1	Supplementary Materials for:
2	How much of the sediment in Gale crater's central mound was fluvially
3	transported?
4	
5	Bradley J. Thomson, Debra L. Buczkowski, Larry S. Crumpler, Kimberly D. Seelos, and
6	Caleb I. Fassett
7	
8 9 10 11 12	This PDF file includes: Supplementary Text Tables S1 to S4 Figures S1 to S5
13	Supplementary Text
14	This supplementary section provides: (1) a description of the data used; (2) an
15	expanded description of our methodology; (3) an explanation of the uncertainty analysis
16	and error propagation, and (4) an expanded description of the paleotopography scenario.
17	Supplementary tables and figures are appended to the end of the text.
18	(1) Source Data: We used topography data from the Mars Orbital Laser
19	Altimeter (MOLA) instrument (Smith et al., 1999) and digital elevation models (DEMs)
20	derived from High Resolution Stereo Camera (HRSC) stereo images (Neukum and
21	Jaumann, 2004) to measure morphometric parameters of the Gale sedimentary mound
22	and surrounding fluvial channels. MOLA data are collated into gridded products with a
23	pixel spacing 463 m (equivalent to 128 pixels per degree) (Smith et al., 2003). The polar
24	orbit of MOLA on Mars Global Surveyor resulted in a dense network of altimetry
25	measurements at high latitudes, but equatorial gaps of up to ten km. As the crater Gale

lies near the equator (it is centered at 5.4°S, 137.8°E), we used higher resolution HRSC
DEMs supplement the MOLA gridded data. Table S1 lists the HRSC DEMs used in this
work. Four DEMs with a pixel spacing of 75 m cover the entire crater (Fig. S1) and a
region to the west and constitute our main sources of topographic information; a higher
resolution DEM (50 m pixel spacing) covers the western portion of the mound.

31 (2) Methodology, valley network measurements: First, the positions of valley 32 networks required some manual adjustment (Fig. S2). Hynek et al. (2010) identified and 33 mapped valley networks across Mars using THEMIS daytime infrared data mosaicked at 34 231 m/pixel supplemented with MOLA topography. Since this time, both uncontrolled 35 (Edwards et al., 2011) and controlled 100 m/pixel THEMIS mosaics have been produced 36 (Fergason et al., 2013), resulting in small offsets in the position of mapped features. Most 37 features are only a few hundred meters offset from their previous mapped locations 38 (typically 2-4 pixels), but some positional errors of up to 30 pixels or 3 km are apparent 39 (Fergason et al., 2013).

40 Each mapped valley network segment was assigned a stream order (Strahler, 41 1957) by Hynek et al. (2010); these were slightly revised as appropriate. In this system, 42 channels without tributaries are designated as 1st-order. Where two 1st-order channels 43 join, the downstream segment is designated an order of 2; where two second-order 44 channels join, a segment of order 3 is formed; and so forth. However, if two channels join 45 that are of different order (e.g., a segment of order 1 merging with an order 2 segment), 46 the order of the downstream segment is not incremented – it retains the highest order of 47 the upstream channel segment. To assess the eroded volume in Farah Vallis and the 48 upstream valley network, we measured segment shapes in a region to the southwest of

Gale that lies within HRSC DEM h1960_0000_da4. Profiles across selected channel segments were extracted from the DEM, and these were analyzed to determine the channel's cross-sectional area. The area was taken to be the difference between the current ground level and an assumed pre-erosional surface. **Table S2** provides a listing of all profiles collected and measured for this project. Three to four profiles were averaged together for each segment. As expected, lower-order channel segments typically have smaller cross-sectional areas than higher-order segments.

56 All cross-sectional area measurements of a given stream order (i.e., 1, 2, 3, or 4) 57 were then combined using length-weighted median values that represent typical channel 58 dimensions of each type (**Table S3**). The combined sum of these figures represents our 59 estimate of total eroded volume of this valley network. One complicating factor is that for 60 second-order valley network segments in particular, we observed that the choice of 61 method for determining "typical" cross-sectional area values resulted in a notably large 62 difference. As evidenced in the histogram below (Figure S3), the population of area 63 values is distinctly bi-model. While the majority of areas are <0.1 km², a subset of larger features fall between 0.1 and 0.6 km². The median cross-sectional area is 0.12 km², near 64 65 the low end of the range of values. The weighted, length-normalized median area is 0.36 km², which is a factor of three higher. To address this, we have computed the weighted 66 median of 2nd order channel segments less than and greater than 0.1 km² separately. The 67 weighted median of small channels is 0.0435 km², and the weighted median of large 68 channels is 0.388 km². A simple average of these two weighted medians is 0.216 km², 69 70 and this is the value we adopt in Table S3. A bar chart that summarizes the water-71 mobilized sediment volume compared to the mound volume is given in Figure S4.

