^4\mathrm{P}_{3/2} - (^1\mathrm{D}_2)5p \ ^2\mathrm{P}_{3/2}^{\circ} \end{array}$
15146. 09 15185. 29 15185. 43 15186. 82 15258. 39			$\begin{array}{c} 4d^{\prime\prime}{}^{2}{\rm D}_{5/2}\!-\!4f^{\prime}[3]_{7/2}^{2} \\ 5p^{\prime\prime}{}^{2}{\rm P}_{3/2}^{3}\!-\!6s^{\prime\prime}{}^{2}{\rm P}_{1/2} \\ 4d{}^{4}{\rm F}_{3/2}\!-\!4f[1]_{3/2}^{3/2} \\ 4d{}^{4}{\rm F}_{3/2}\!-\!4f[1]_{1/2}^{3/2} \\ 4d^{\prime\prime}{}^{2}{\rm D}_{5/2}\!-\!4f^{\prime}[4]_{7/2}^{2} \end{array}$	21787. 24 21801. 59 22720. 19 22865. 65 23404. 15	$ \begin{array}{r} 469 \\ 6 \\ 60 \\ 950 \\ 12 \end{array} $	4588. 59 4585. 57 4400. 17 4372. 18 4271. 58	$\begin{array}{c} 5p \ ^4\mathrm{P}_{5/2}^{5}-4d \ ^4\mathrm{D}_{5/2} \\ 5p' \ ^2\mathrm{S}_{1/2}^{5}-6s \ ^4\mathrm{P}_{3/2} \\ 5p' \ ^4\mathrm{S}_{3/2}^{3}-4d' \ ^2\mathrm{P}_{3/2} \\ 5p \ ^4\mathrm{P}_{5/2}^{5}-4d \ ^4\mathrm{D}_{7/2} \\ 4d \ ^4\mathrm{F}_{5/2}-6p \ ^4\mathrm{D}_{3/2}^{5} \end{array}$
15311. 39 15359. 71 15422. 36 15539. 25 15570. 03	63 22 28 3 219	6529. 30 6508. 76 6482. 32 6433. 56 6420. 84	$egin{array}{l} 4d\ ^4\mathrm{F}_{3/2}-4f\ [2]_{5/2}^{\circ}\ 4d\ ^4\mathrm{F}_{3/2}-4f\ [3]_{5/2}^{\circ}\ 4d\ ^4\mathrm{D}_{7/2}-6p\ ^4\mathrm{P}_{3/2}^{\circ}\ 5p'\ ^2\mathrm{D}_{5/2}^{\circ}-4d''\ ^2\mathrm{D}_{5/2}\ 5p\ ^4\mathrm{P}_{3/2}^{\circ}-4d\ ^4\mathrm{P}_{1/2} \end{array}$	23513. 15 23732. 96 24099. 97	206 40 2	4251. 78 4212. 40 4148. 25	$egin{array}{lll} 5p\ ^4\mathrm{P}_{3/2}^3 - 4d\ ^4\mathrm{D}_{5/2} \ 5p\ ^4\mathrm{P}_{1/2}^3 - 4d\ ^4\mathrm{D}_{1/2} \ 4d\ ^4\mathrm{F}_{5/2} - 6p\ ^4\mathrm{D}_{5/2}^5 \end{array}$