

**Supplementary Material to: Electronic Properties of the
Coronene Series from Thermally-Assisted-Occupation Density
Functional Theory**

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TABLE S1. Comparison of the singlet-triplet energy gap E_{ST} (in kcal/mol) of n -coronene, calculated using spin-unrestricted TAO-LDA with the 6-31G(d) and 6-31G basis sets.

n	6-31G(d)	6-31G	Difference
2	51.31	52.41	-1.10
3	30.05	30.81	-0.76
4	17.68	18.16	-0.48
5	9.96	10.21	-0.25
6	5.49	5.59	-0.10
7	3.18	3.21	-0.03
8	2.02	2.03	-0.01
9	1.41	1.40	0.01

TABLE S2. Singlet-triplet energy gap E_{ST} (in kcal/mol) of n -coronene, calculated using spin-unrestricted TAO-LDA, KS-LDA, and KS-B3LYP with the 6-31G basis set. Here, the experimental data is taken from the literature [1].

n	TAO-LDA	KS-LDA	KS-B3LYP	Expt.
2	52.41	60.21	61.10	55.35 [1]
3	30.81	40.77	43.23	
4	18.16	28.94	31.35	
5	10.21	20.82	22.05	
6	5.59	14.86	13.31	
7	3.21	10.31	2.91	
8	2.03	6.78	-9.06	
9	1.40	3.77		
10	1.05	0.53		
11	0.82	1.34		

TABLE S3. Vertical ionization potential IP_v (in eV), vertical electron affinity EA_v (in eV), and fundamental gap E_g (in eV) for the lowest singlet state of n -coronene, calculated using spin-unrestricted TAO-LDA with the 6-31G basis set. Here, the experimental data (given in parentheses) are taken from the literature [2–4].

n	IP_v	EA_v	E_g
2	6.72 (7.29 [2])	0.58 (0.50 [3, 4])	6.14 (6.79 [2–4])
3	5.81	1.70	4.11
4	5.30	2.35	2.95
5	4.99	2.77	2.22
6	4.80	3.04	1.76
7	4.68	3.22	1.46
8	4.60	3.35	1.26
9	4.55	3.44	1.11
10	4.51	3.51	0.99
11	4.47	3.57	0.90

TABLE S4. Symmetrized von Neumann entropy S_{vN} for the lowest singlet state of n -coronene, calculated using spin-restricted TAO-LDA with the 6-31G basis set.

n	S_{vN}
2	0.01
3	0.15
4	0.54
5	1.29
6	2.39
7	3.77
8	5.30
9	6.93
10	8.65
11	10.44

TABLE S5. Active orbital occupation numbers (HOMO–8, ..., HOMO–1, HOMO, LUMO, LUMO+1, ..., and LUMO+8) for the lowest singlet state of n -coronene, calculated using spin-restricted TAO-LDA with the 6-31G basis set. For brevity, HOMO is denoted as H, LUMO is denoted as L, and so on.

n	2	3	4	5	6	7	8	9	10	11
H–8	2.000	2.000	2.000	2.000	2.000	1.997	1.990	1.971	1.933	1.870
H–7	2.000	2.000	2.000	1.999	1.996	1.987	1.965	1.927	1.869	1.791
H–6	2.000	2.000	2.000	1.999	1.996	1.987	1.965	1.927	1.868	1.790
H–5	2.000	2.000	2.000	1.999	1.992	1.970	1.917	1.828	1.707	1.574
H–4	2.000	2.000	1.998	1.989	1.964	1.913	1.833	1.728	1.609	1.491
H–3	2.000	2.000	1.998	1.989	1.964	1.913	1.832	1.727	1.608	1.490
H–2	2.000	2.000	1.997	1.975	1.907	1.781	1.624	1.474	1.350	1.257
H–1	1.999	1.988	1.948	1.865	1.743	1.602	1.467	1.353	1.264	1.198
H	1.999	1.988	1.948	1.864	1.742	1.601	1.466	1.352	1.263	1.197
L	0.001	0.012	0.052	0.136	0.262	0.411	0.557	0.685	0.787	0.863
L+1	0.001	0.012	0.052	0.136	0.261	0.409	0.556	0.684	0.786	0.862
L+2	0.000	0.000	0.003	0.025	0.097	0.234	0.408	0.578	0.719	0.823
L+3	0.000	0.000	0.002	0.010	0.032	0.074	0.144	0.241	0.357	0.478
L+4	0.000	0.000	0.002	0.010	0.031	0.074	0.144	0.240	0.356	0.477
L+5	0.000	0.000	0.000	0.001	0.005	0.021	0.061	0.136	0.247	0.378
L+6	0.000	0.000	0.000	0.001	0.004	0.012	0.031	0.065	0.117	0.189
L+7	0.000	0.000	0.000	0.001	0.004	0.012	0.031	0.064	0.117	0.188
L+8	0.000	0.000	0.000	0.000	0.001	0.002	0.007	0.022	0.052	0.104

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- [1] R. Abouaf and S. Diaz-Tenderoa, *Phys. Chem. Chem. Phys.*, 2009, **11**, 5686–5694.
 - [2] G. Mallocci, C. Joblin, and G. Mulas, *Astron. Astrophys.*, 2007, **462**, 627–635.
 - [3] M. A. Duncan *et al.*, *Chem. Phys. Lett.*, 1999, **309**, 49–54.
 - [4] G. Chen, R. G. Cookscor, E. Corpuz, and L. T. Scott, *J. Am. Soc. Mass Spectrosc.*, 1996, **7**, 619–627.