72 (3) Error propagation. Our eroded valley network volume estimate is the sum of
73 four products (equation S1):

74
$$V_{vn} = a_1L_1 + a_2L_2 + a_3L_3 + a_4L_4.$$
 (S1)

Here, V_{vn} is the total cumulative volume of the valley network, a_1 to a_4 are the mean cross-sectional areas of stream order 1 to 4, respectively, and L_1 to L_4 are the total cumulative lengths of stream order types 1 to 4, respectively (**Table S3**). The uncertainty in V_{vn} is dominated by the uncertainty in the mean areas; uncertainty in the length values is likely <1% and they are neglected in equation S2. The uncertainty in the volume calculation can be estimated using fractional uncertainties:

81
$$\delta V_{\nu n} = V_{\nu n} \sqrt{\frac{\delta a_1}{a_1} + \frac{\delta a_2}{a_2} + \frac{\delta a_3}{a_3} + \frac{\delta a_3}{a_3}}$$
(S2)

In Eq. S2, δa_1 to δa_3 are the uncertainty in the cross-sectional area averages, taken here to be the standard deviations. This results in a formal uncertainty of 12.5×10^3 km³, a value that is 150% larger than the volume of 8×10^3 km³. Despite the fact that this valley network volume estimate is uncertain to within a factor of two, it is still almost an order of magnitude smaller than the mound volume (which approaches ~ 10^5 km³).

87 The mound volume was determined by tabulating the difference between the 88 current surface elevation and the assumed base elevation of -4.5 km, yielding a volume of 89 9414 km³. To estimate the uncertainty in our estimate of the volume of the mound, we 90 buffered the enclosed polygon that represents the mapped outermost extent of the mound 91 (Thomson et al., 2011) one MOLA pixel (463 m) inward and outwards to obtain 92 minimum (9301 km³) and maximum (9520 km³) error estimates, respectively. Thus we obtain our reported value of $9.4\pm0.1\times10^3$ km³. However, there is an additional 93 94 component of uncertainty related to the assumed base elevation surface under the Gale

mound. Below we conduct a sensitivity analysis to explore how the mound volume (and
by extension, the lower elevation level) varies given the choice of this parameter.

In the topographic profile of Gale given in **Figure S5**, the north-south topographic asymmetry of the crater is evident. The crater floor in the northern section of the crater lies more than 1 km deeper than the floor to the south of this crater. This asymmetry could be partly attributed to intrinsic factors such as Gale's formation on the northwarddipping dichotomy boundary or post-impact tilt, or it could reflect a depositional or erosional asymmetry that affected the infilling material.

103 We consider four potential base level surfaces to explore their effect on the 104 volume calculations presented in this paper. First, we consider mound with a flat base at 105 -4.5 km. Second, we consider a plane fit to the measured mound base elevation levels. 106 Finally, following the approach of *Gabasova and Kite* (2018), we scale a radially 107 averaged profile of a fresh example of complex martian crater (Tooting, 27.86 km in 108 diameter) and a profile of a relatively fresh peak-ring crater (Galle, 223.5 km in diameter) 109 to match the diameter of Gale using depth-diameter scaling relationships from *Tornabene* 110 et al. (2013). These basal surfaces are given in Fig. S5; the resulting mound volumes are 111 tabulated below.

As given in **Table S4**, three of the four basal surfaces considered (i.e., a flat plane, scaled Tooting crater, and scaled Galle basin) give reasonably consistent results, i.e., the volume of the mound is ~9.4 to 9.6×10^3 km³. For the third basal surface considered, however, the mound volume is ~6×10³ km³, which is about one-third less than the other cases. While this result does not alter the conclusion that the overwhelming majority of the mound in Gale is not attributable to aqueous transport processes, the elevation level that constitutes the boundary between the lower and intermediate elevation zone is more than 0.8 km higher than the other cases. Therefore, we consider this to be reflective the overall uncertainty associated with this elevation value. Accordingly, we have identified a range of elevation values on the mound in Figures 1 and 3.

122

(4) Paleotopography scenario.

