Supplementary Information for

Energy Dispersive X-ray Absorption Spectroscopy

with an Inverse Compton Source

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S1 Possible elements which can be investigated with the MuCLS

As stated in the main manuscript, the energy of our ICS source can be tuned freely in the range from 15 keV to 35 keV which can cover the absorption edges highlighted in the following figure:

1																	18
1 H hydrogen																	He helium
1.008 [1.0078, 1.0082]	2		ŀ	C-edg	ve							13	14	15	16	17	4.0026
3 Li Ithium 6.94 [6.938, 6.997]	4 Be beryllium 9.0122		I	L-edg	ge		5 B boron 96.81 [10.806, 10.821]	6 C carbon 12.011 [12.009, 12.012]	7 N nitrogen 14.007 [14.006, 14.008]	8 O 15.999 (15.999, 16.000)	9 F fluorine 18.998	10 Ne neon 20.160					
Na sodum	Mg magnesium		Data unavailable from the cited source AI Si P 16 CI AV														
22.990	24.305 [24.304, 24.307]	3	4	5	6	7	8	9	10	11	12	26.962	28.065 [26.064, 28.066]	30.974	32.06 [32.059, 32.076]	35.45 [35.446, 35.457]	39.95 [39.792, 39.963]
19 K potassium	Ca calcium	21 Sc scandium	22 Ti titanium	V V vanadium	Cr chromium	25 Mn manganese	Fe iron	Co cobalt	28 Ni nickel	Cu copper	30 Zn zinc	Ga gallum	Ge germanium	AS arsenic	34 Se selenium	35 Br bromine	36 Kr krypton
39.098	40.078(4)	44.956	47.867	50.942	51.996	54.938	55.845(2)	58.933	58.693	63.546(3)	65.38(2)	69.723	72.630(8)	74.922	78.971(8)	[79.901, 79.907]	83.798(2)
Rb rubidium	38 Sr strontium	39 Y yttrium	Zr zirconium	41 Nb niobium	42 Mo molybdenum	43 Tc technetium	44 Ru ruthenium	45 Rh rhodium	Pd palladium	47 Ag silver	48 Cd cadmium	49 In indium	50 Sn tin	Sb antimony	52 Te tellurium	53 iodine	Xe xenon
85.468	87.62	88.905	91.224(2)	92.906	95.95		101.07(2)	102.91	106.42	107.87	112.41	114.82	118.71	121.76	127.60(3)	126.90	131.29
55 Cs caesium	56 Ba barium	57-71 Ianthanoids	72 Hf hafnium 176.49(2)	73 Ta tantalum	74 W tungsten 183.84	75 Re rhenium	76 OS osmium	77 Ir iridium 192.22	78 Pt platinum	79 Au gold 196.97	80 Hg mercury 200.59	81 TI thallium 204.38 (204.38)	82 Pb lead	83 Bi bismuth	84 Po polonium	At astatine	Rn radon
87 Fr	Ra 88	89-103 actinoids	104 Rf	105 Db	106 Sg	107 Bh	108 Hs	109 Mt	110 Ds	111 Rg	112 Cn	113 Nh	114 FI	115 Mc	116 Lv	117 Ts	118 Og
francium	radium		rutherfordium	dubnium	seaborgium	bohrium	hassium	meitnerium	darmstadtium	roentgenium	copernicium	nihonium	flerovium	moscovium	livermorium	tennessine	oganesson
			57 La Ianthanum	Ce cerium	59 Pr praseodymium	60 Nd neodymium	Pm promethium	Sm samarium	63 Eu europium	64 Gd gadolinium	65 Tb terbium	66 Dy dysprosium	67 Ho holmium	68 Er erbium	Tm thulium	YD ytterbium	71 Lu lutetium
			138.91	140.12	140.91	144,24		150.36(2)	151.96	157.25(3)	158.93	162.50	164.93	167.26	168.93	173.05	174.97
			AC actinium	90 Th thorium 232.04	91 Pa protactinium 231.04	92 U uranium 238.03	93 Np reptunium	94 Pu plutonium	95 Am americium	96 Cm curium	97 Bk berkelium	98 Cf californium	99 Es einsteinium	fermium	101 Md mendelevium	102 No nobelium	103 Lr Iawrencium

Figure S1 Absorption edges that can be covered by the energy range (15 keV - 35 keV) available at the MuCLS. The absorption edge data is from [1] and the periodic table is adapted from [2]

S2 Warping correction

If the crystal bending is not perfectly, one may observe that the absorption fringes are not parallel to each other. In this case, a warping correction can be applied as shown in Figure S2. The same warping parameters generated from the standard sample (such as a Ag foil) can be applied to other Ag containing materials using the same crystal.