123 Topography of the southern half of Gale was used to estimate the current average 124 crater rim height of 0.58±0.41 km (Grotzinger et al., 2015). Based on measurements of 125 other morphologically fresh craters between 40° S and 40° N (Robbins and Hynek, 2012), 126 the initial rim height of Gale is estimated to have been ~ 1.6 km. The difference between 127 the observed and expected rim height suggests ~1 km of net vertical erosion (Grotzinger 128 et al., 2015), but the authors concede that this value is likely an overestimate based on 129 preservation of Gale's ejecta blanket (Irwin et al., 2005). Using a geometric model of the 130 potential sediment yield from this magnitude of wall erosion, Grotzinger et al. (2015) estimate that $\sim 9 \times 10^3$ km³ of sediment could have delivered to the crater floor, a value that 131 132 we note is roughly equivalent to the present-day mound volume. Assuming that the 133 density of eroded and deposited material is the same (an assumption made by Grotzinger et al. [2015]), this yields 0.5 to 0.6 km of deposition over the 18,250 km² area of the 134 135 crater floor, exclusive of the central peak region.

Based on rover observations of the exposed stratigraphy in Aeolis Mons, it also
has been estimated that there was at least ~1 km of sediment present above the current
crater floor that has since been eroded away (Grotzinger et al., 2015; Fedo et al., 2017).
The preferred interpretation of the exposed stratigraphy is that it represents fluvial-deltaic
deposits that undergo lateral facies transitions into fine-grained lacustrine deposits. The

inferred sediment transport direction in these deposits is generally southward, i.e., toward
the current mound, a geometry that would necessitate the presence of a now-eroded
sequence of sediments in the gap between the crater walls and current mound.

144 The precise geometry of this now-eroded sedimentary sequence is necessarily 145 speculative. The simplest configuration is to consider this volume of material spread 146 uniformly from the crater wall toward central peak, accounting for a depth of 0.5 to 0.6 147 km of material (after Grotzinger et al., 2015). Additional sources of sediment considered in this current analysis collectively constitute $\sim 1.3 \times 10^3$ km³, a volume that would account 148 149 for a 70 m thick deposit if spread uniformly over the same area or about 300 m thick if 150 tallied within the context of the existing mound topography. The lower elevation 151 boundary would be -3.7 km in a flat plane is assumed for the basal boundary of the 152 mound (Table S5).

153 There are several complicating factors with the deep burial and exhumation 154 scenario. One paradoxical element of this scenario is that it would appear to require more 155 sediment that visible means of transport into Gale would permit. Eolian processes such as 156 the settling of dust or volcanic ash could be one potential mechanism that could 157 contribute additional sediment into Gale. However, the grain-size segregation necessary 158 to create the observed fine laminae is more likely in water, and therefore the settling of 159 fine grains directly from the atmosphere "is excluded... ... as a primary sediment 160 accumulation mechanism [of the lower mound]" (Grotzinger et al., 2015). Evaporite 161 sequences, perhaps delivered by groundwater, are another means to introduce material, 162 but rover evidence suggests that the evaporitic material is pore-filling, not matrix-forming 163 (e.g., Rampe et al., 2017).

164	The observed depth of Gale presents another potential complication to this
165	scenario—local sediment sources and sinks inside the crater are out of balance by a factor
166	of 2 to 4 in favor of the latter. Again using measurements of morphologically fresh
167	craters (Robbins and Hynek, 2012), the initial depth of a crater Gale's size is estimated to
168	have been 4.0 to 5.0 km; the current depth is 3.1 ± 0.7 km (Grotzinger et al., 2015). The
169	difference between the observed depth and expected depth range is 0.9 to 1.9 km,
170	suggesting the crater has been shallowed by sedimentary infill. Yet the volume of
171	sediment potentially shed from the walls is not sufficient to account for the inferred 1 to 2
172	km of shallowing.
173	Finally, two recent rover-based observations support the notional that former
174	burial depths have been minimal in Gale. The first is observations that the compressive
175	strength of rock drilled by the MSL rover range from very weak to medium strong (Peters
176	et al., 2018). The presence of very weak rocks (i.e., with inferred compressive strengths
177	~5 MPa) is not straightforwardly compatible with deep burial and exhumation. Second,
178	accelerometer data on MSL has been used to infer the near-surface density of the terrain
179	over which the rover has traversed (Lewis et al., 2019). The inferred average density
180	value of 1.680 ± 0.180 g/cm ³ is low, and value indicates a high porosity (40±6%). As
181	stated by Lewis et al., "This is a typical value for soils and poorly lithified sediments, but
182	porosity is typically reduced in sedimentary rocks that have experienced burial and
183	compaction."
184	
185	

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Table S1: List of HRSC Digital Elevation Models (DEMs) of Gale and vicinity

Conference2013, Volume 44, p. abstract #2592.

HRSC product ID	Resolution (m)	Comments
h1916_0000_da4	75	Includes eastern portion of Gale
h1927_0000_da4	75	Includes central portion of Gale
h1938_0000_da4	75	Includes western portion of Gale
h1960_0000_da4	75	Includes part of valley network SW of Gale
h5273_0000_da4	75	Includes western half of Gale
h4235_0001_da4	50	Includes western half of Gale

pitted impact melt-bearing craters, in Proceedings Lunar and Planetary Science

Table S2: Valley network segment measurements

Object ID	Stream Order	Center Lat (°N)	Center Lon (°E)	Cross-sectional Area [km²]	Segment Length [km]	Mean Area [km²]	Std. Dev. [km ²]
454 a	1	-7.783	134.338	0.077	21.3	0.08	0.01
454 b	1	-7.861	134.401	0.092			
454 c	1	-7.924	134.456	0.073			
451 a	1	-7.730	134.746	0.060	13.0	0.11	0.07
451 b	1	-7.750	134.744	0.032			
451 c	1	-7.810	134.765	0.158			
409 a	1	-8.432	134.178	0.015	5.6	0.05	0.03
409 b	1	-8.453	134.193	0.070			
409 c	1	-8.481	134.202	0.058			
387 a	1	-8.823	134.106	0.057	21.5	0.18	0.20
387 b	1	-8.665	134.275	0.417			
387 c	1	-8.611	134.313	0.073			
390 a	1	-8.668	134.340	0.328	8.3	0.15	0.16
390 b	1	-8.646	134.340	0.083			
390 c	1	-8.600	134.322	0.040			
397 a	1	-8.611	134.599	0.111	12.3	0.11	0.01
397 b	1	-8.506	134.531	0.096			
397 c	1	-8.491	134.524	0.126			
518 a	1	-8.280	134.016	0.041	5.6	0.03	0.01
518 b	1	-8.275	134.043	0.033			
518 c	1	-8.491	134.524	0.014			
521 a	1	-8.187	134.057	0.023	6.6	0.04	0.03
521 b	1	-8.230	134.072	0.078			
521 c	1	-8.256	134.085	0.020			
520 a	1	-8.349	134.046	0.005	5.6	0.04	0.04
520 b	1	-8.340	134.081	0.083			
520 c	1	-8.321	134.091	0.019			
519 a	1	-8.372	134.095	0.048	7.5	0.05	0.03

Object ID	Stream	Center Lat	Center	Cross-sectional	Segment	Mean Area	Std. Dev.
Object ID	Order	(°N)	Lon (°E)	Area [km ²]	Length [km]	$[km^2]$	$[km^2]$
519 b	1	-8.368	134.101	0.080			
519 c	1	-8.335	134.130	0.024			
408 a	1	-8.415	134.286	0.039	12.2	0.08	0.04
408 b	1	-8.371	134.321	0.096			
408 c	1	-8.316	134.335	0.118			
415 a	1	-8.339	134.418	0.053	9.5	0.06	0.01
415 b	1	-8.294	134.386	0.054			
415 c	1	-8.264	134.333	0.071			
430 a	1	-8.132	134.320	0.059	6.5	0.08	0.02
430 b	1	-8.154	134.355	0.097			
430 c	1	-8.186	134.369	0.088			
402 a	2	-8.477	134.268	0.525	20.4	0.45	0.08
402 b	2	-8.390	134.362	0.452			
402 c	2	-8.450	134.465	0.368			
395 a2	2	-8.594	134.039	0.334	46.5	0.39	0.06
395 b	2	-8.512	134.166	0.388			
395 c	2	-8.515	134.193	0.461			
399 a	2	-8.566	134.318	0.149	14.3	0.25	0.12
399 b	2	-8.542	134.388	0.384			
399 с	2	-8.521	134.414	0.219			
522-523 a	2	-8.293	134.098	0.056	4.2	0.06	0.02
522-523 b	2	-8.307	134.101	0.085			
522-523 с	2	-8.316	134.113	0.040			
522-523 d	2	-8.321	134.128	0.071			
420 a	2	-8.324	134.143	0.204	12.9	0.35	0.14
420 b	2	-8.326	134.179	0.479			
420 c	2	-8.285	134.269	0.358			
516 a	2	-8.105	134.455	0.044	8.0	0.04	0.02
516 b	2	-8.084	134.477	0.057			
516 c	2	-8.059	134.513	0.036			
516 d	2	-8.039	134.513	0.011			
422-424 a	2	-8.233	134.312	0.089	5.3	0.08	0.03
422-424 b	2	-8.207	134.345	0.043			
422-424 c	2	-8.197	134.363	0.098			
426-427-438 a	2	-8.185	134.393	0.023	7.9	0.10	0.13
426-427-438 b	2	-8.171	134.422	0.037			
426-427-438 с	2	-8.123	134.442	0.252			
404-405 a	3	-8.477	134.500	0.230	29.1	0.30	0.14
404-405 b	3	-8.448	134.588	0.192			
404-405 c	3	-8.309	134.630	0.509			
404-405 d	3	-8.156	134.639	0.268			
524 a	3	-8.037	134.545	0.110	8.1	0.15	0.04
439 b	3	-8.072	134.592	0.159			
529 c	3	-8.084	134.616	0.179			
528 a	4	-8.057	134.662	0.223	6.2	0.28	0.09
528 b	4	-8.039	134.684	0.237			
528 c	4	-8.038	134.716	0.377			
429 a	4	-8.062	134.753	0.400	16.1	0.27	0.12
429 b	4	-8.116	134.792	0.165			
429 c	4	-8.148	134.832	0.234			
431 a	4	-8.139	134.966	0.466	6.4	0.63	0.29
431 b	4	-8.121	134.994	0.456			
431 c	4	-8.097	135.005	0.965			