Figure S2 Four points are chosen manually (see yellow points) on the fringes of the original absorption image. Then the fringes are automatically aligned by transforming yellow points positions to red points

positions.



S3 Results with different total exposure times

Figure S3 (a) Ag foil absorption images of averaging 10 frames, 5 frames and 1 frame (from top to bottom) (b) XANES spectra for averaging different number of frames

The XANES results that have been shown in the main manuscript were averages of 10 frames of 30 s exposure time for I_s and I_0 . Figure S2 shows the absorption images and the corresponding spectra of Ag using the same Si(311) reflection with averaging different number of frames. Also for a total exposure time down to 1 minute (30 s with the sample and 30 s without sample), the spectrum is still in good agreement with a reference spectrum taken at the synchrotron.

S4 EXAFS measurement

The EXAFS spectrum in the main manuscript was obtained by stitching two absorption images measured for two angles with a Si(311) diffraction order. These two images are shown in Figure S4.1 (a) and (b).



Figure S4.1 (a) Absorption image measured at the crystal angle covering the K-edge.(b) Another absorption image measured after rotating the crystal by 0.05 degree.

The limitation of the current method is that distinct features are needed in the second image and some spectrum overlap (e.g., at least 1 peak/feature in common) is required in order to stitch them together. However, there are also other measurement methods instead of stitching two spectra, which can overcome this limitation:

1. To use other diffraction orders such as Si(111) and adjust the bending curvature, which increases the maximum energy range in a single image, but at the cost of a degraded energy resolution. However, since the primary signal of EXAFS mainly arises from the low frequency components, the energy resolution does not have to be as high in EXAFS as it needs to be in XANES. Switching between Si(311) and Si(111) diffraction orders is feasible by simply rotating the crystal. Therefore, one can combine both

high resolution and lower resolution, e.g., to measure XANES by using the Si(311) diffraction order and then rotate the crystal to the Si(111) diffraction order for the EXAFS measurement.

2. Instead of rotating to just one other angle and thus measuring two spectra, which are shifted with respect to each other by a large fraction of the maximum energy range and therefore might be difficult to stitch, a scanning fashion can be adopted (see example in Figure S4.2). The Laue crystal (bending radius \sim 11 m, Si(311) diffraction order) was rotated in very fine steps of 0.001 degree per step, 100 steps in total. At each step, the energy range is extended by a small fraction (\sim 3 eV, 1.2 % of maximum energy range) with a big fraction of the energy range overlapping with the previous one. The scanning with fine steps not only extends the energy range smoothly, but also increases the data statistics in the overlapping regions. Therefore, there is no need to average multiple image frames at each step to increase the statistics (the spectra shown in the main manuscript are averaged of 10 image frames). If better optimized fine steps are chosen e.g., 10 steps in total, the total acquisition time is unchanged.



Figure S4.2 Extending the energy range by rotating the crystal in fine steps. In this example, we used the same Si(311) diffraction order of the Laue crystal (bending radius ~11 m, maximum energy range ~250 eV). The crystal was rotated in total by 0.1 degree in 100 steps (30 s exposure time at each step for both the I_s and I₀). Each spectrum at each step was shifted by a linear function. The final overlapped 100 spectra (the last figure) were not calibrated or normalized, illustrating the stability and feasibility of the scanning fashion.

Finally, regarding the EXAFS energy range, it should be mentioned that the x-ray intensity of the 26 keV energy configuration drops quickly after 26 keV and there is no flux left above 26.5 keV (see Figure S4.3). For Ag whose K-edge is at ~25.5 keV, the energy range for EXAFS can be further extended by switching to another MuCLS energy configuration e.g., 27 keV. Some preliminary tests have been done on fast

energy configuration switching lately, during which 1 keV energy changes could be realized in less than 5 s.



Figure S4.3 MuCLS measured source spectrum of 26 keV and 27 keV configurations (measured with energy dispersive detector, KETEK AXAS-D, KETEK GmbH, Munich, Germany). The blue line highlights the Ag K-edge. Yellow and green regions correspond to 1 keV, 1.5 keV energy range, respectively.

Reference

- [1] http://skuld.bmsc.washington.edu/scatter/AS_periodic.html
- [2] https://iupac.org/what-we-do/periodic-table-of-elements/