	Object ID	Stream Order	Center Lat (°N)	Center Lon (°E)	Cross-sectional Area [km ²]	Segment Length [km]	Mean Area [km²]	Std. Dev. [km ²]
	432 a	4	-8.097	135.053	0.307	53.5	0.21	0.09
	432 b	4	-8.013	135.235	0.154			
	432 c	4	-7.999	135.276	0.158			
-	Table S3: Cumulative volume by stream order							

Stream Order	Normalized Median Area [km²]	Stand. Dev. [km ²]	Total Length [km]	Volume [km ³]
1	0.07	0.08	2970	218
2	0.22	0.17	1203	260
3	0.23	0.13	574	132
4	0.31	0.21	645	198
			Total Cumulativ	ve Volume [km ³] 808

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2	+.	7

Table S4: Gale basal elevation options with modern topography given in Fig. S4

Assumed basal topography	Current mound volume [km ³]	Lower elevation level [km]	Ratio volume moved by water to mound volume	Fraction of volume moved by water
Flat base	9410	-4.21	1:7.54	13.3%
Inclined plane	6070	-3.47	1:4.87	20.5%
Tooting crater, scaled	9480	-4.25	1:7.59	13.2%
Galle basin, scaled	9630	-4.27	1:7.71	13.0%



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Table 55: Gale basal elevation options with interred paleolopograp	basal elevation options with inferred paleotopograp
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Assumed basal topography	Mound paleovolume [km ³]	Lower elevation level [km]
Flat base	18400	-3.67
Inclined plane	15100	-3.34
Tooting crater, scaled	18500	-3.66
Galle basin, scaled	18600	-3.67





Figure S1. Shaded relief map of Gale crater from HRSC-derived topography. (a) Interior channels and inverted channels are given in pink and orange respectively [from Le Deit et 258 259 al., 2013]; larger inward-draining channels are outlined in gray. (b) Same view as Fig. 1a 260 indicating sub-divided interior watersheds. The N=705 separate watersheds are assigned 261 colors based on their mean flow direction.



263 264

Figure S2. Comparison of valley networks mapped by Hynek et al (2010) (given in blue) with those re-mapped in this study (given in yellow). White border marks extent of 265 266 HRSC DEM H1960_0000; background is MOLA shaded relief map.



Cross-sectional area (km²) **Figure S3.** Histogram of the cross-sectional areas of second-order valley network segments. X-axis labels are bin centers; tick marks are at increments of 0.1 km². 269



Figure S4. Comparison of the volume of sediment mobilized by water with the total

volume of sediment in the mound.



distance from center [km]
 Figure S5. Two-segment topographic profile of Gale crater from north to south in black.

Given in red is flat base elevation level of -4.5 km; the dashed green line is an inclined plane fit the mound base. The orange and blue curves are radially averaged profiles of

Tooting crater and the Galle basin, respectively, from *Gabasova and Kite* (2018) that

were scaled to Gale's diameter (*Tornabene et al.*, 2013).

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